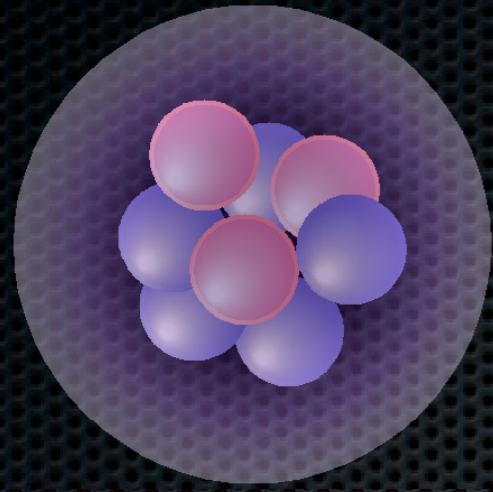


Study of the Properties of Atomic Nuclei with RI Beam

Maya Takechi, Niigata University

Atomic Nuclei



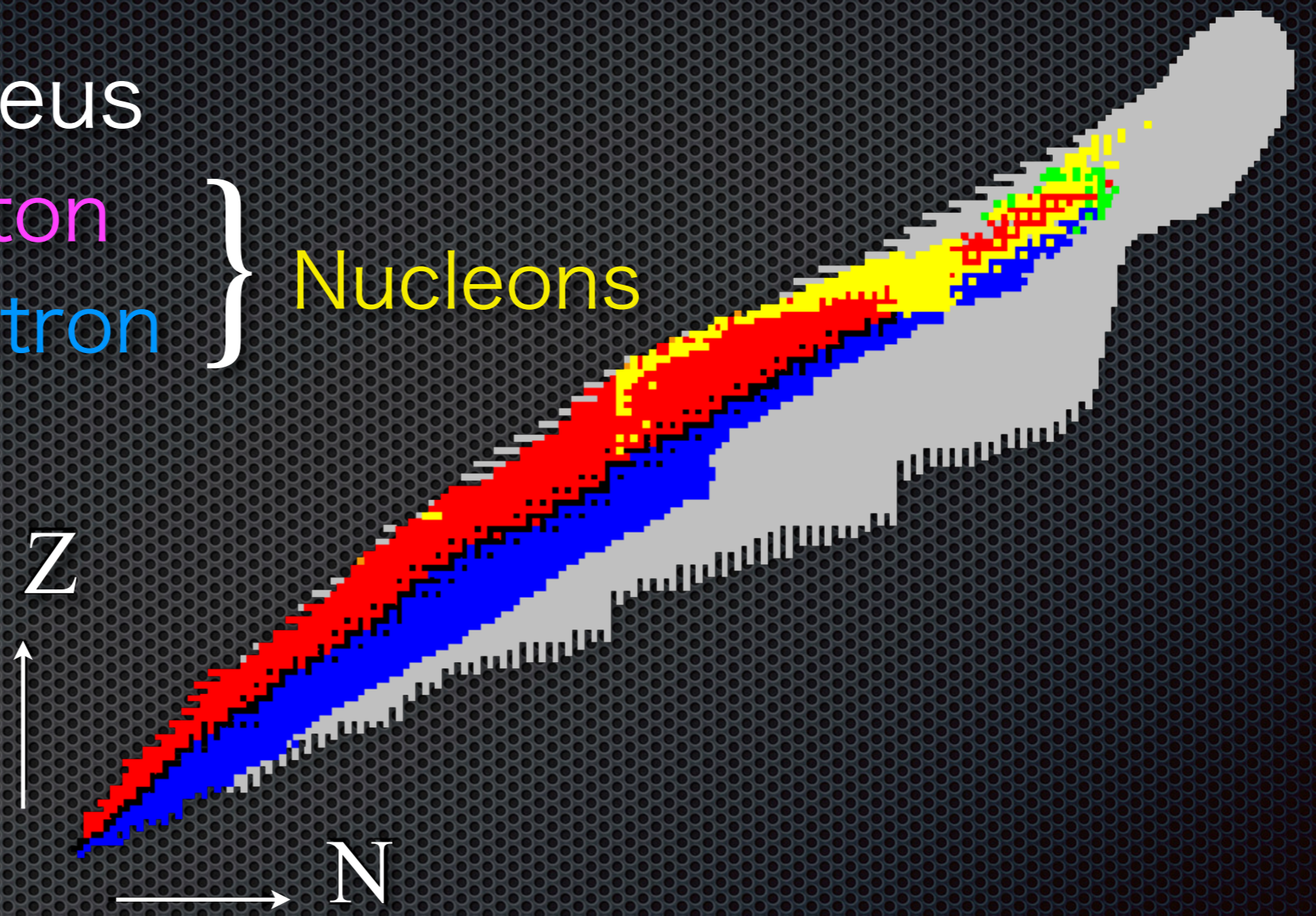
Nucleus

Proton

Neutron

} Nucleons

Proton Number Z
Neutron Number N
Mass Number A
 $A = N + Z$



Stable Nuclei ~300, Unstable Nuclei ~8000

Experimentally confirmed

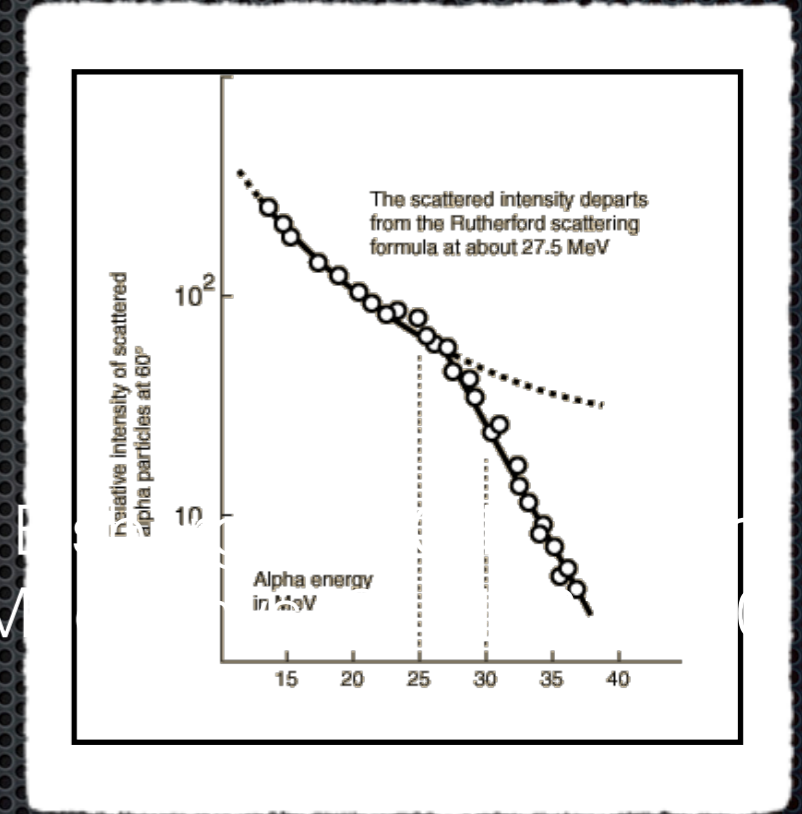
Unstable Nuclei ~ 3000

Nuclear Property : Size, Density

Known Properties from the Study of Stable Nuclei

Radius Alpha particle scattering on Stable Nuclei

Nuclear Radii ~ 1~10 fm



R.M. ...
Rev. M...

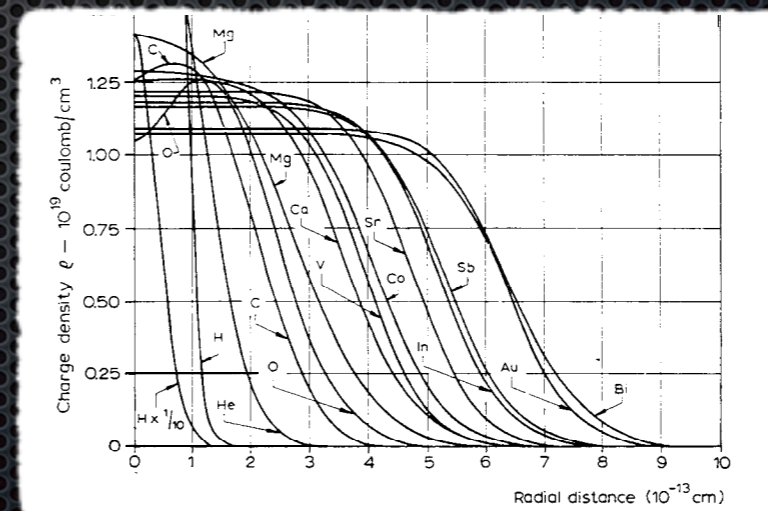
Electron Scattering

Nuclear radii follow the function of mass number

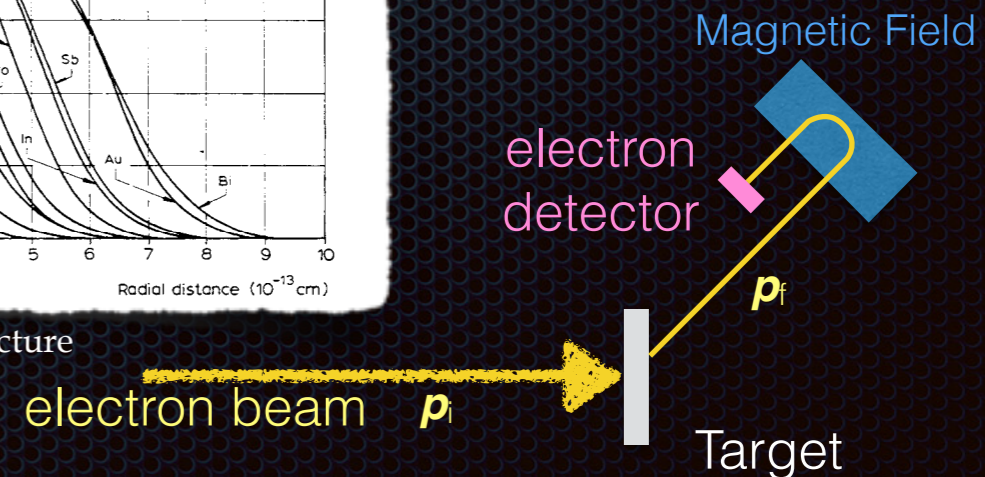
$$R \propto r_0 \cdot A^{1/3}$$

Saturation Density

$$m \rho_0 \approx 2.8 \times 10^{14} \text{ g} \cdot \text{cm}^{-3}$$



R. Hofstadter Nobel Lecture



Nuclear Property : Mass, Binding Energy

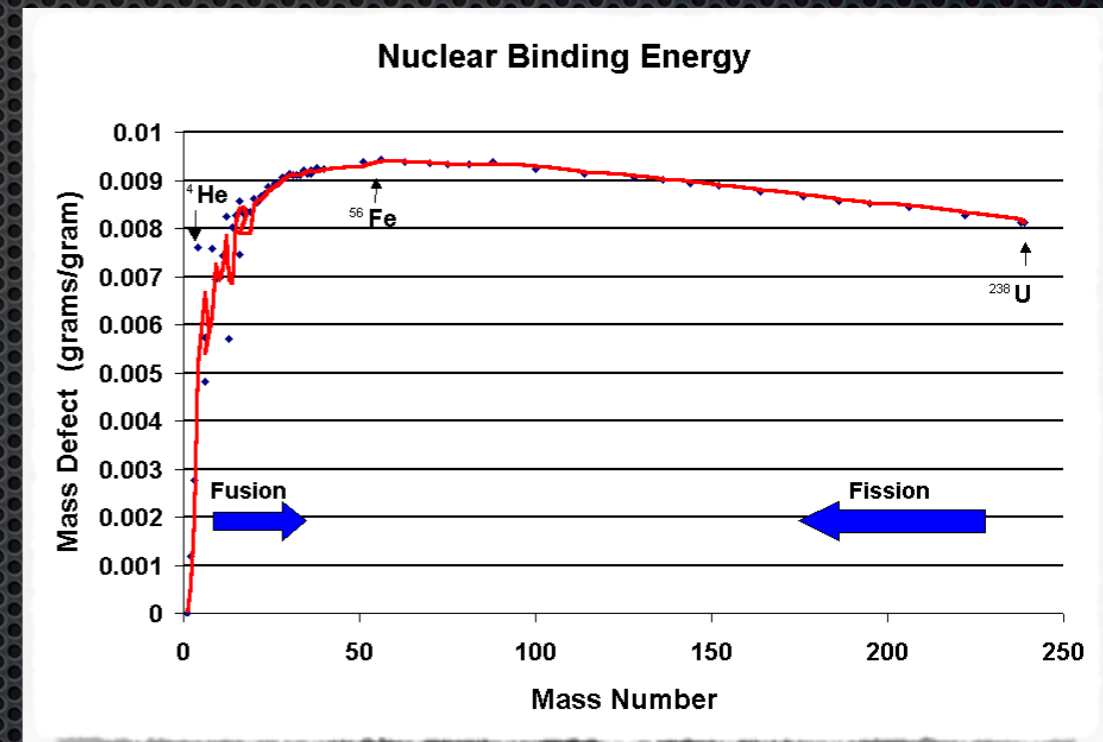
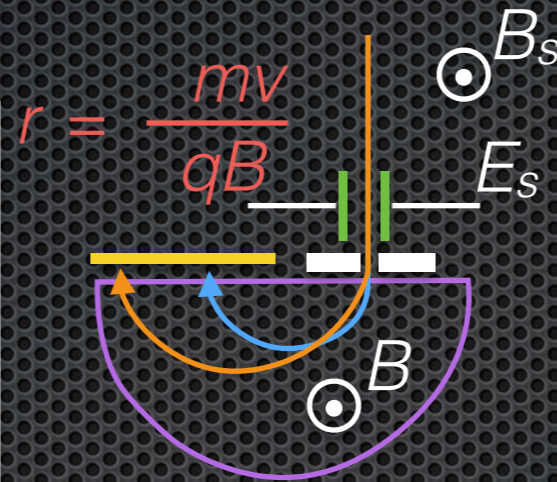
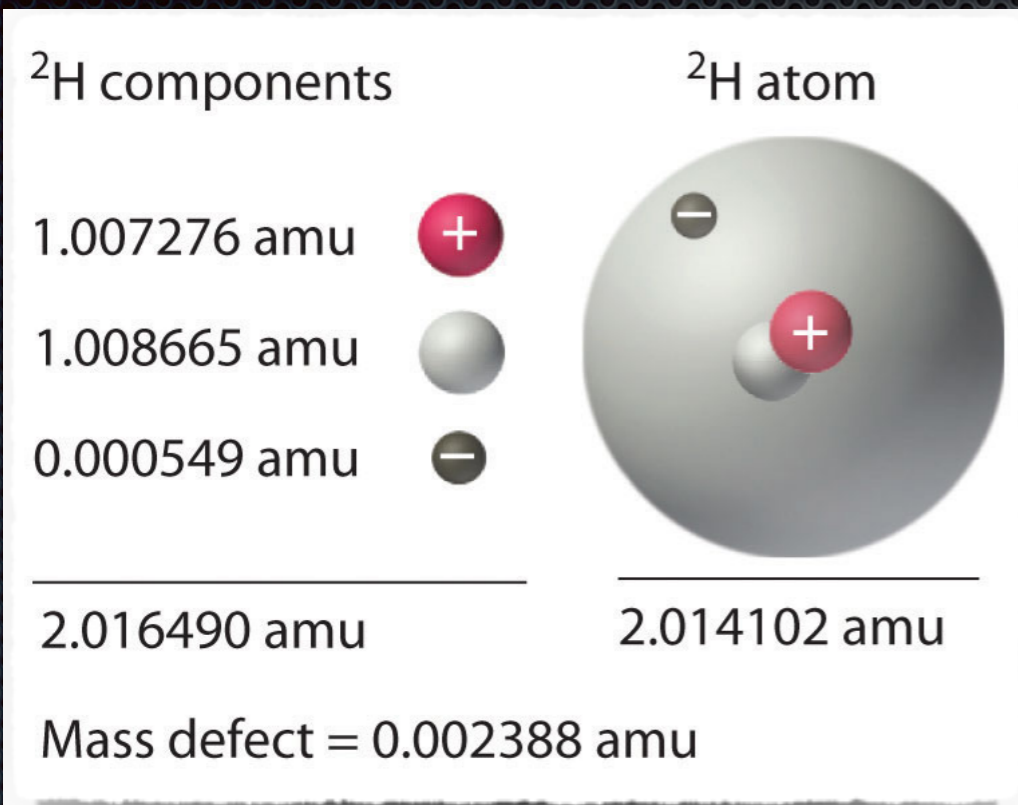
Ion source

Known Properties from the Study of Stable Nuclei

Mass Measurements by Mass spectrometer

Bainbridge Mass spectrograph (1930s)

Mass Defect



<http://chemwiki.ucdavis.edu/>

<http://www.chem.fsu.edu/>

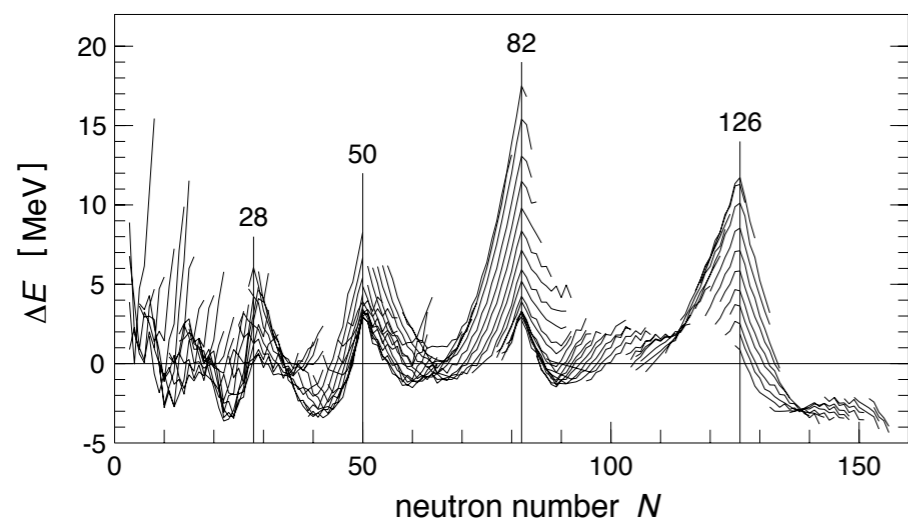
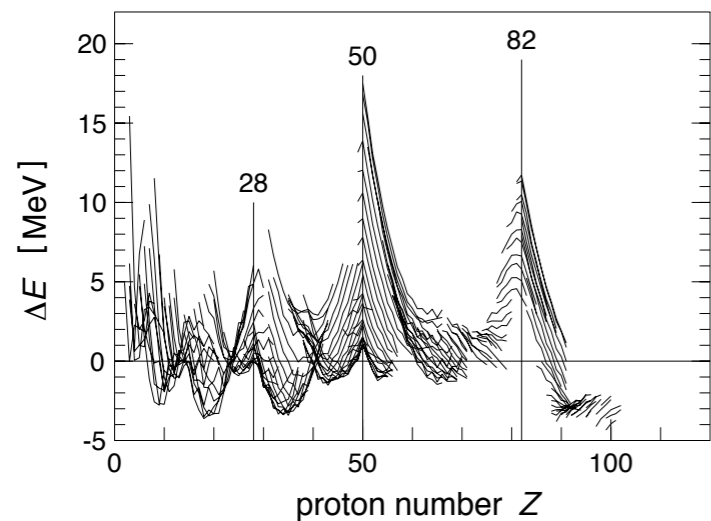
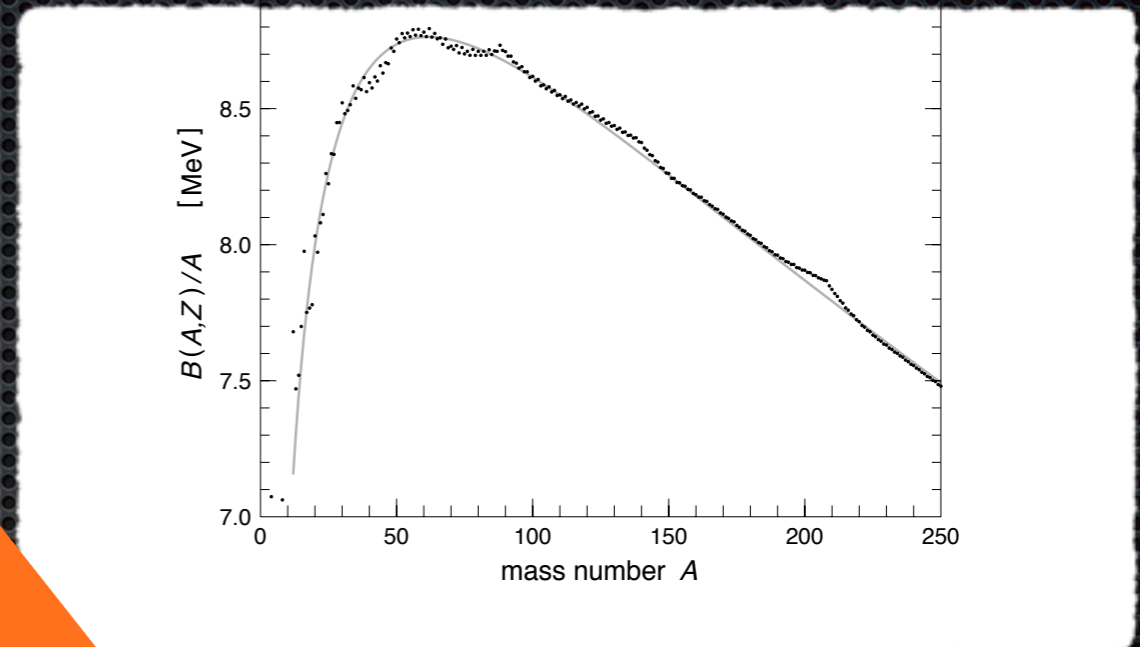
$$M(A, Z) = Z \cdot m_p + N \cdot m_n - B(A, Z)$$

Binding Energy / nucleon
~8 MeV

Nuclear Structure : Magic Number

Figures are from the lecture note of Prof. K. Muto, Tokyo Institute of Technology

$B(A, Z) = b_{\text{vol}} A$	体積エネルギー
$- b_{\text{surf}} A^{2/3}$	表面エネルギー
$- b_{\text{Coul}} \frac{Z^2}{A^{1/3}}$	Coulomb エネルギー
$- b_{\text{sym}} \frac{(A - 2Z)^2}{A}$	対称エネルギー
$- \Delta(A)$	対エネルギー



$$\Delta E = M_{\text{Expt}} - B(A, Z)$$

Nuclei are more strongly bound when the number of constituent nucleons are certain magic numbers

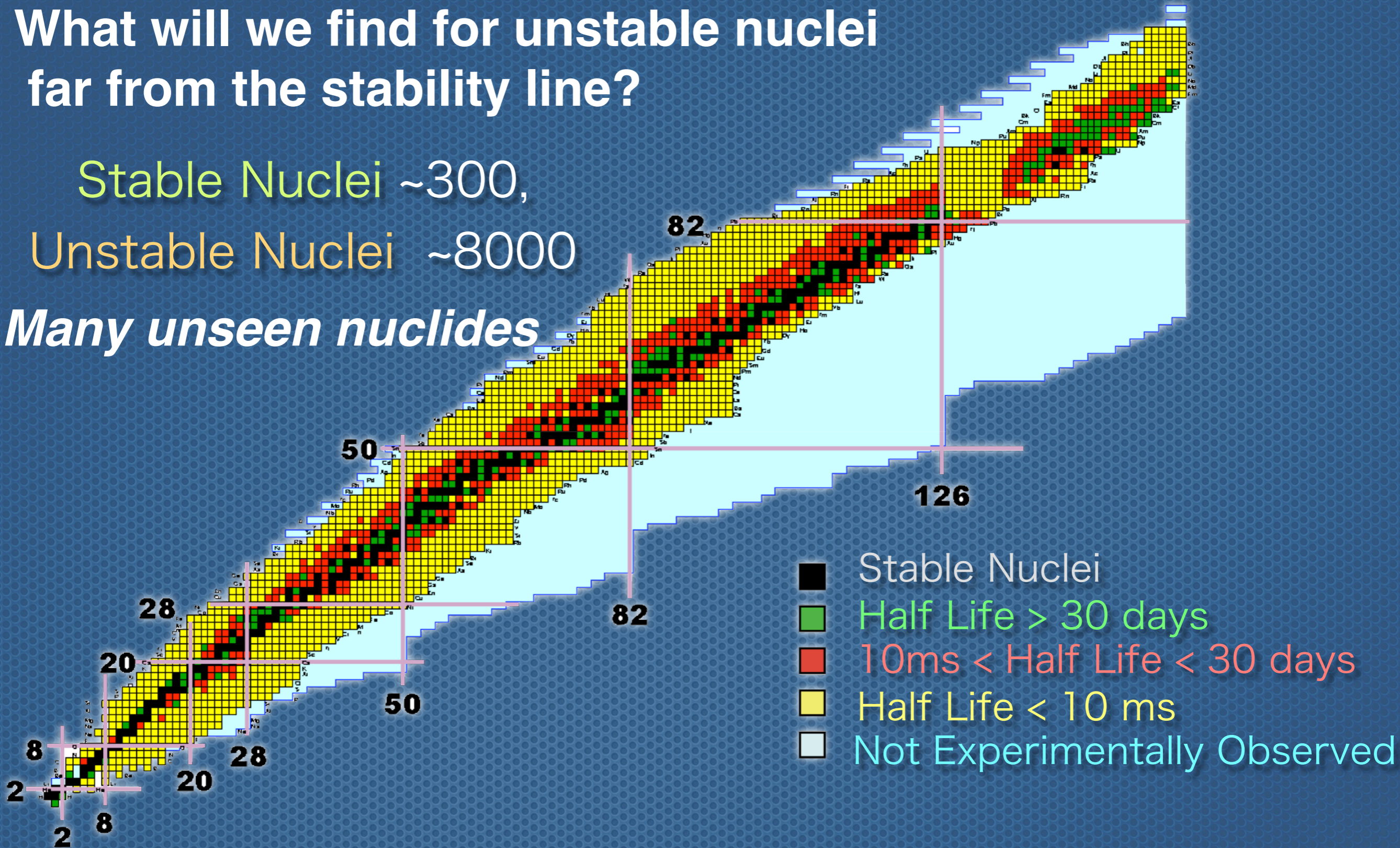
N or $Z = 8, 20, 28, 50, 82, 128$
Closed Shell

What will we find for unstable nuclei far from the stability line?

Stable Nuclei ~300,

Unstable Nuclei ~8000

Many unseen nuclides

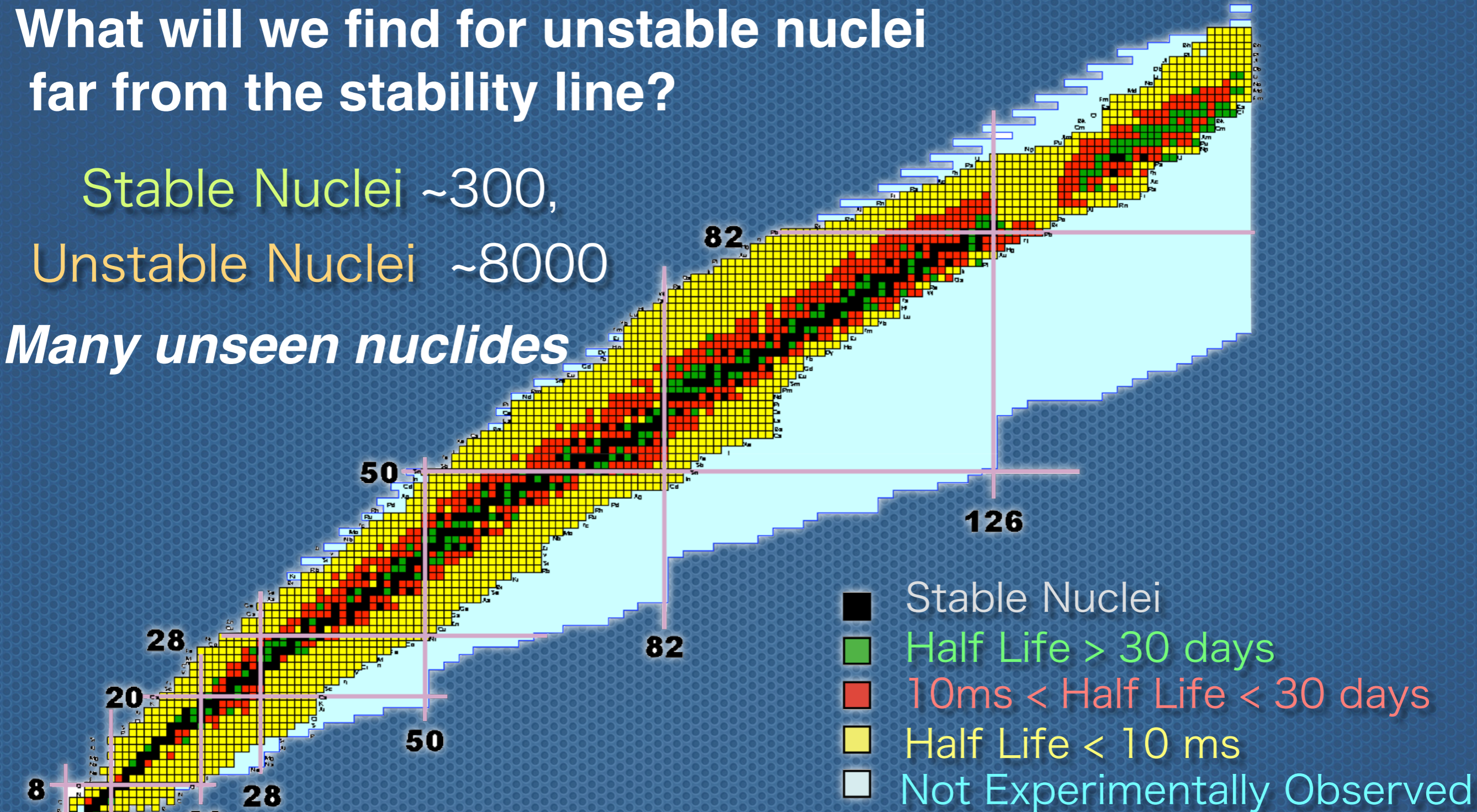


What will we find for unstable nuclei far from the stability line?

Stable Nuclei ~300,

Unstable Nuclei ~8000

Many unseen nuclides



- Stable Nuclei
- Half Life > 30 days
- 10ms < Half Life < 30 days
- Half Life < 10 ms
- Not Experimentally Observed

Loosely-bound, Drip Line Nuclei

Weak Binding Energy

How is the structure ?

Magic Number ?

$$R \propto r_0 \cdot A^{1/3} \quad ?$$

$$m \rho_0 \approx 2.8 \times 10^{14} \text{ g} \cdot \text{cm}^{-3} \quad ?$$

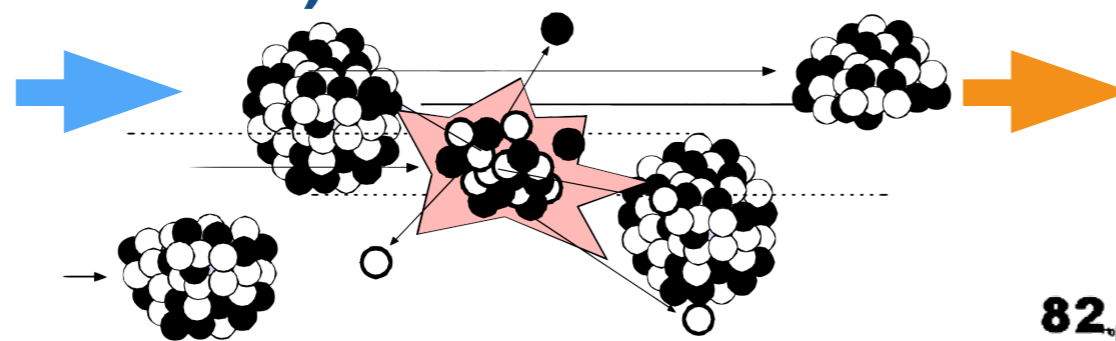
How can we study Unstable Nuclei?

Experimental Approach to Unstable Nuclei

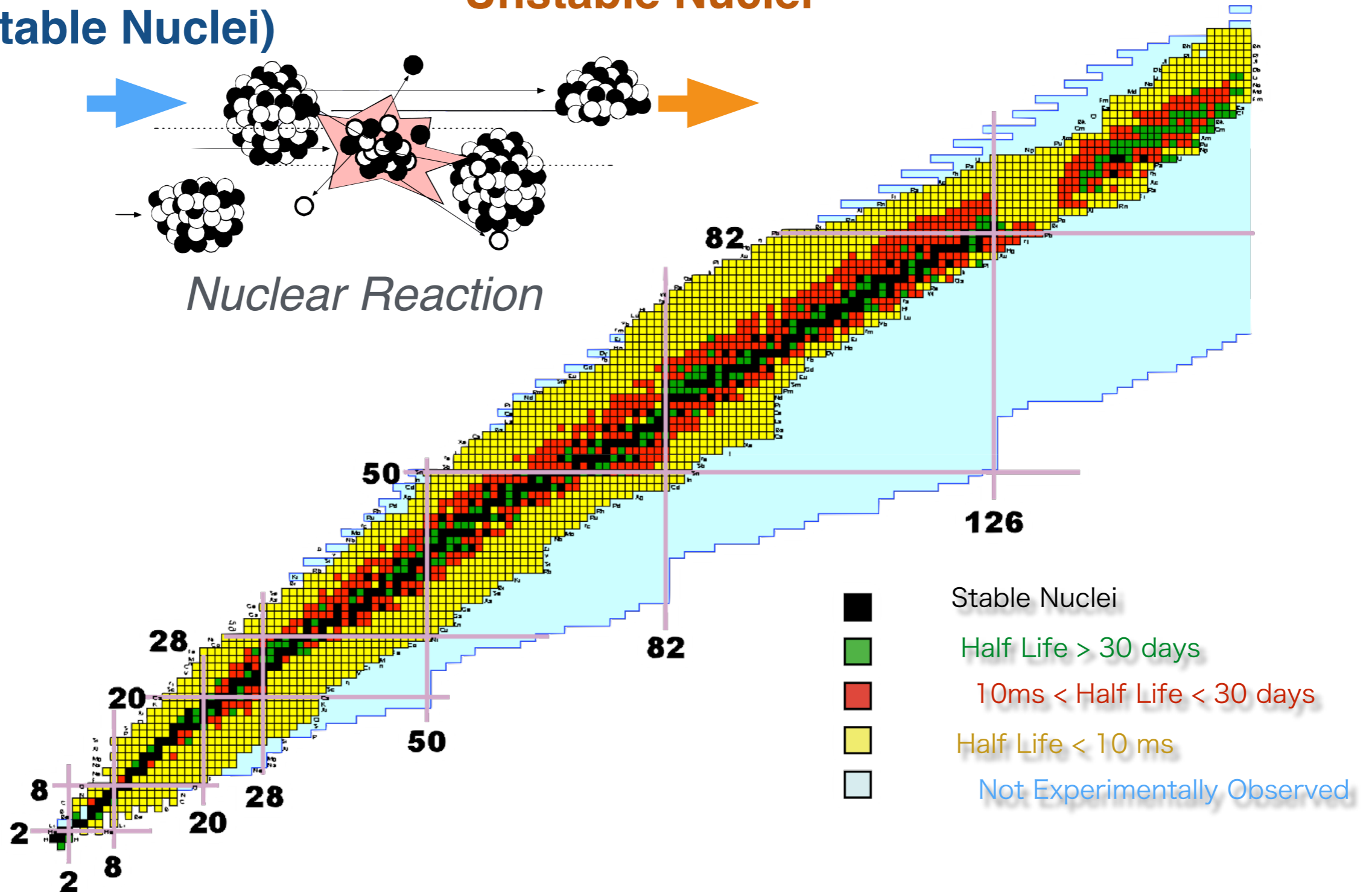
Accelerator

Heavy Ion Beam
(Stable Nuclei)

Unstable Nuclei



Nuclear Reaction



Accelerator

What is the accelerator?

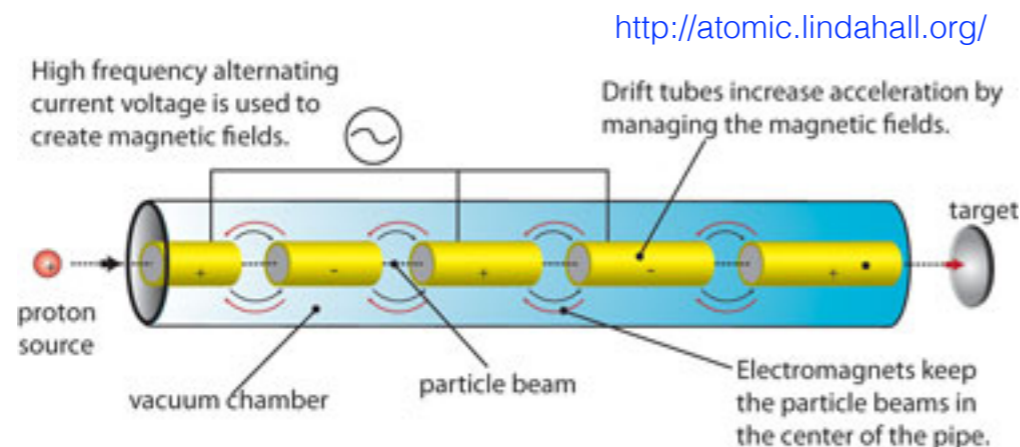
The machine which give certain kinetic energy to charged particles by accelerating and controlling them with the use of electro magnetic field.

Unit of Energy

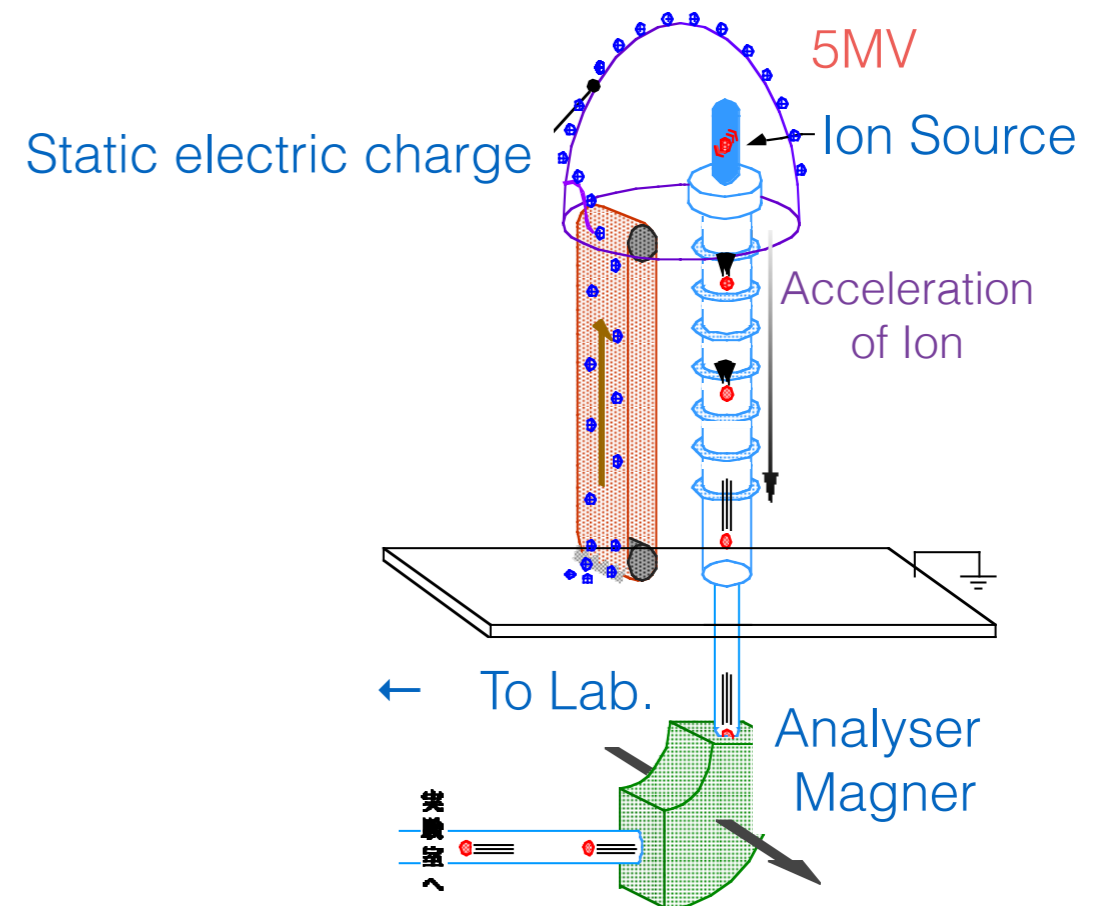
eV : Amount of energy gained by the charge of a single electron (1.6×10^{-19}) accelerated by an electric potential difference of 1V.

$$1 \text{ eV} = 1.6 \times 10^{-19} \text{ J}$$

- Electrostatic Accelerator
- Linear Accelerator



- Circular Accelerator



Van de Graaff

Circular Accelerator

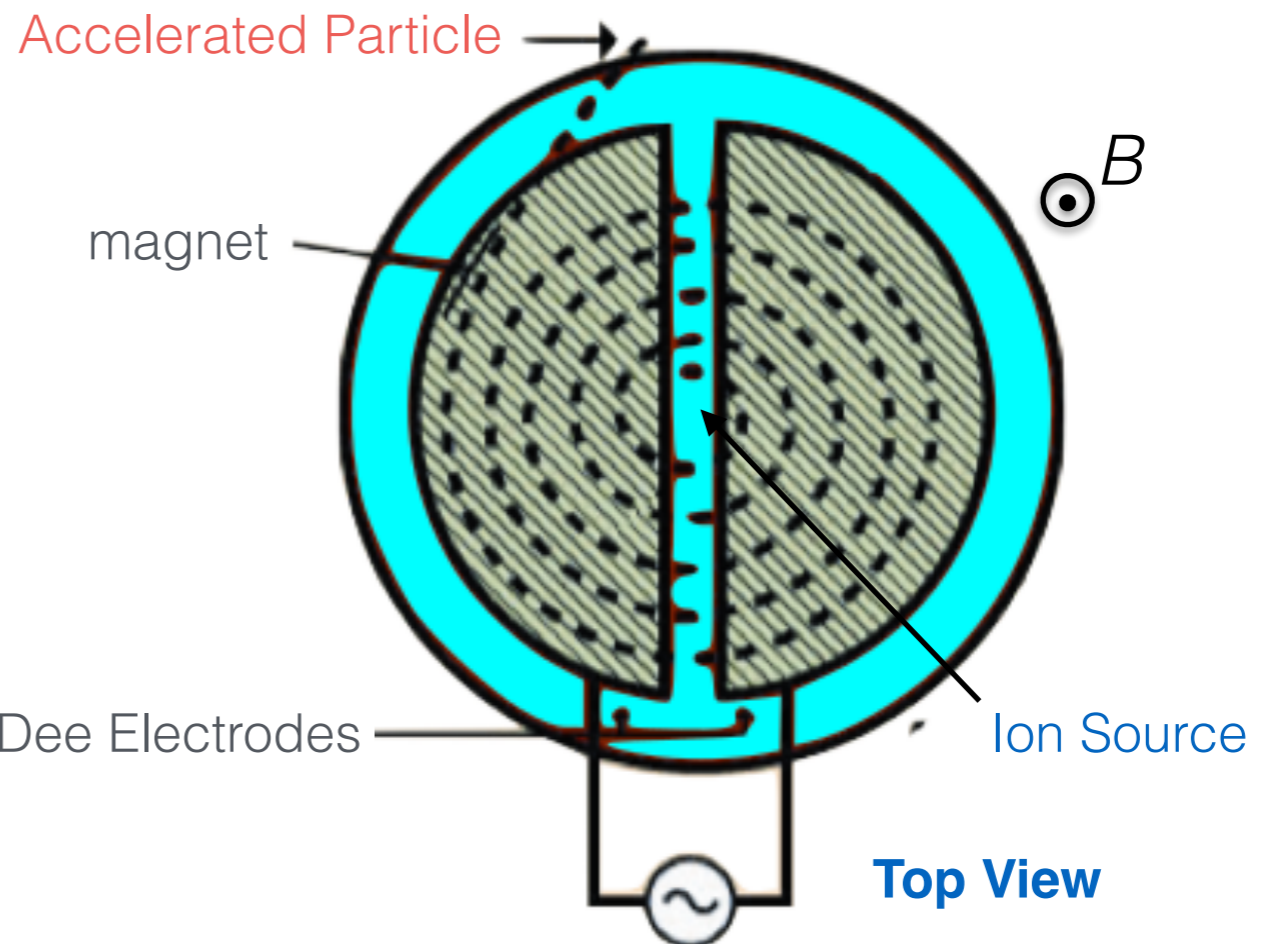
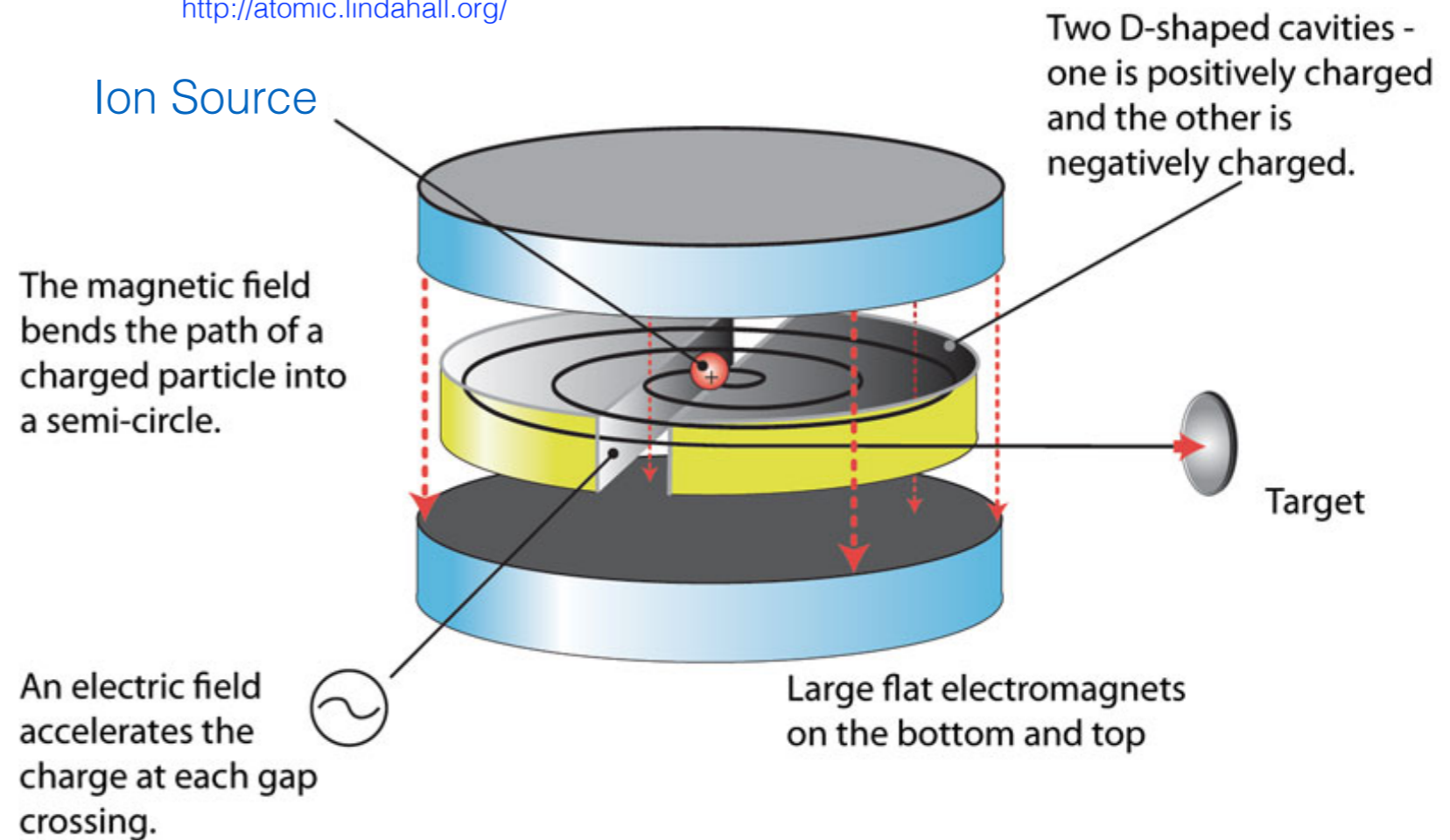
<http://atomic.lindahall.org/>

$$qvB = \frac{mv^2}{r}$$

$$r = \frac{mv}{qB}$$

$$T = \frac{2\pi m}{qB}$$

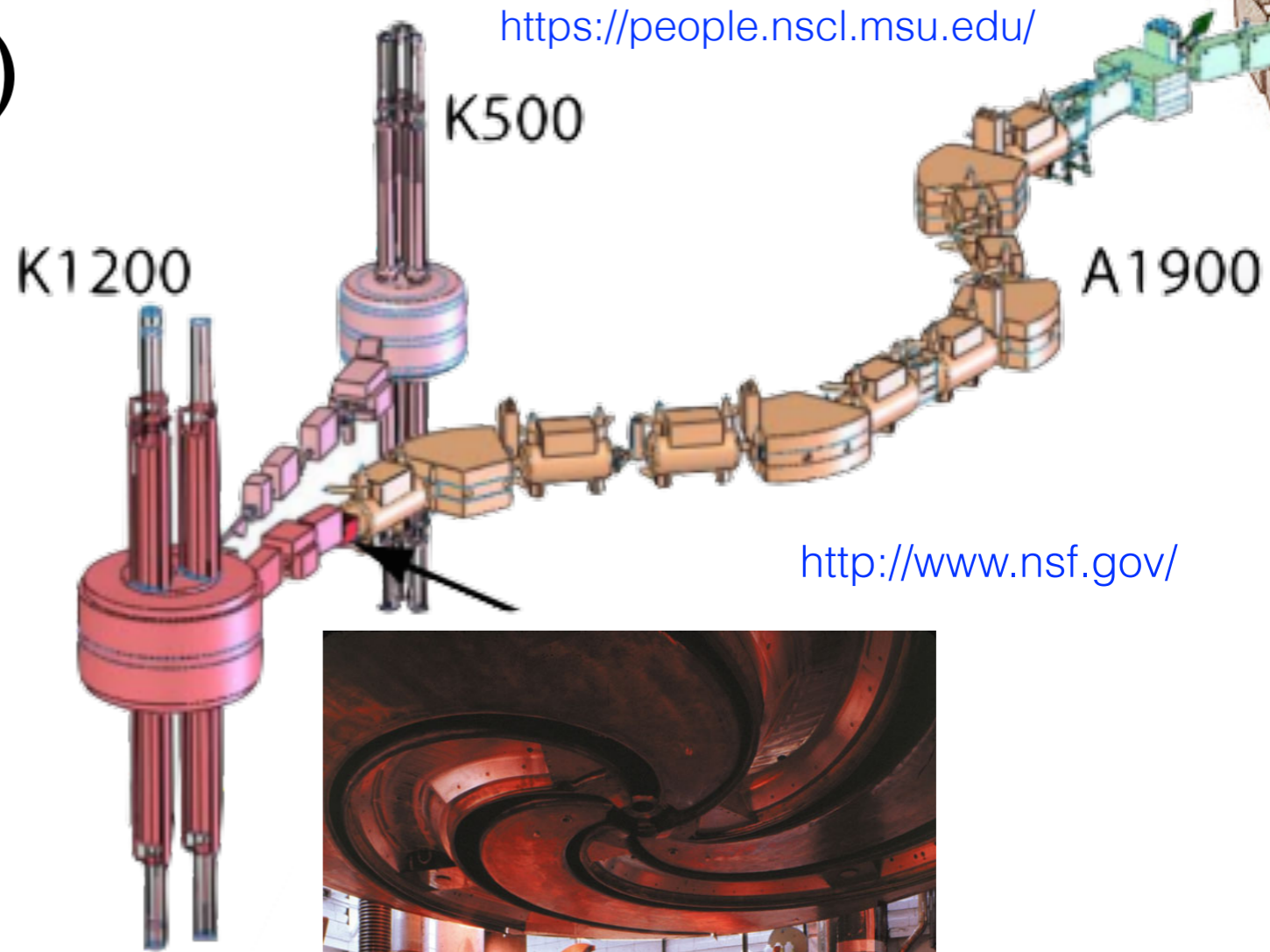
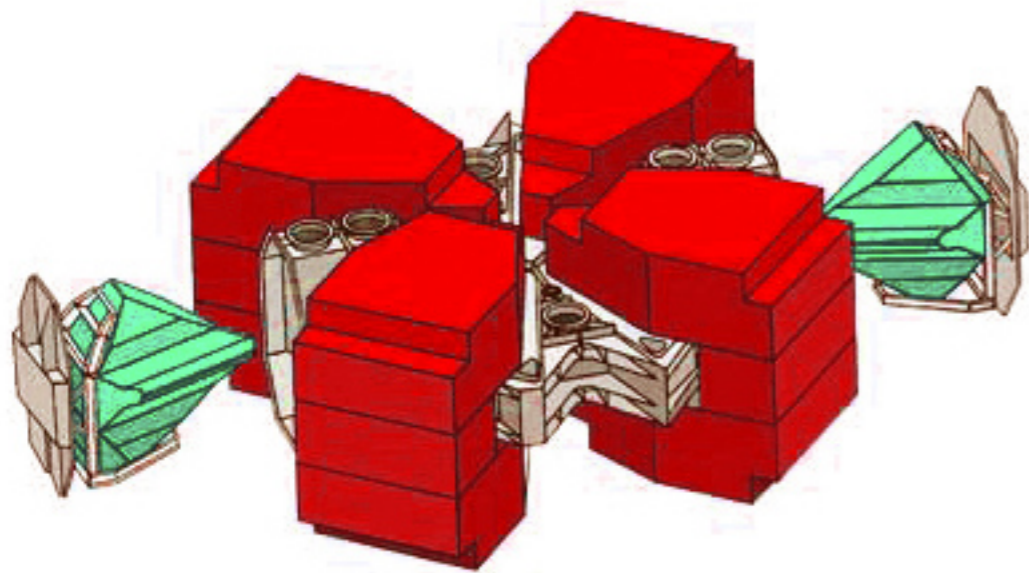
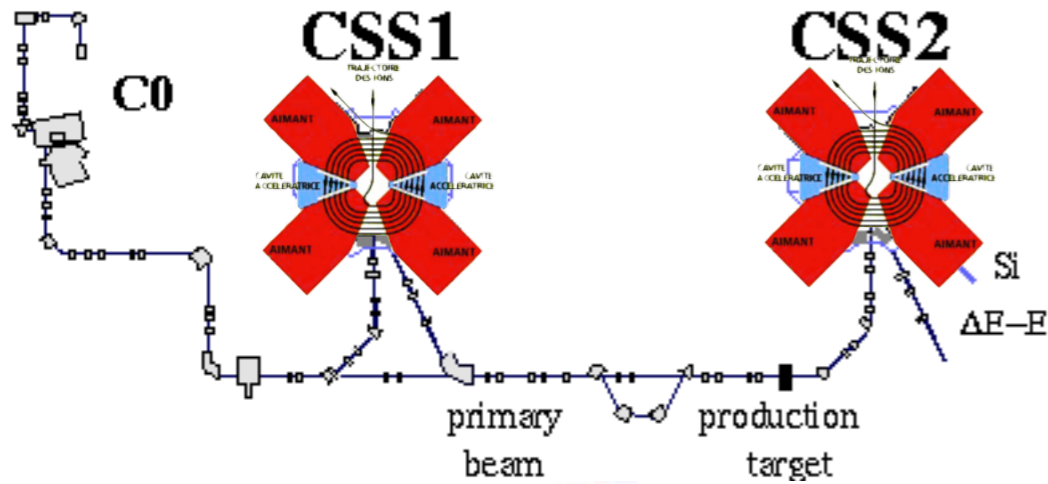
T is always constant,
independent of
particles velocity



Cyclotron Facilities

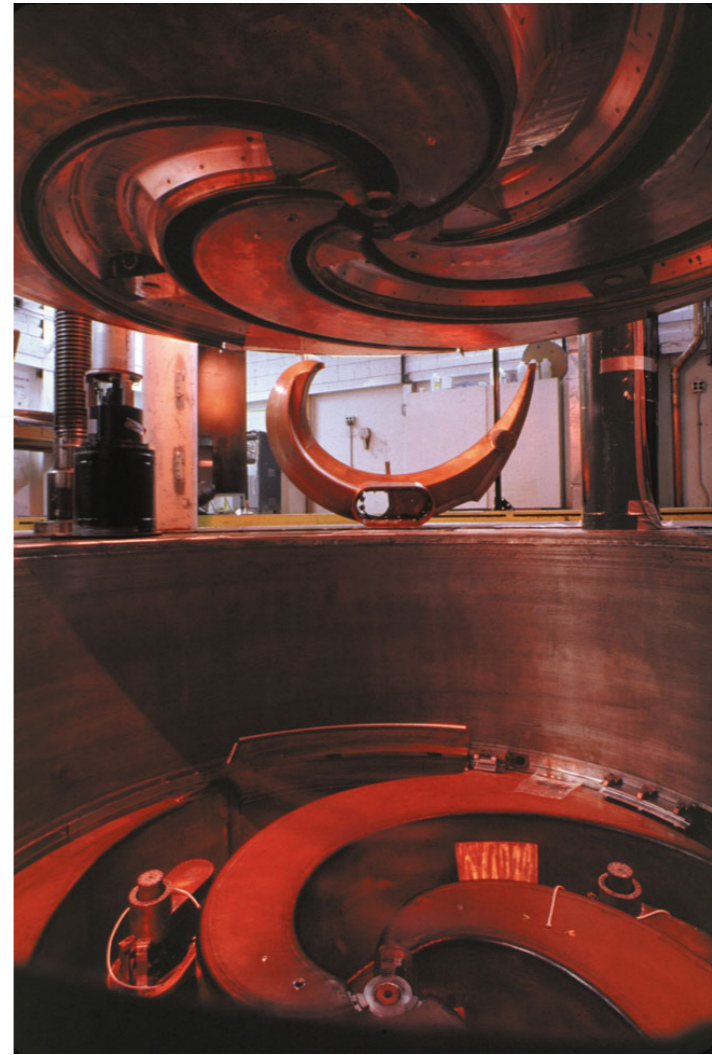
NSCL K500, K1200 (USA)

GANIL CSS1, CSS2



<https://people.nsl.msui.edu/>

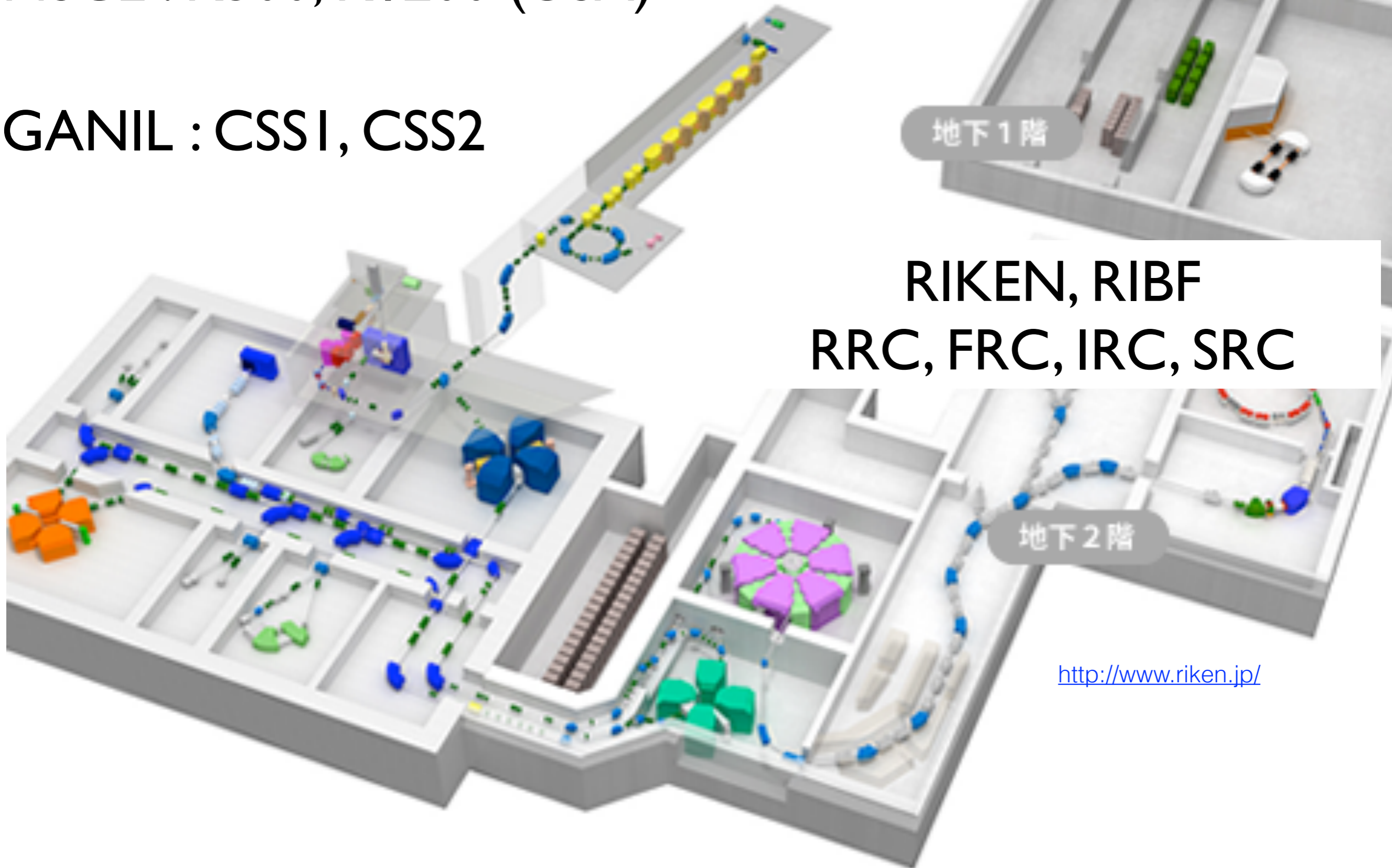
<http://www.nsf.gov/>



Cyclotron Facilities

NSCL : K500, K1200 (USA)

GANIL : CSS1, CSS2



地下1階

RIKEN, RIBF
RRC, FRC, IRC, SRC

地下2階

<http://www.riken.jp/>



Synchrotron Accelerator

At relativistic energy

$$qvB = \frac{mv^2}{r}$$

$$T = \frac{2\pi m}{qB}$$

~~*T is always constant, independent of particles velocity*~~

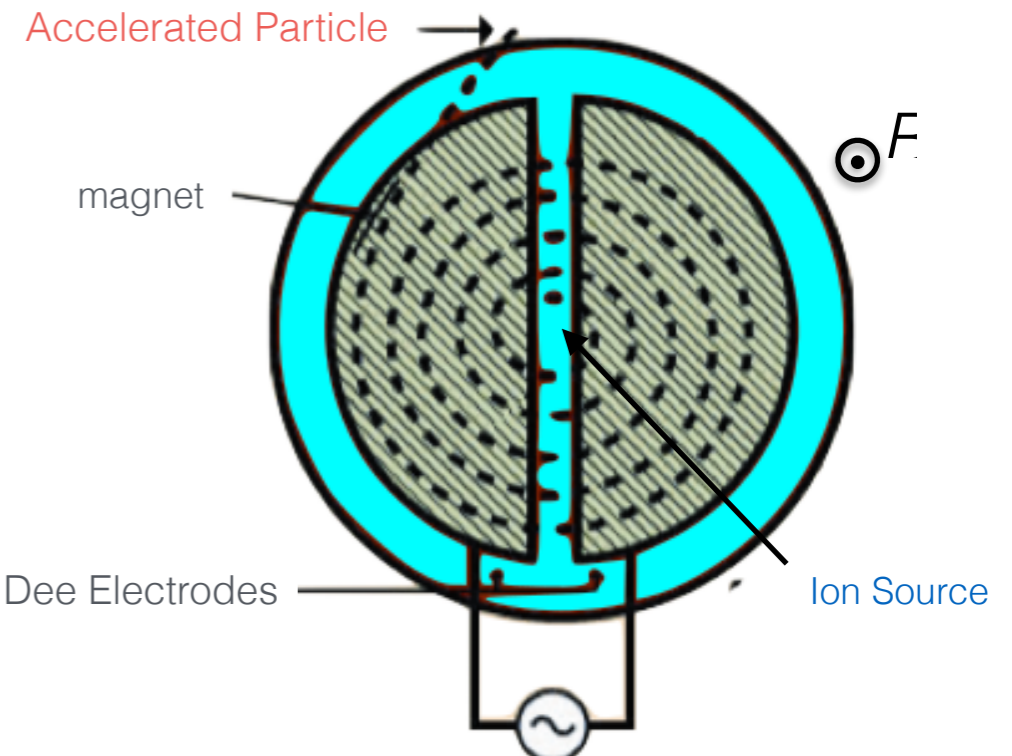
A bunch of Beam

Accelerating cavity

Electric field applied high and varying frequency to accelerate particles appropriate timing

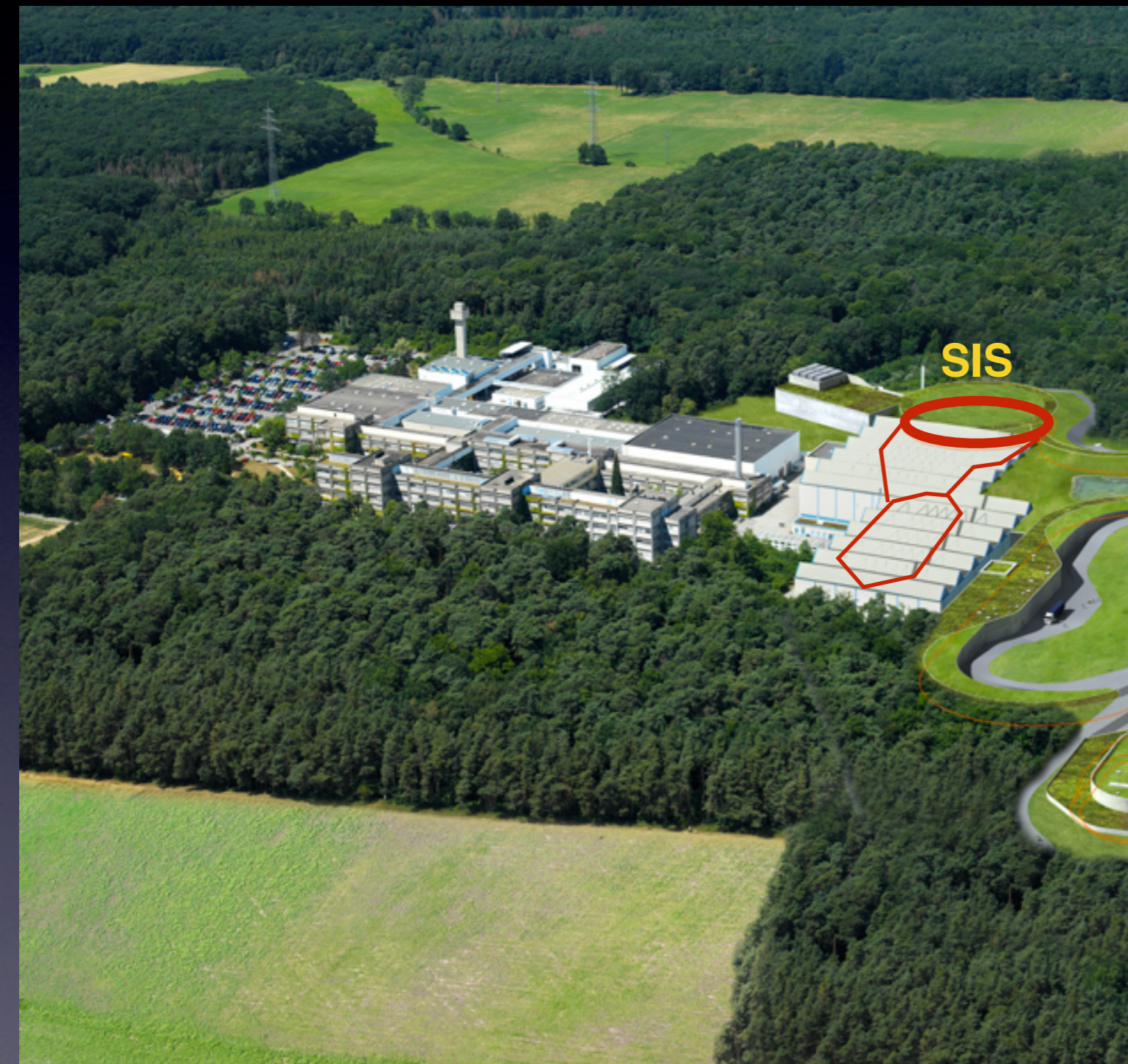
Particles are staged in the ring, and travel always on the same fixed orbit

Cyclotron : High **Intensity**
Synchrotron : High **Energy**

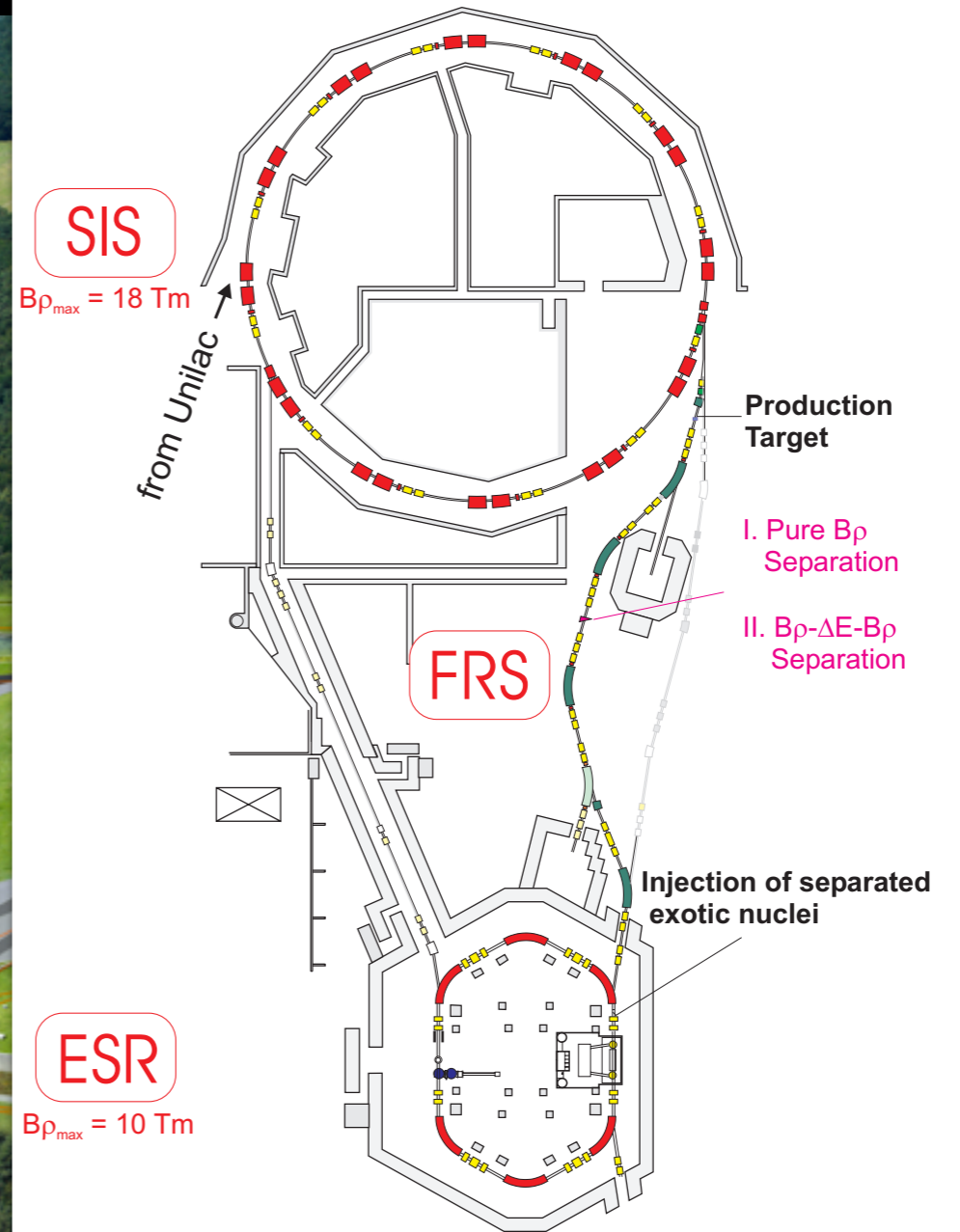


Cyclotron : Continuous Acceleration

• GSI, SIS



Mass and Lifetime Measurements of Stored Exotic Nuclei



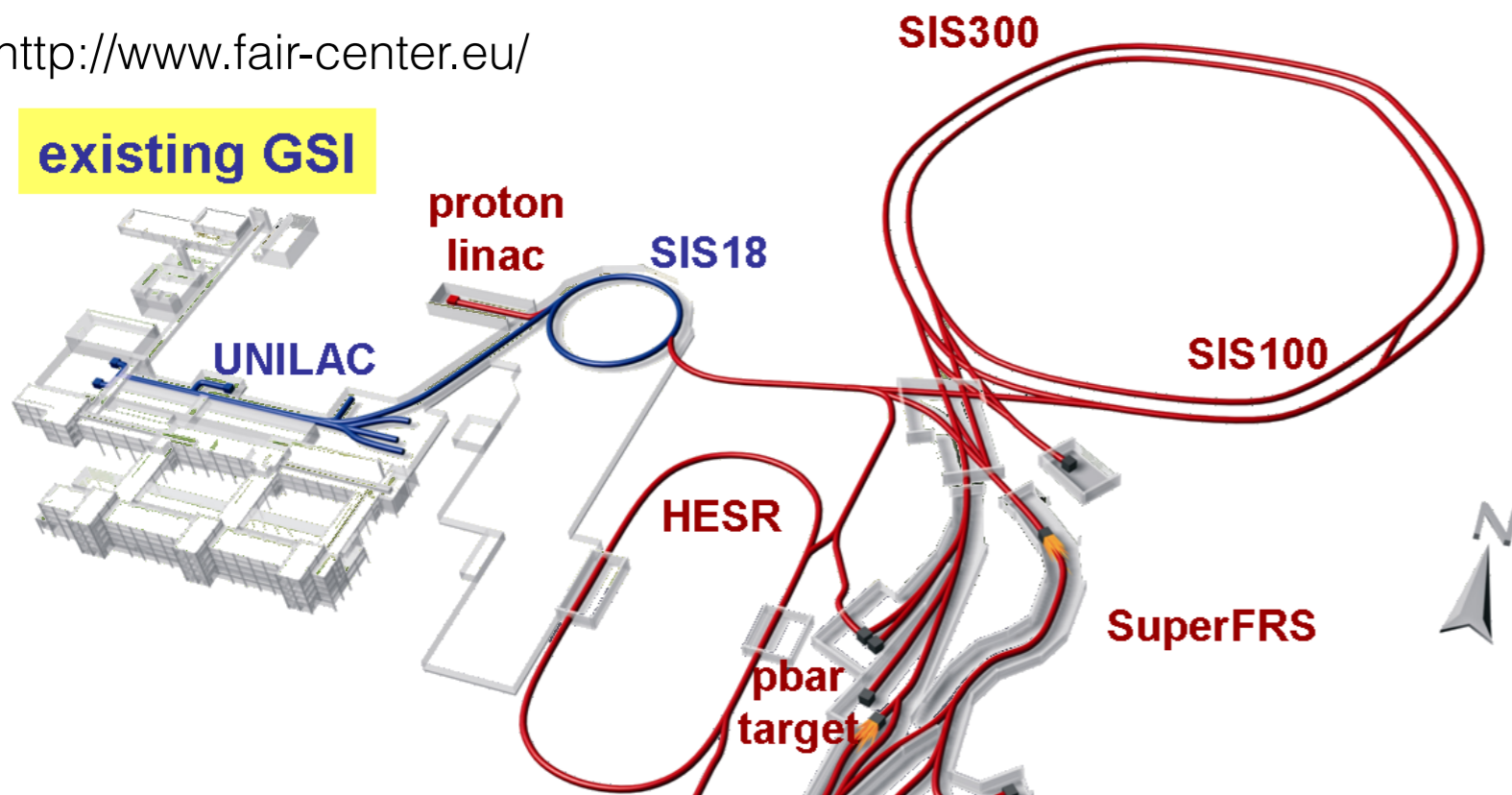
F. Bosch, Lect. Notes Phys. 651, 137 (2004)

- GSI, SIS and then SIS100, SIS300



<http://www.fair-center.eu/>

existing GSI

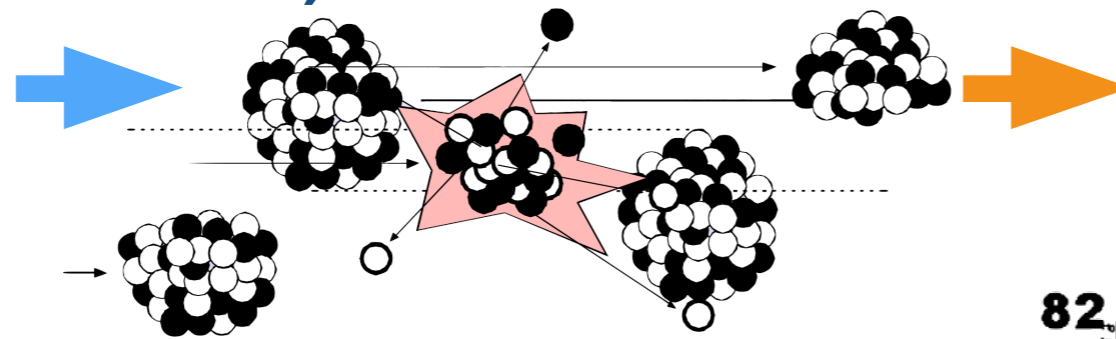


Experimental Approach to Unstable Nuclei

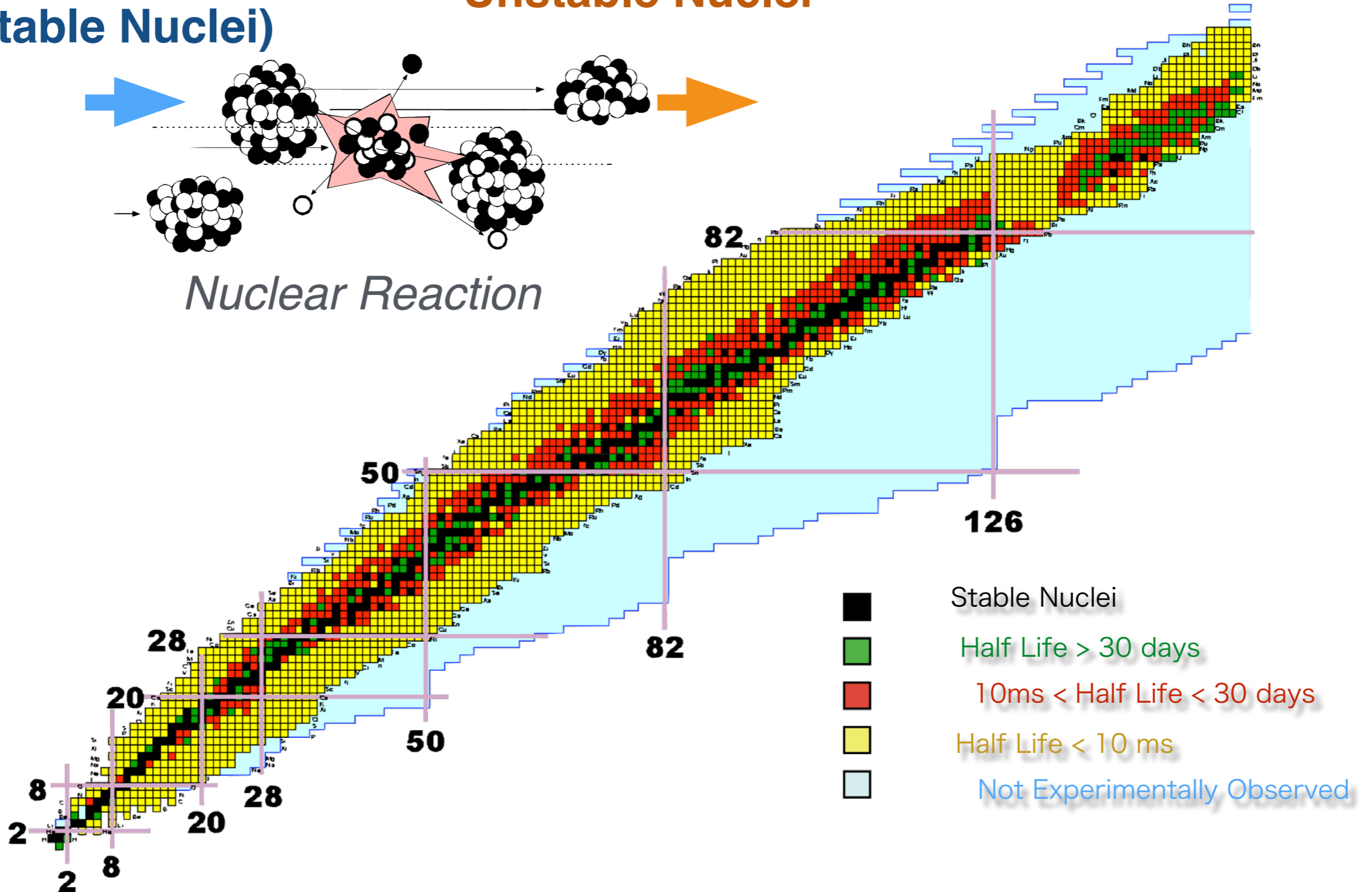
Accelerator

Heavy Ion Beam
(Stable Nuclei)

Unstable Nuclei



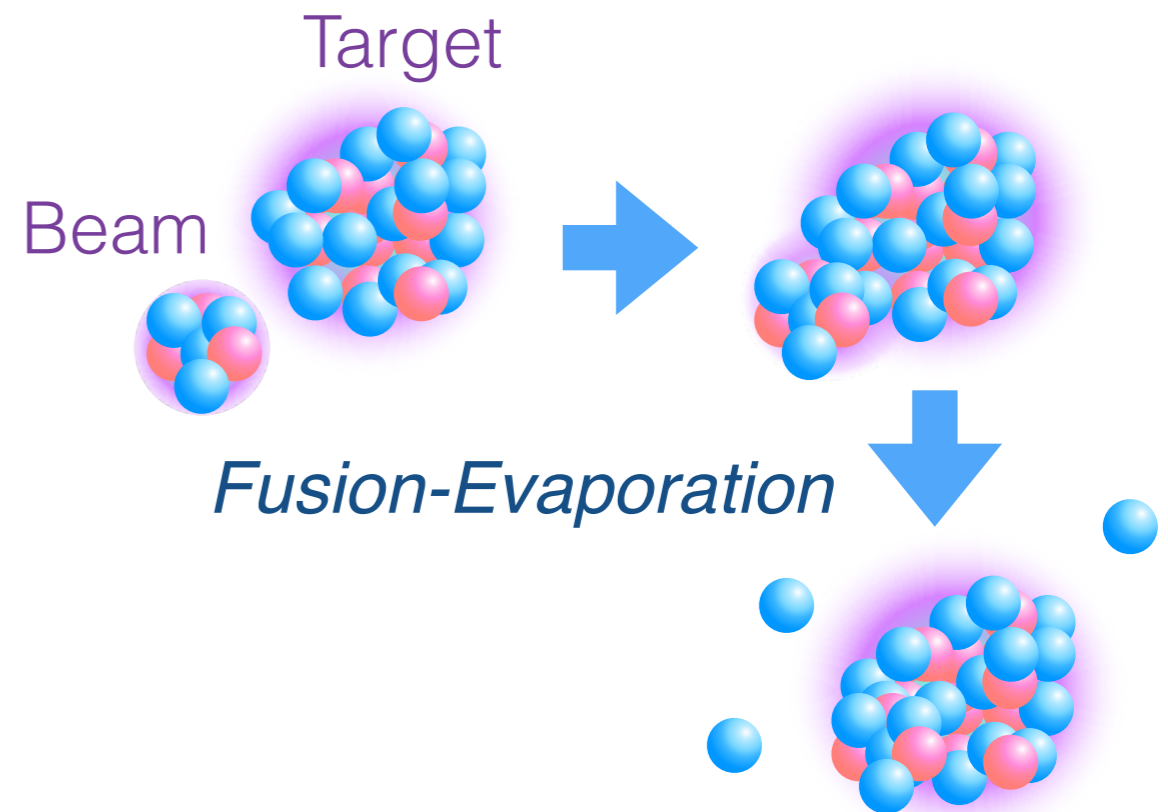
Nuclear Reaction



Production of Unstable Nuclei as Radioactive Isotope (RI) Beam

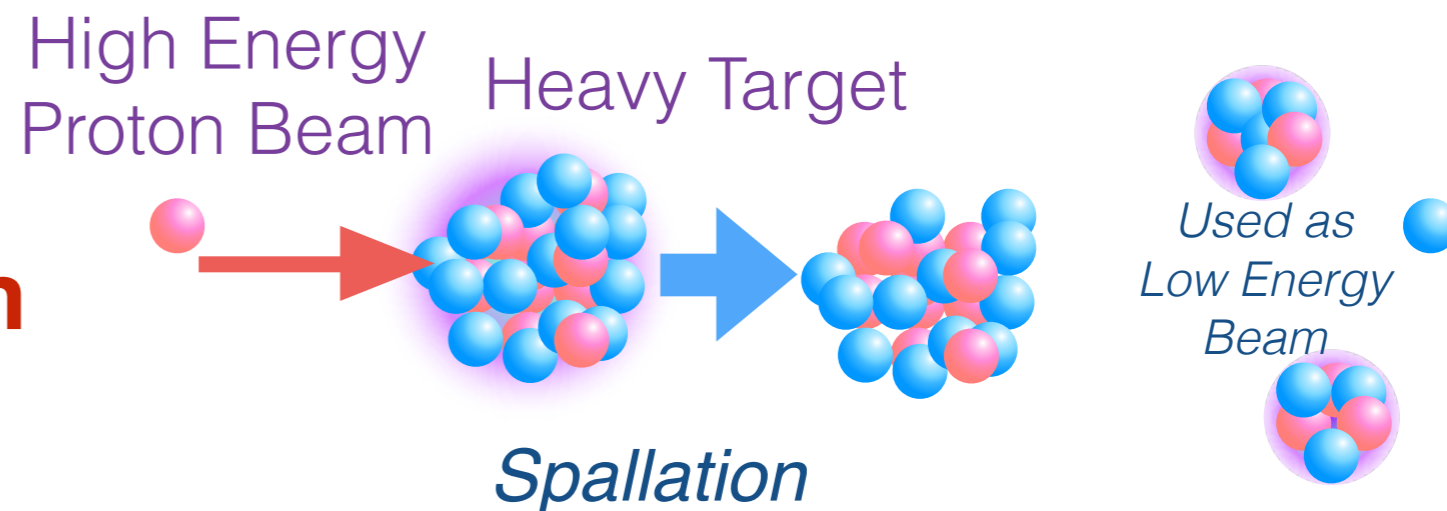
Production of Unstable Nuclei

- Fusion-evaporation
- Fission
- Transfer Reaction



Production of **Unstable Nuclei Beam**

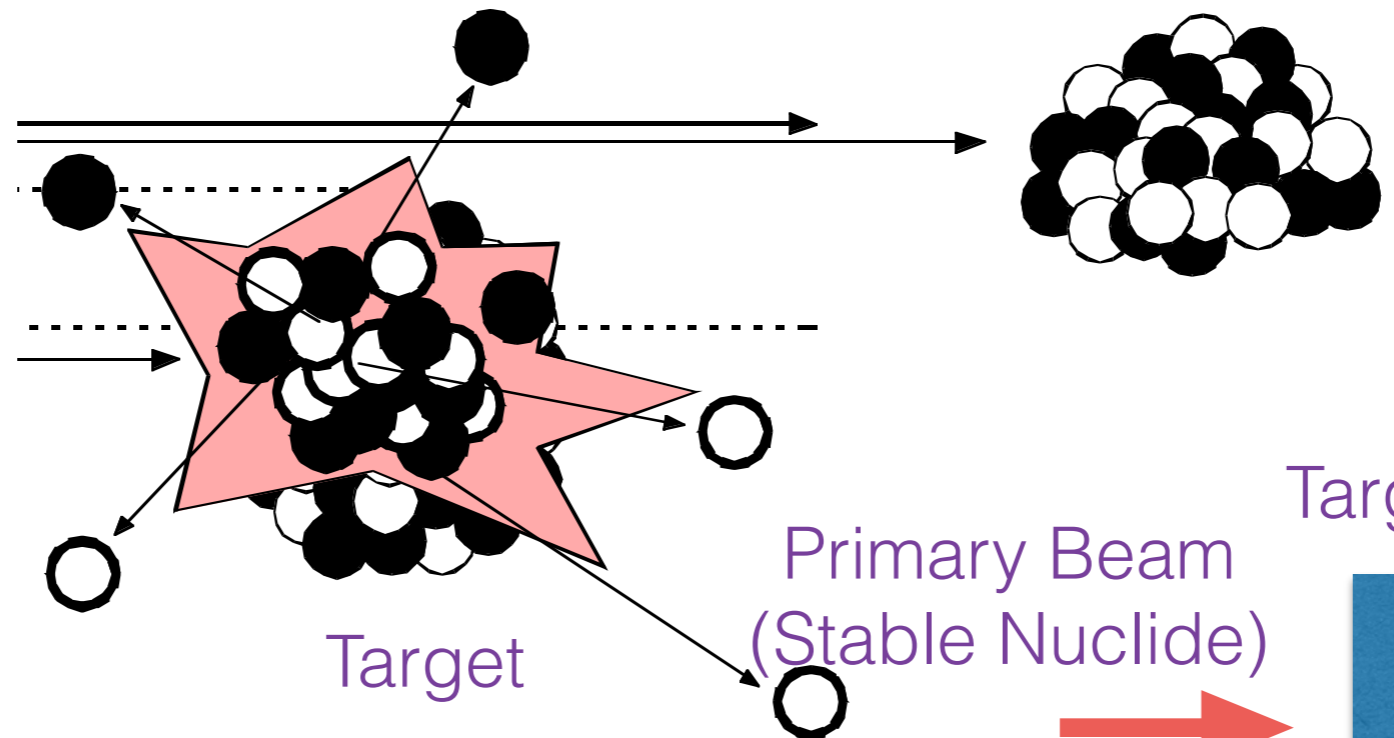
- Spallation → ISOL
- **Projectile Fragmentation**
- **Fission Ablation**



Primary Beam
(Stable Nuclide)

Projectile Fragmentation

Secondary Beam
(Unstable Nuclide)



Primary Beam
(Stable Nuclide)

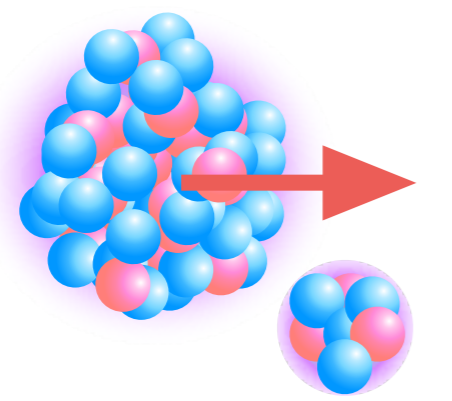
Target

Secondary Beam
(Unstable Nuclide)

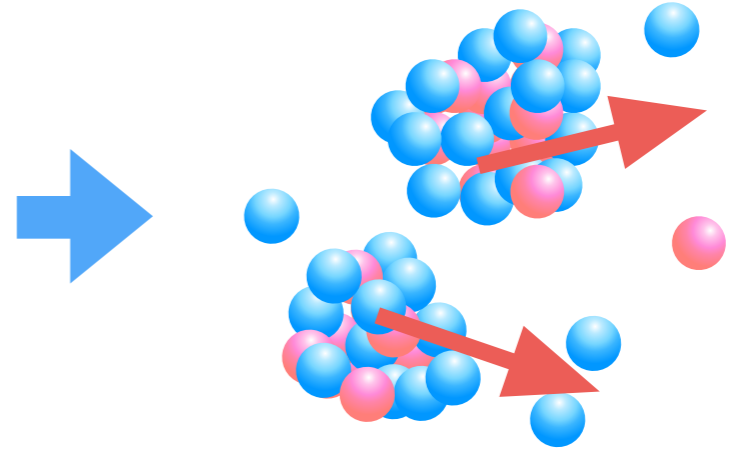
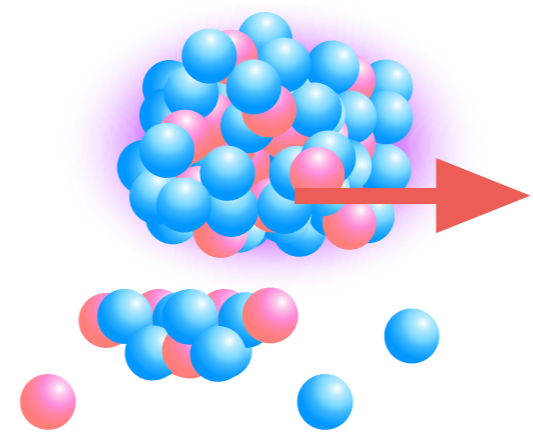
^{238}U Beam

Fissile Nucleus

Fission Fragments

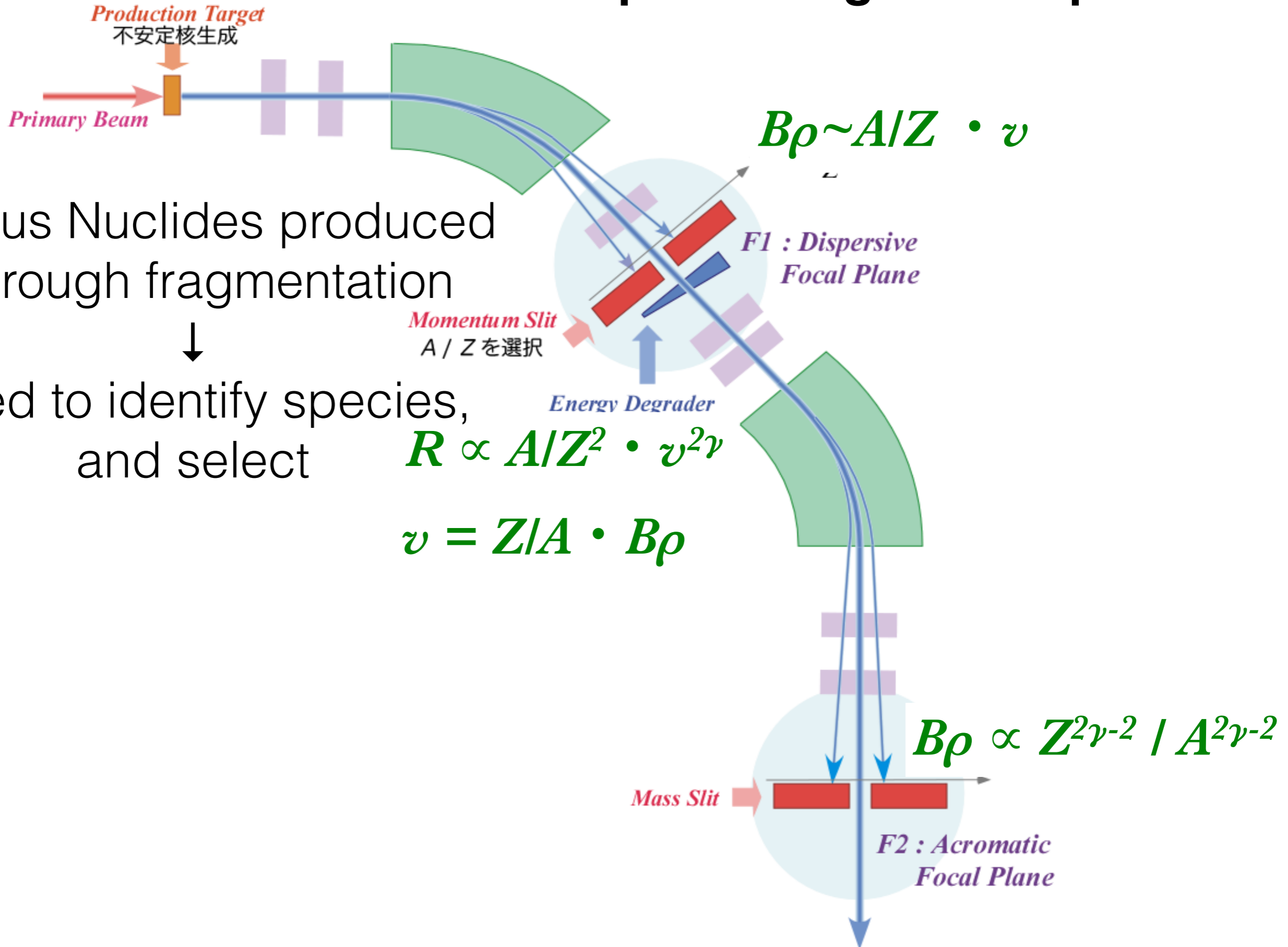


Target



Fission Abrasion

Principle of Fragment Separator

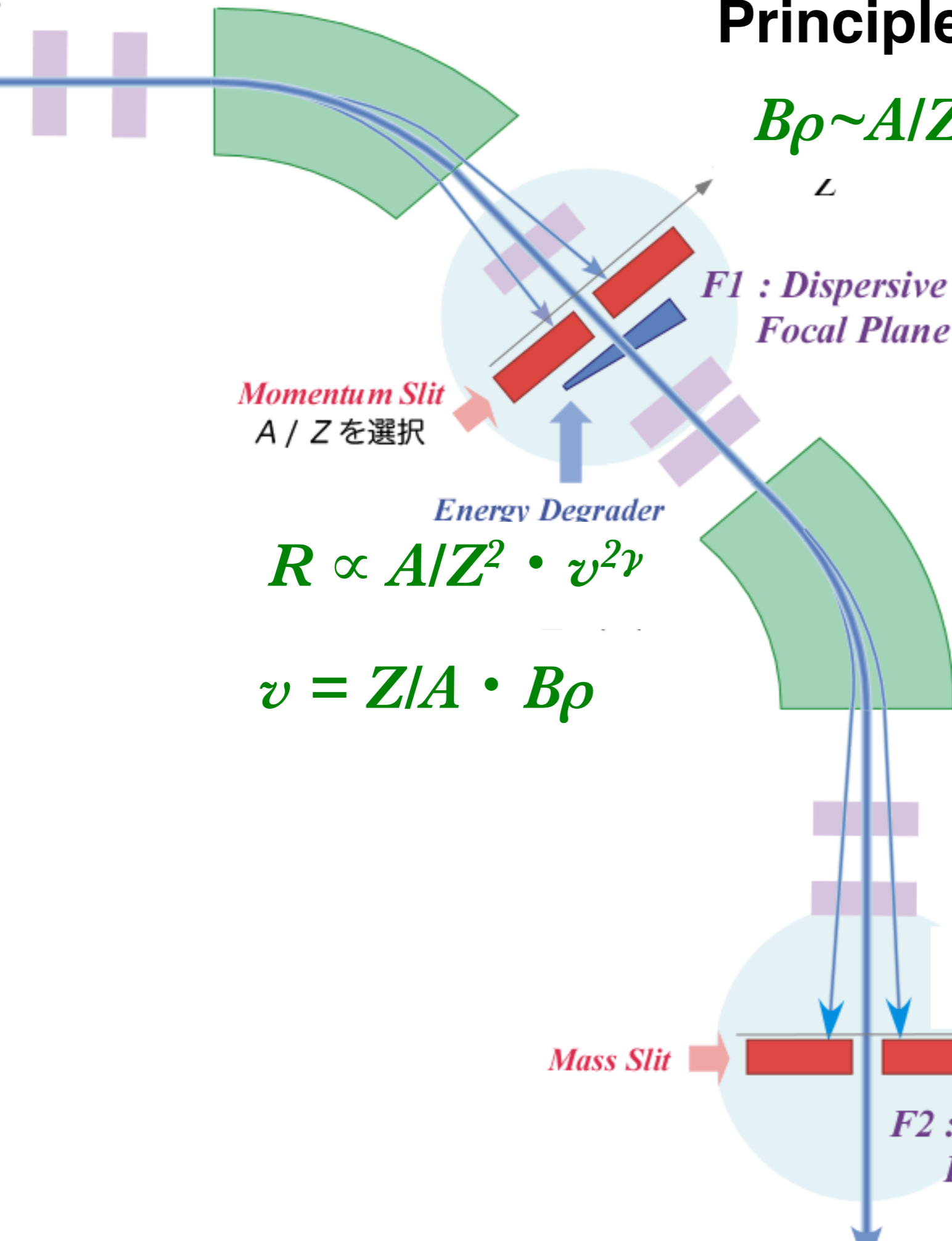


Various Nuclides produced thorough fragmentation



Need to identify species, and select

Principle of Fragment Separator



$$B\rho \sim A/Z \cdot v$$

L

For the more precise Particle Identification

Detectors

Plastic Scintillation Counter
Time Of Flight (TOF) : v

Position Sensitive Detector
Beam Position : $B\rho$

Energy-Loss (ΔE) Detector
 $\Delta E \propto Z^2/v^2$

$$B\rho \propto Z^2 v^{-2} / A^2 v^{-2}$$

Momentum Slit
A / Z を選択

Energy Degradator

F1 : Dispersive Focal Plane

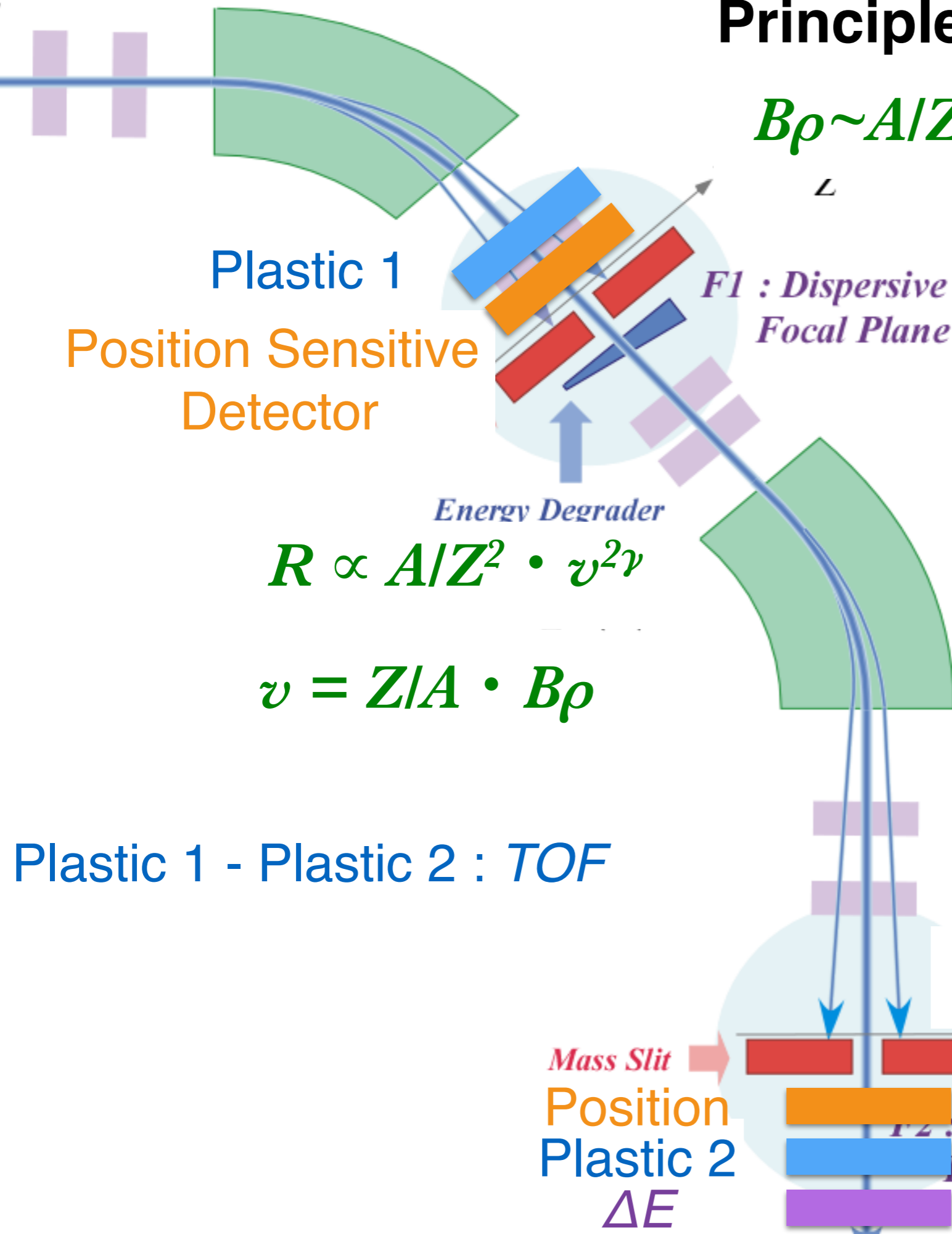
$$R \propto A/Z^2 \cdot v^2 \gamma$$

$$v = Z/A \cdot B\rho$$

Mass Slit

F2 : Acromatic Focal Plane

Principle of Fragment Separator



$$B\rho \sim A/Z \cdot v$$

L

For the more precise Particle Identification

Detectors

Plastic Scintillation Counter
Time Of Flight (TOF) : v

Position Sensitive Detector
Beam Position : $B\rho$

Energy-Loss (ΔE) Detector
 $\Delta E \propto Z^2/v^2$

$$B\rho \propto Z^2 v^{-2} / A^2 v^{-2}$$

Plastic 1

Position Sensitive Detector

F1 : Dispersive Focal Plane

Energy Degradator

$$R \propto A/Z^2 \cdot v^2 \gamma$$

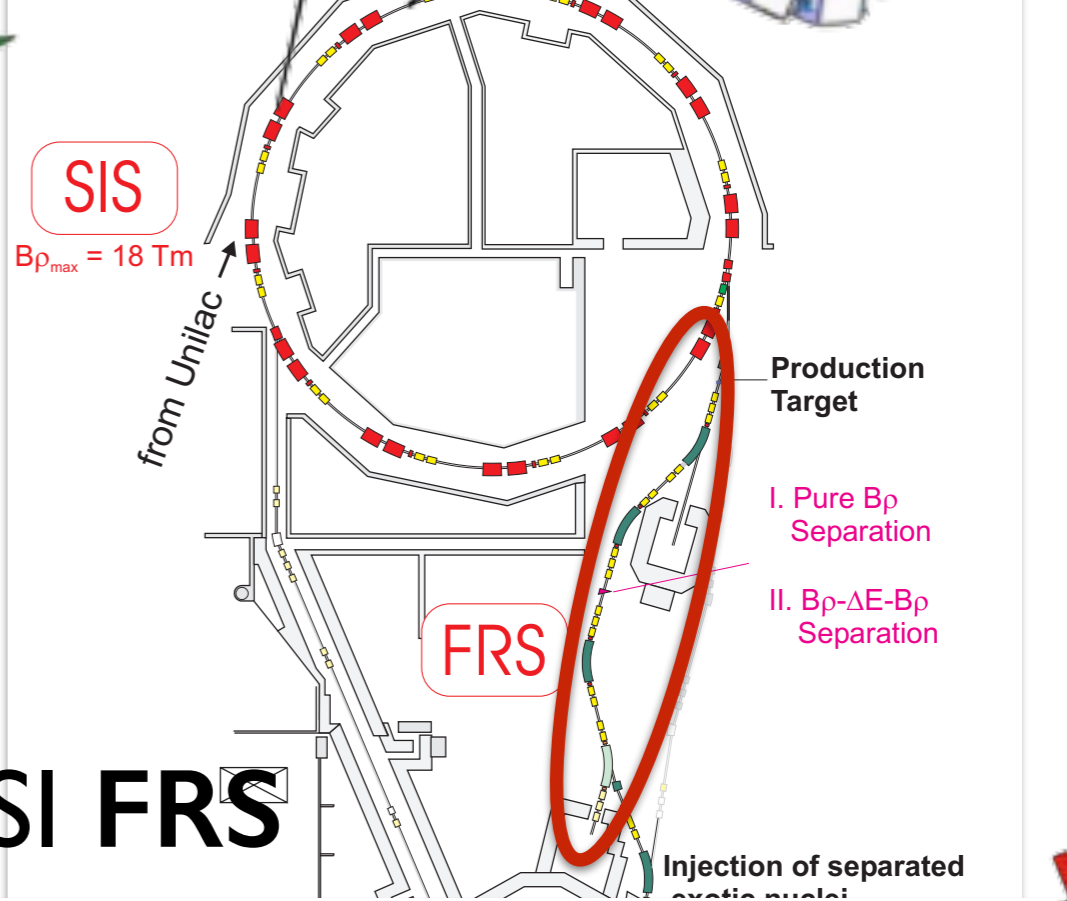
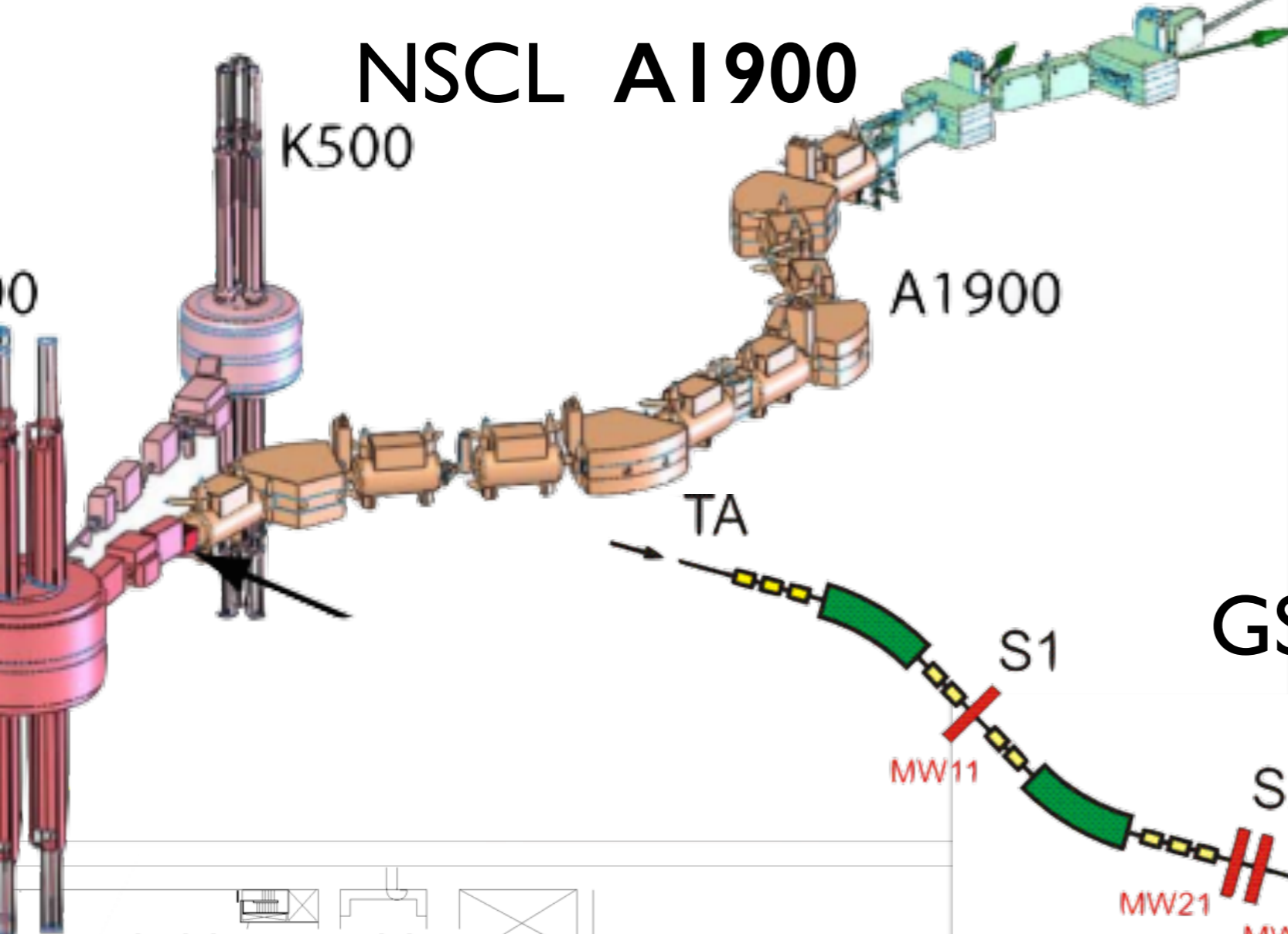
$$v = Z/A \cdot B\rho$$

Plastic 1 - Plastic 2 : TOF

Mass Slit
Position Sensitive Detector
Plastic 2
 ΔE

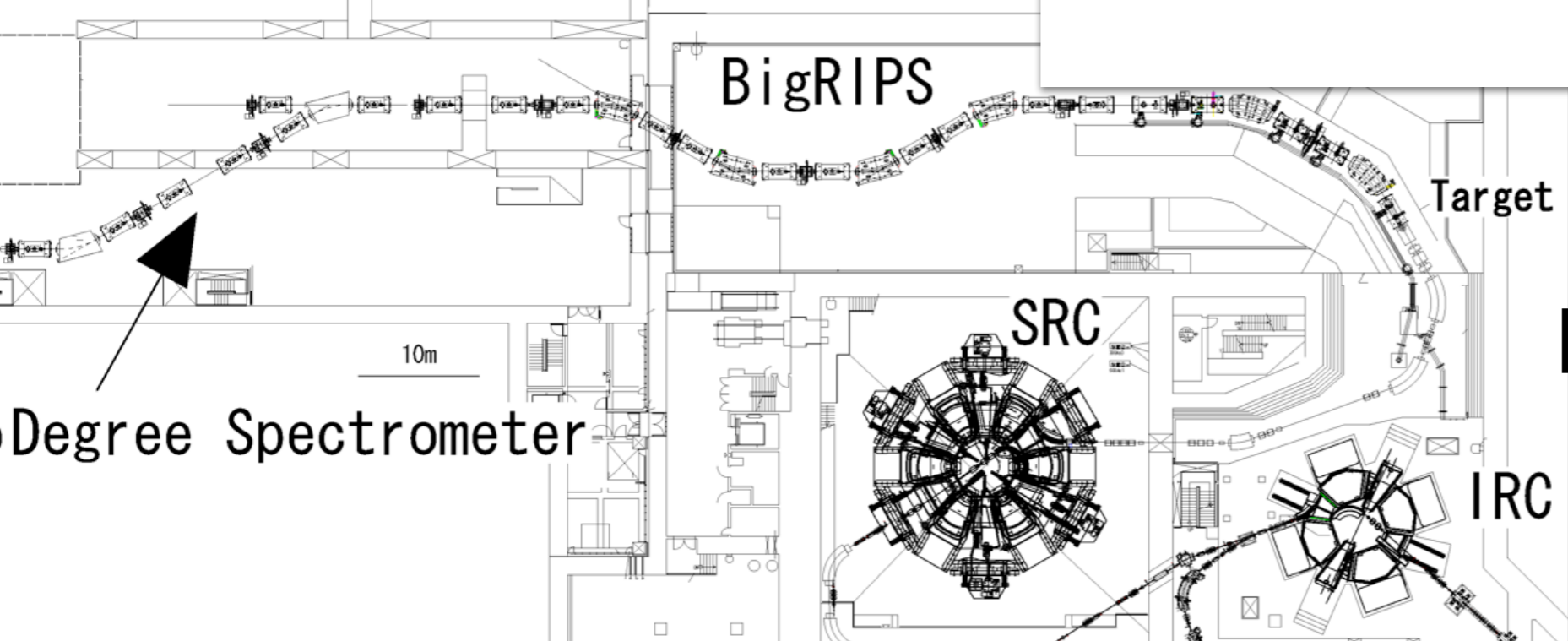
F2 : Acromatic Focal Plane

NSCL A1900



GSI FRS

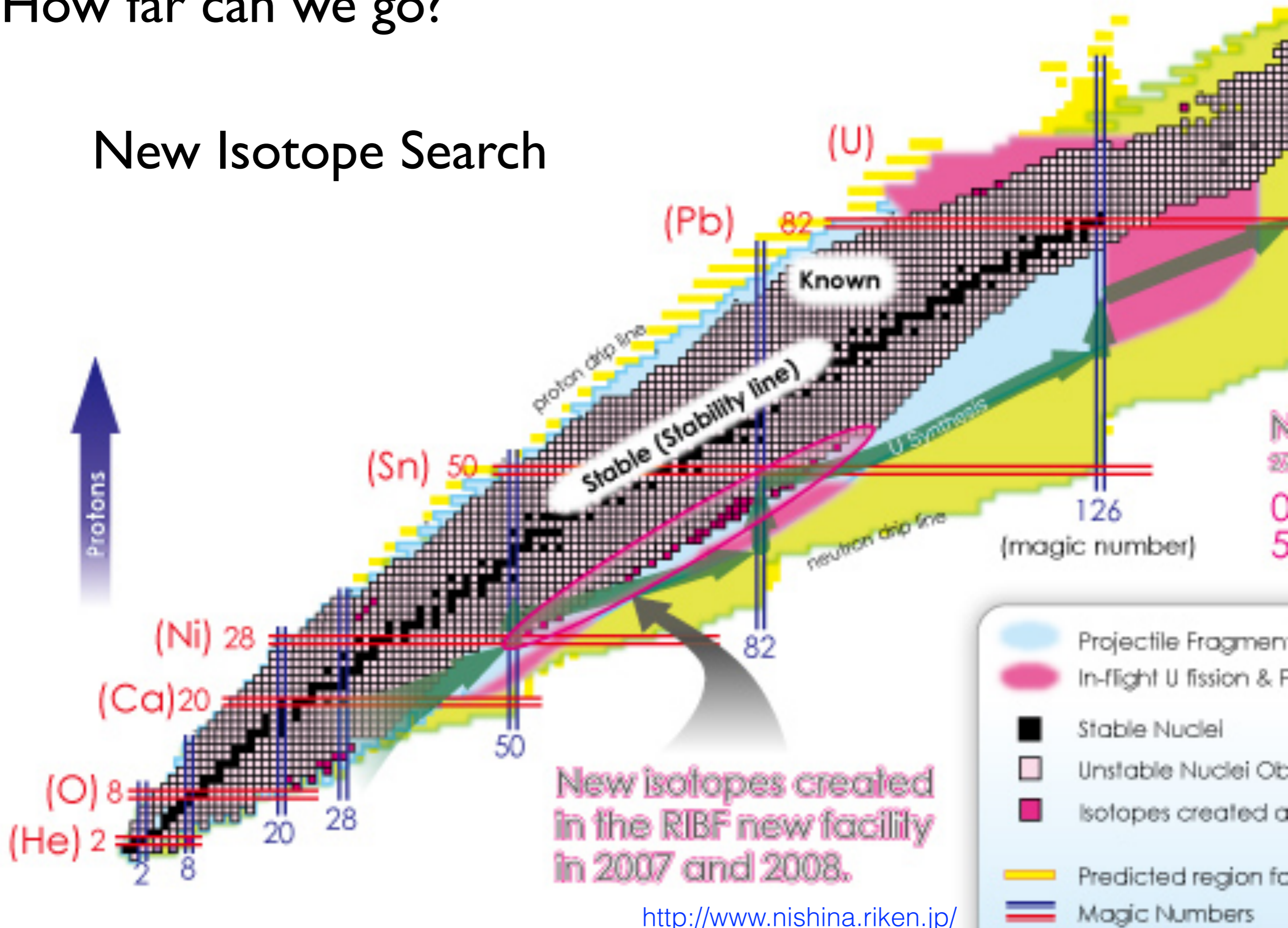
beam delivery line



RIKEN BigRIPS

How far can we go?

New Isotope Search

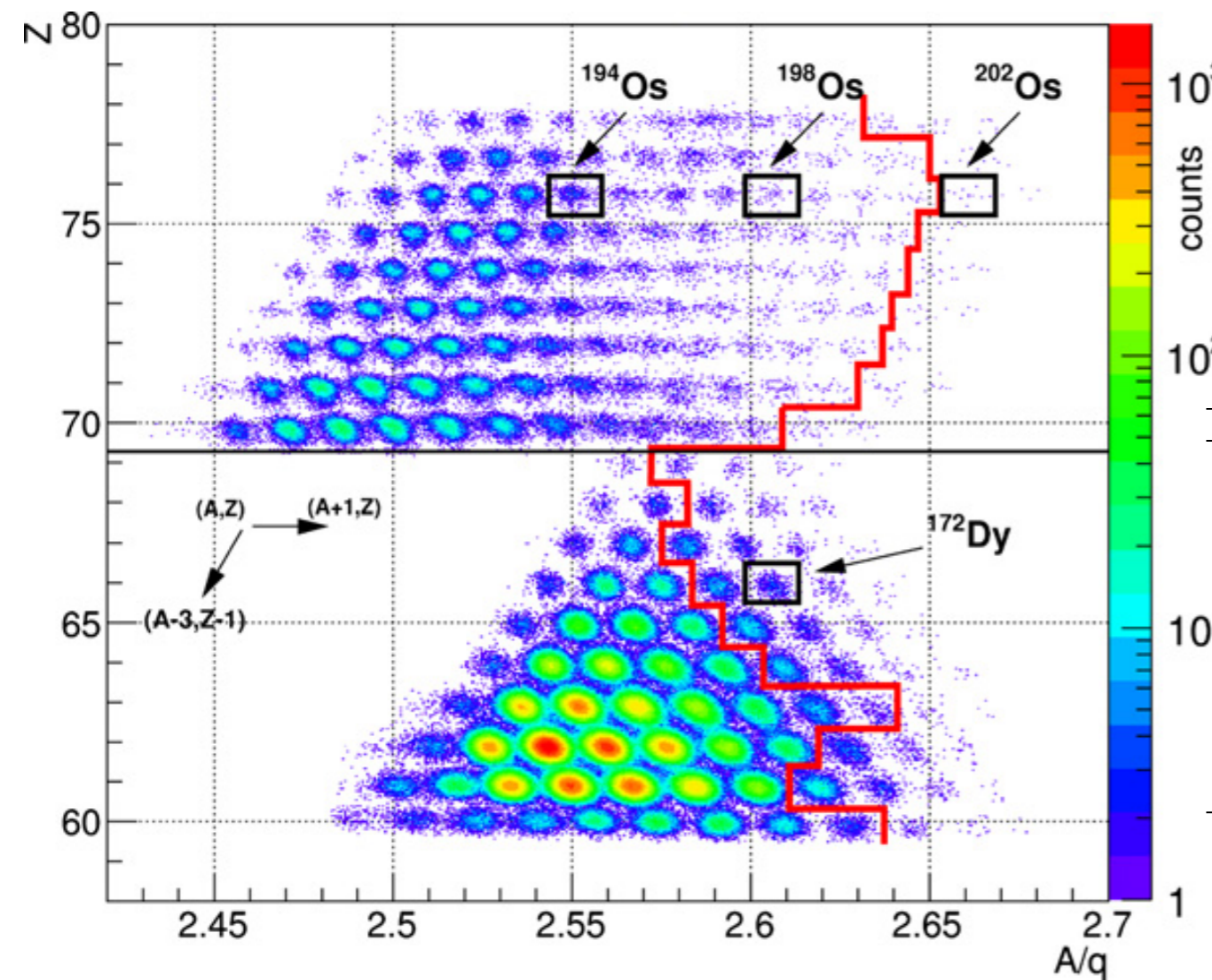
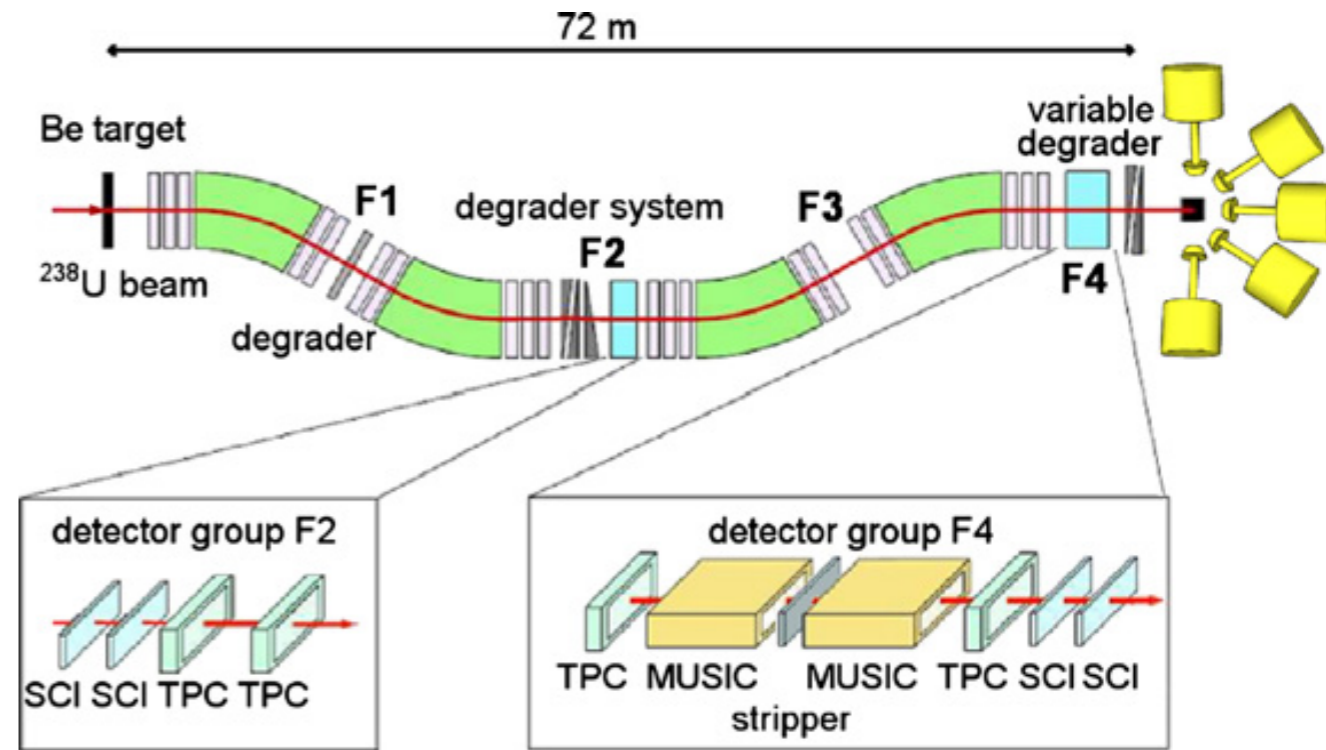


Recent Works for New Isotope Search

J. Kurcewicz et al., Phys. Lett. B 717 (2012) 371–375

GSI (2012)

^{238}U 1 GeV/u on Be Target fission abrasion



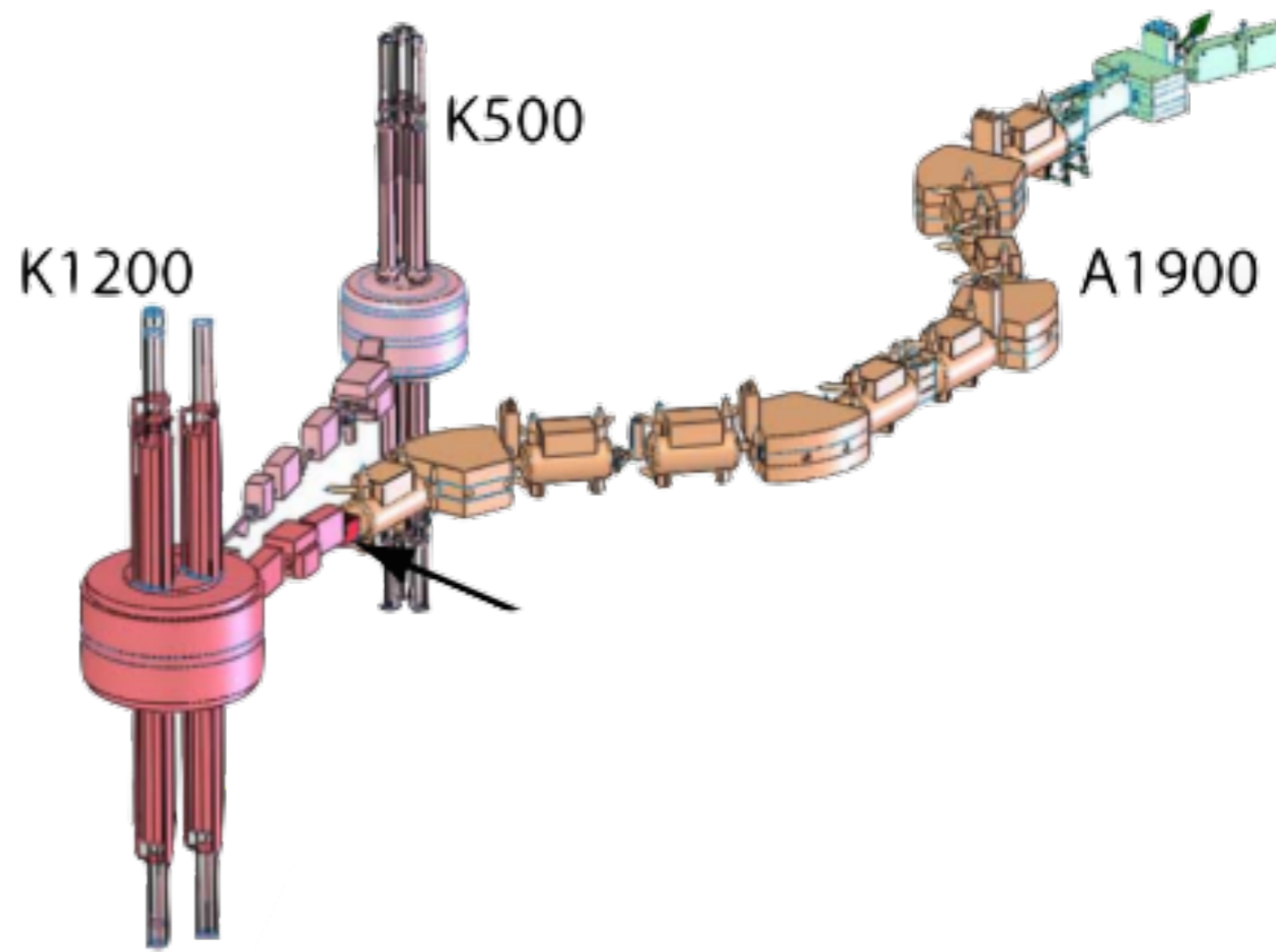
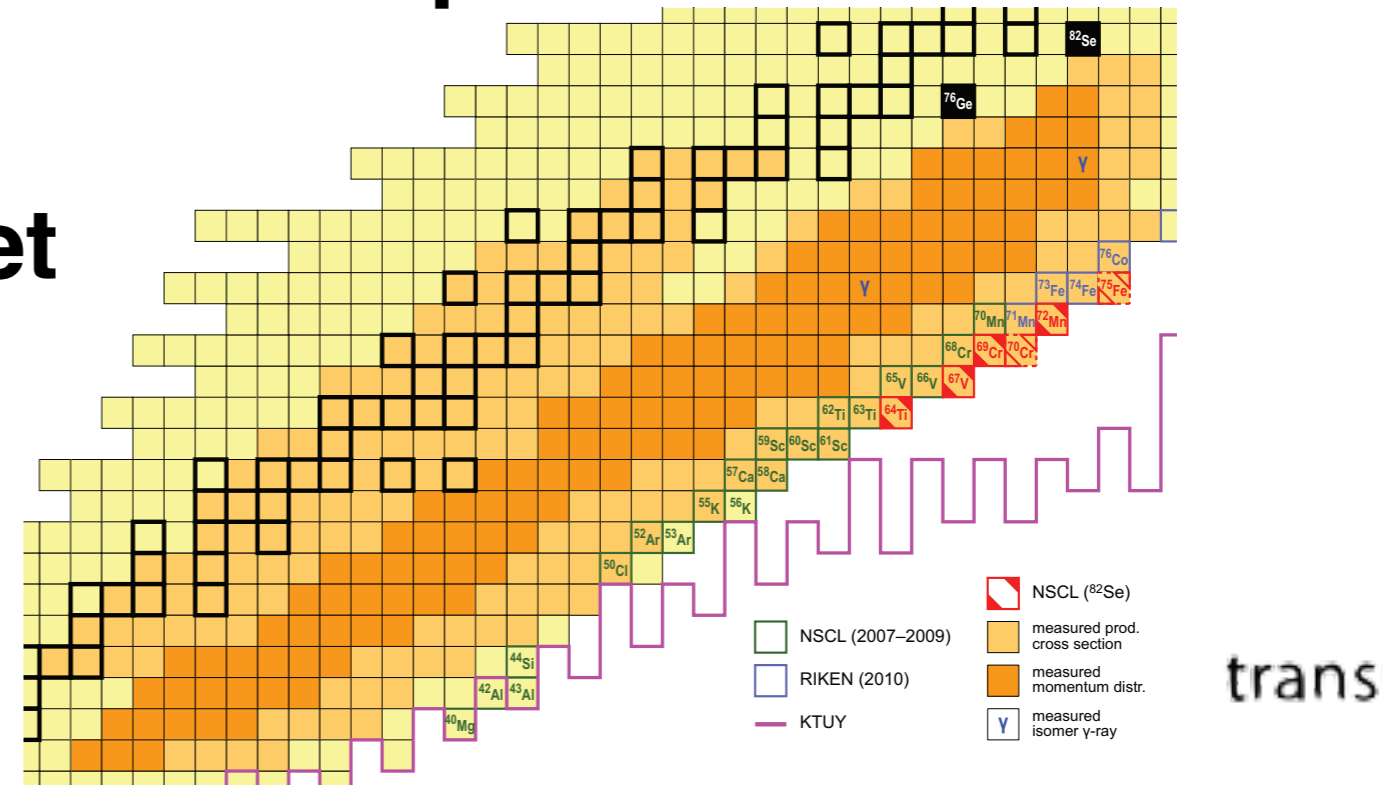
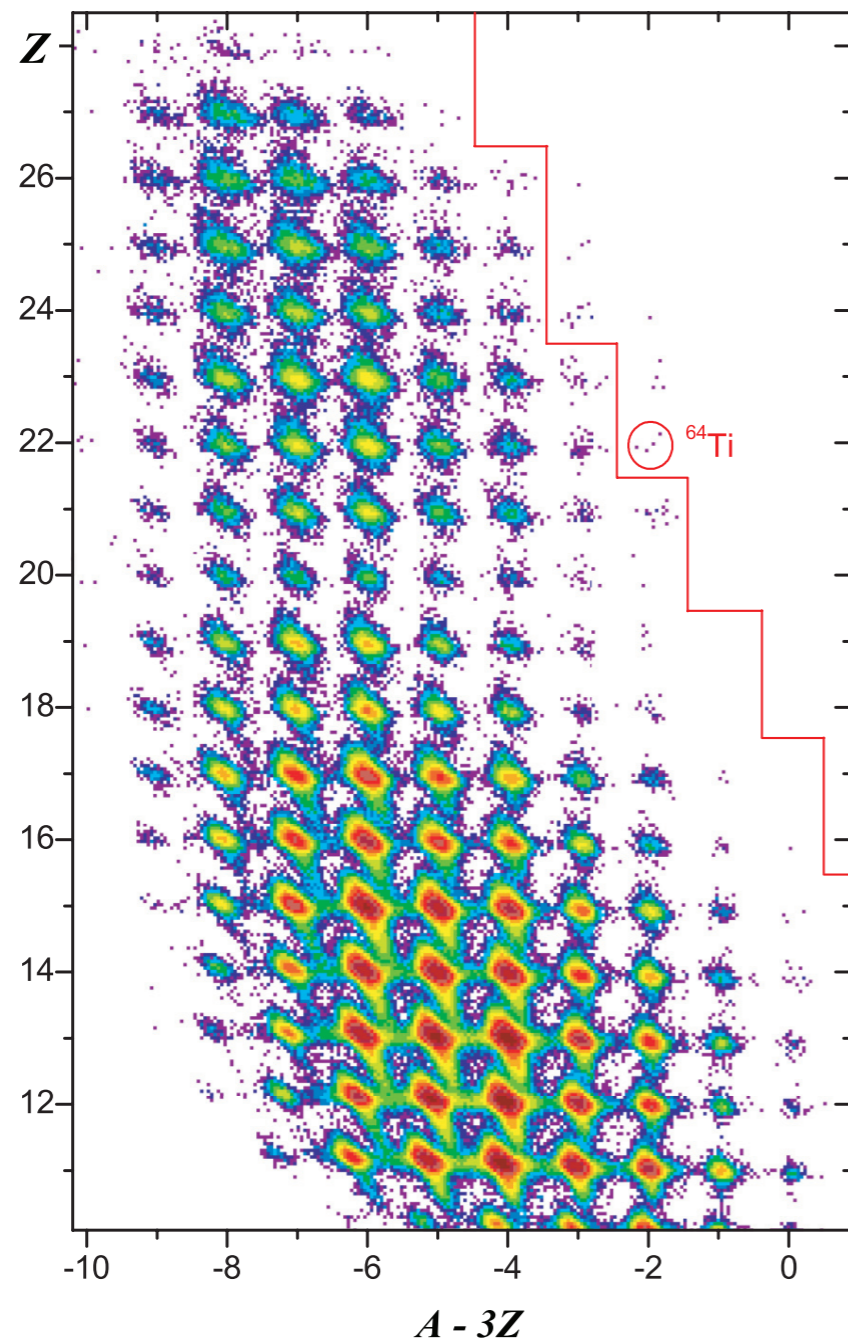
Isotope	σ (nb)	Isotope	σ (nb)	Isotope	σ (nb)	Isotope	σ (nb)
$^{157}\text{Nd}^*$	980(40)	^{168}Gd	78(5)	^{176}Er	68(5)	^{188}Lu	0.010(3)
$^{158}\text{Nd}^*$	201(11)	^{169}Gd	10.6(15)	^{177}Er	18(2)	$^{190}\text{Hf}^*$	0.027(13)
^{159}Nd	39(4)	^{170}Gd	2.6(8)	^{178}Er	5.5(9)	^{193}Ta	0.017(5)
^{160}Nd	9.5(22)	^{169}Tb	751(28)	$^{178}\text{Tm}^*$	24(3)	^{194}Ta	0.0037(19)
^{161}Nd	3.0(17)	^{170}Tb	99(6)	^{179}Tm	1.21(18)	$^{195}\text{W}^*$	0.049(1)
^{160}Pm	518(36)	^{171}Tb	14(2)	^{180}Tm	4.5(9)	^{196}W	0.018(4)
^{161}Pm	161(9)	^{172}Tb	1.0(4)	^{181}Tm	0.6(3)	^{197}W	0.0034(17)
^{162}Pm	25(3)	^{171}Dy	441(18)	$^{181}\text{Yb}^*$	2.3(3)	$^{198}\text{Re}^*$	0.028(7)
^{163}Pm	4.5(15)	^{172}Dy	121(7)	$^{182}\text{Yb}^*$	0.45(10)	^{199}Re	0.0076(27)
^{163}Sm	134(11)	^{173}Dy	18(2)	^{183}Yb	0.21(5)	^{202}Os	0.0044(20)
^{164}Sm	42(4)	^{174}Dy	1.9(6)	^{184}Yb	0.028(9)	^{203}Os	0.0025(18)
^{165}Sm	7.8(16)	^{173}Ho	341(15)	^{185}Yb	0.007(3)	^{205}Ir	0.003(2)
^{167}Eu	7.1(12)	^{174}Ho	98(6)	$^{185}\text{Lu}^*$	0.22(7)	^{206}Pt	0.033(11)
^{168}Eu	2.0(8)	^{175}Ho	22(2)	$^{186}\text{Lu}^*$	0.15(4)	^{207}Pt	0.008(3)
^{167}Gd	625(23)	^{176}Ho	2.2(6)	^{187}Lu	0.043(9)	^{208}Pt	0.0027(15)

Recent Works for New Isotope Search

NSCL (2013)

^{82}Se 139 MeV/u on Be Target Projectile Fragmentation

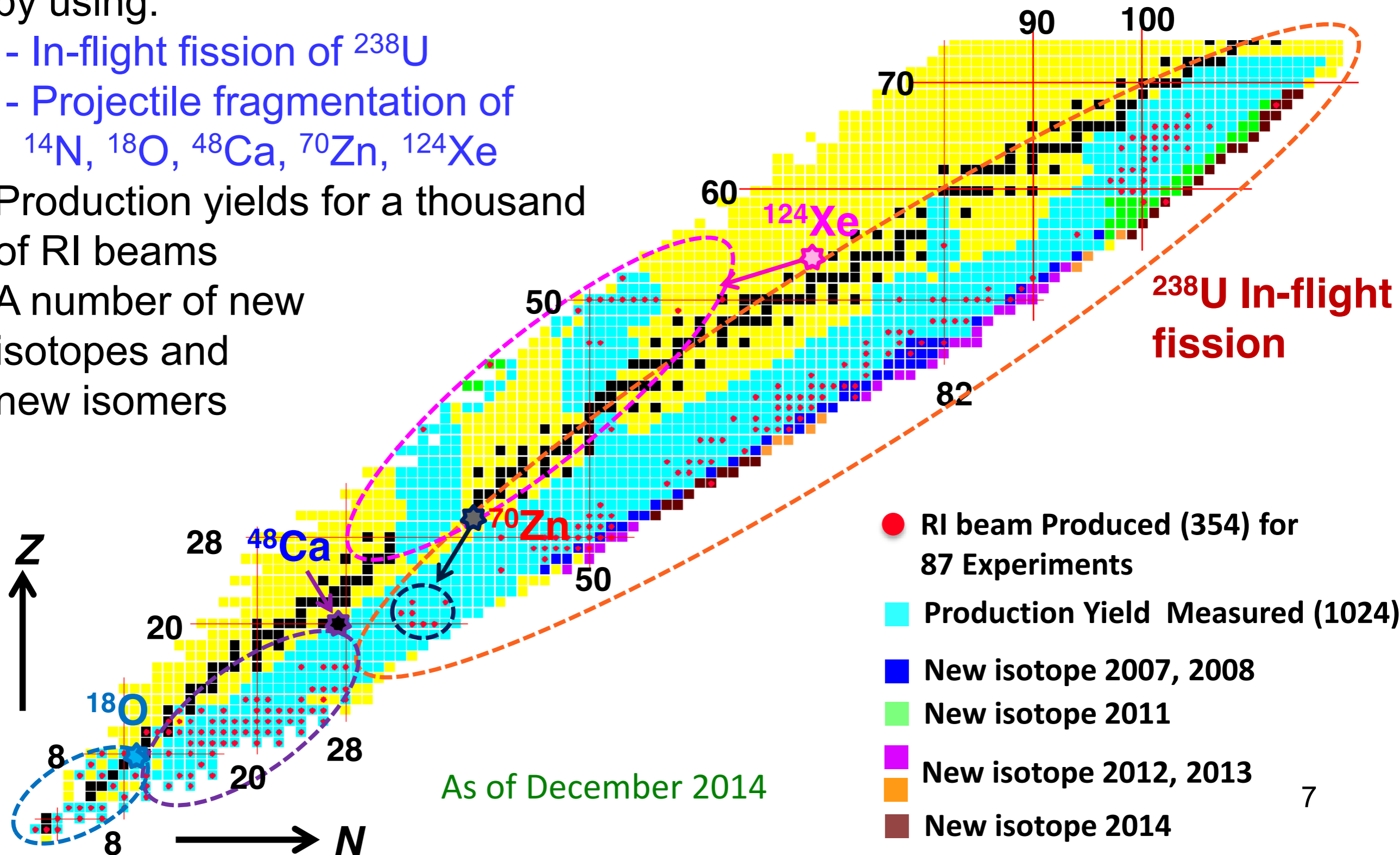
O. B. Tarasov et al., Phys. Rev. C **87**, 054612 (2013)



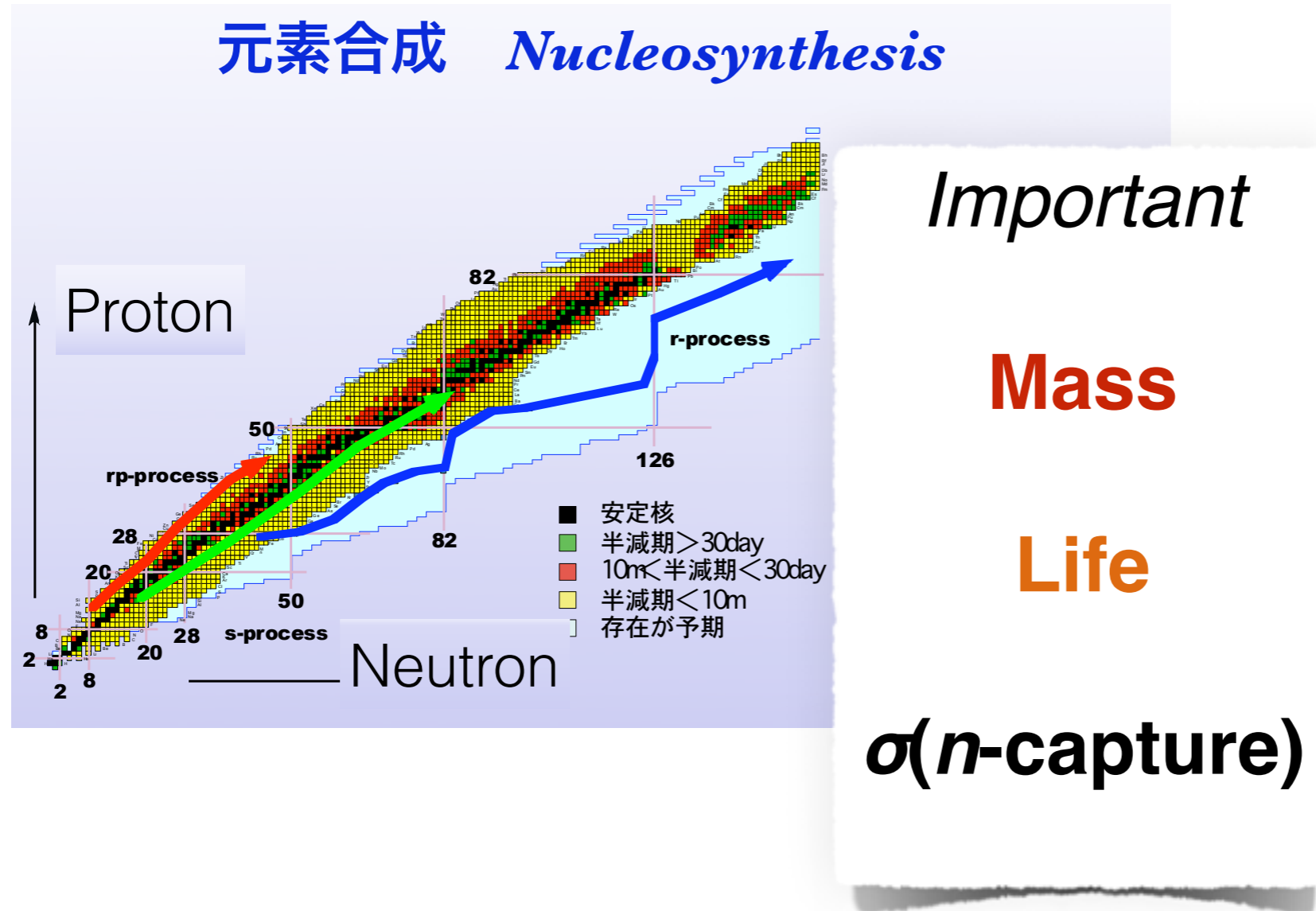
RI beams produced at BigRIPS (May 2007 – Dec. 2014)

From presentation file of Prof. T. Kubo, RIKEN

- We have produced a total of 354 RI beams and delivered to 87 experiments.
- by using:
 - In-flight fission of ^{238}U
 - Projectile fragmentation of ^{14}N , ^{18}O , ^{48}Ca , ^{70}Zn , ^{124}Xe
- Production yields for a thousand of RI beams
- A number of new isotopes and new isomers



Mass and Lifetime Measurements



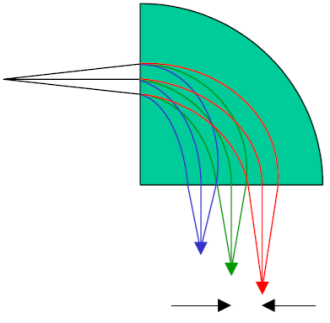
Mass Measurements of unstable nuclei
Lifetime of neutron-rich nuclei

Lifetime of bare nuclei • • • Re/Os clock

Direct MASS Measurements

Mass Spectrometer

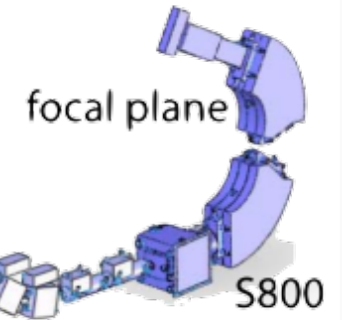
From the lecture note of Prof. A. Gade, NSCL



Mass separator
(spectrograph,
spectrometer)

Dispersion

$$D = \Delta x \cdot m / \Delta m$$



$$B\rho = \gamma m/q \cdot (dx/dt)$$

transfer line

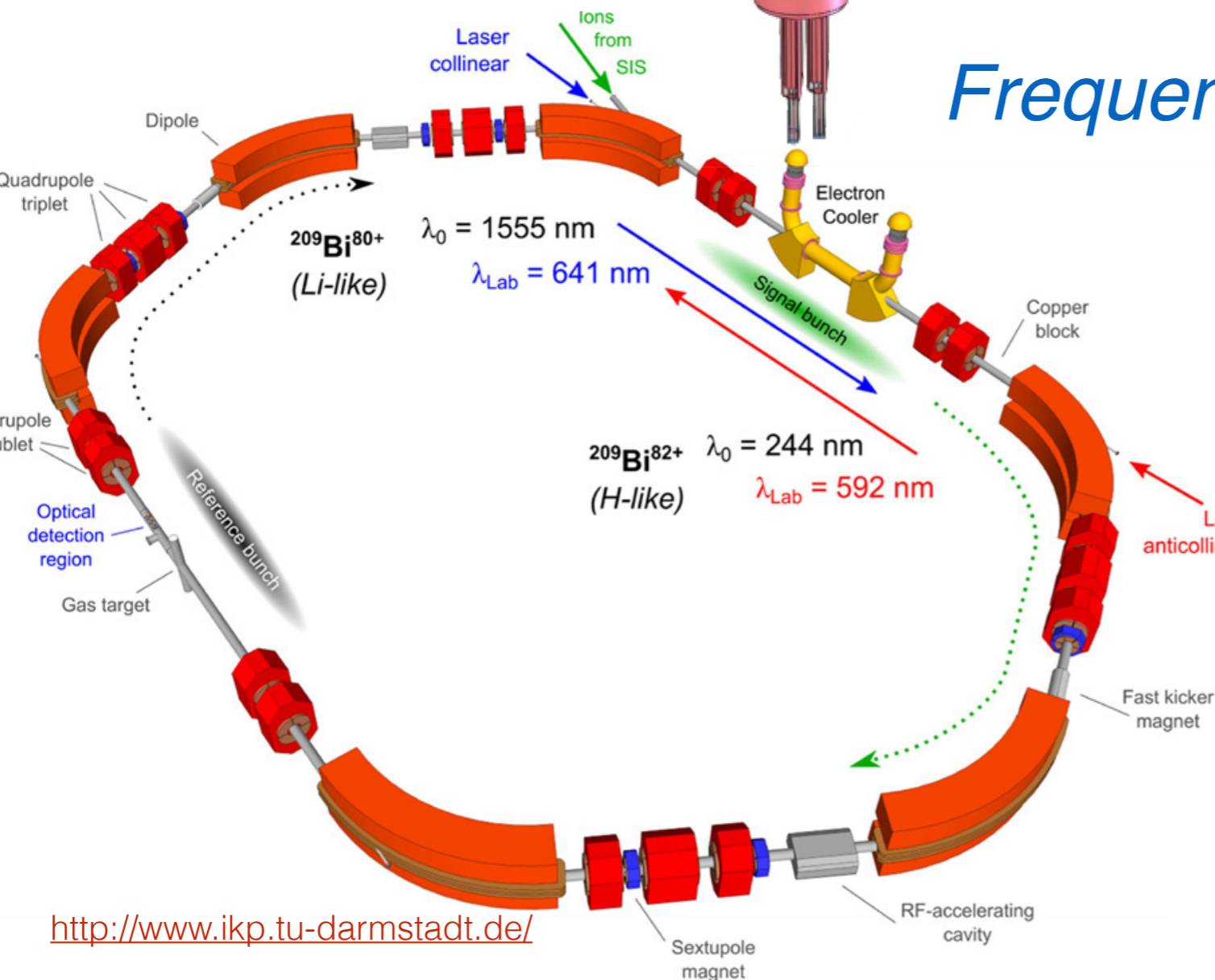
analysis line

S800

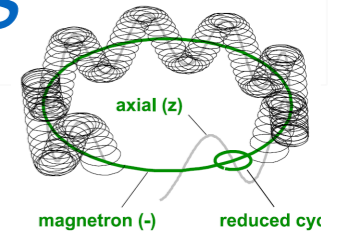
TOF Spectrometer

From the lecture note of Prof. A. Gade, NSCL

Frequency Measurements



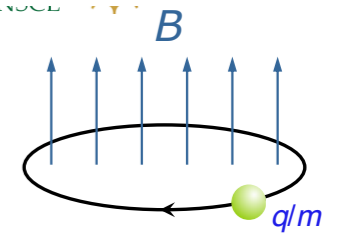
$$f_c = \frac{1}{2\pi} \cdot \frac{q}{m} \cdot B$$



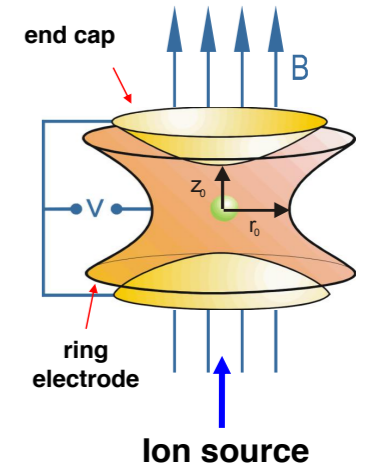
$$f_+ + f_- = f_c = \frac{1}{2\pi} \cdot \frac{q}{m} \cdot B$$

Penning Trap

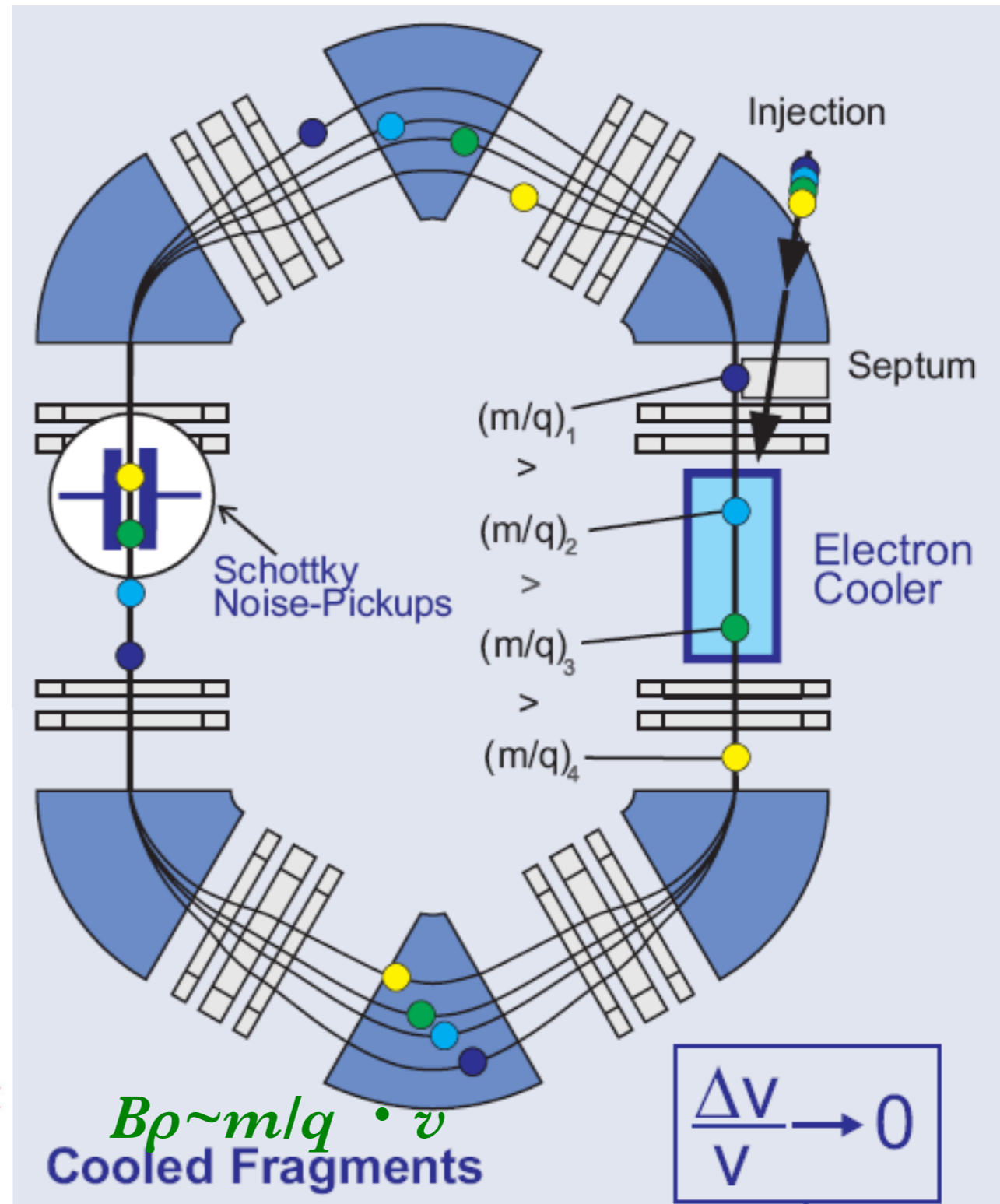
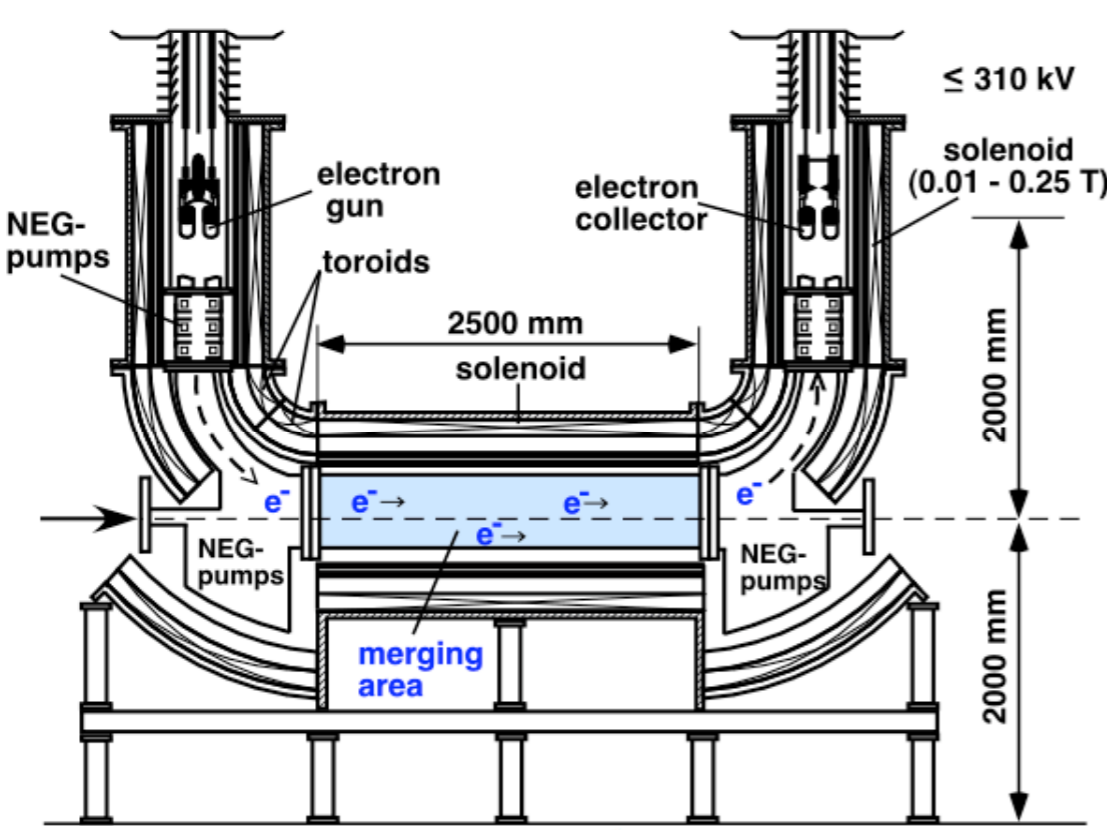
From the lecture note of Prof. A. Gade, NSCL



Storage Ring

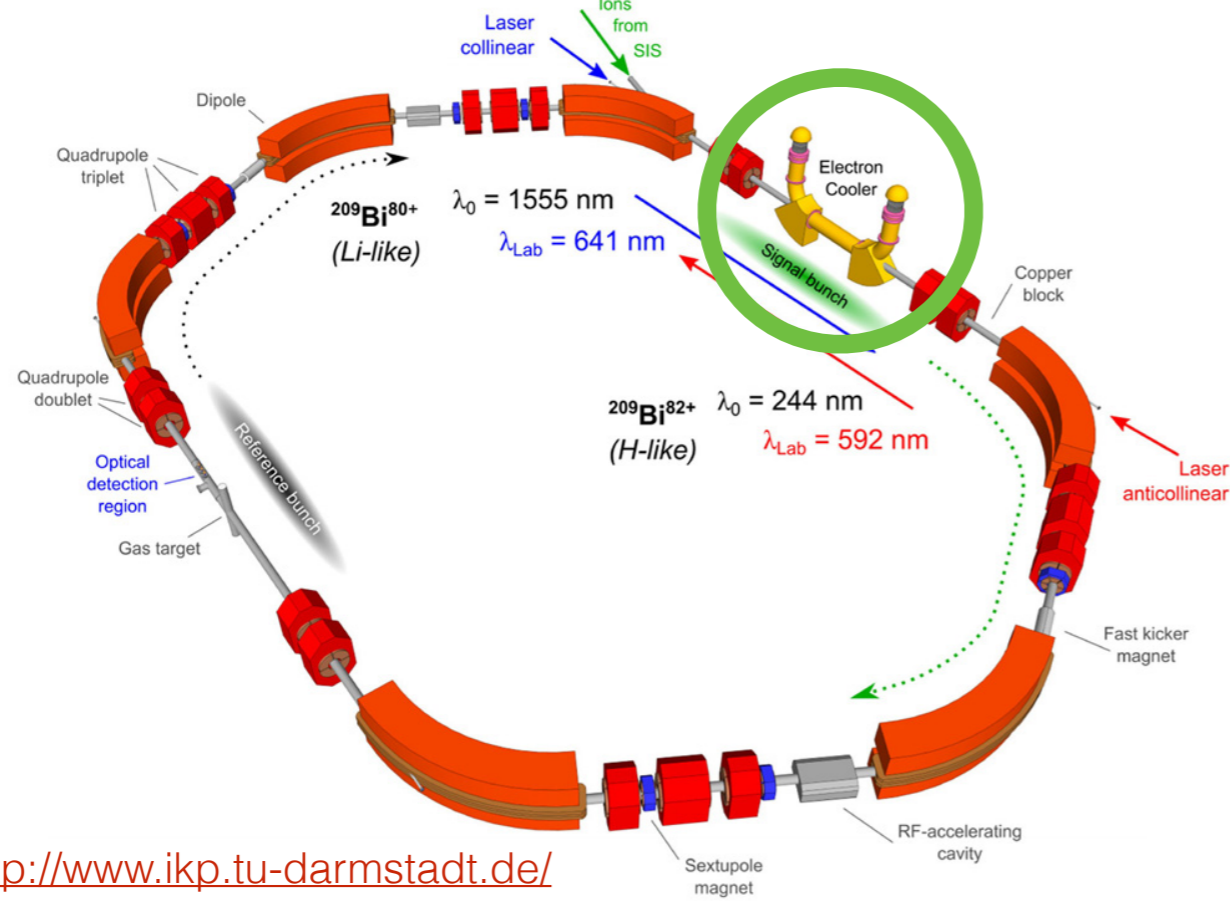


Mass Measurements with Experimental Storage Ring (ESR at GSI)



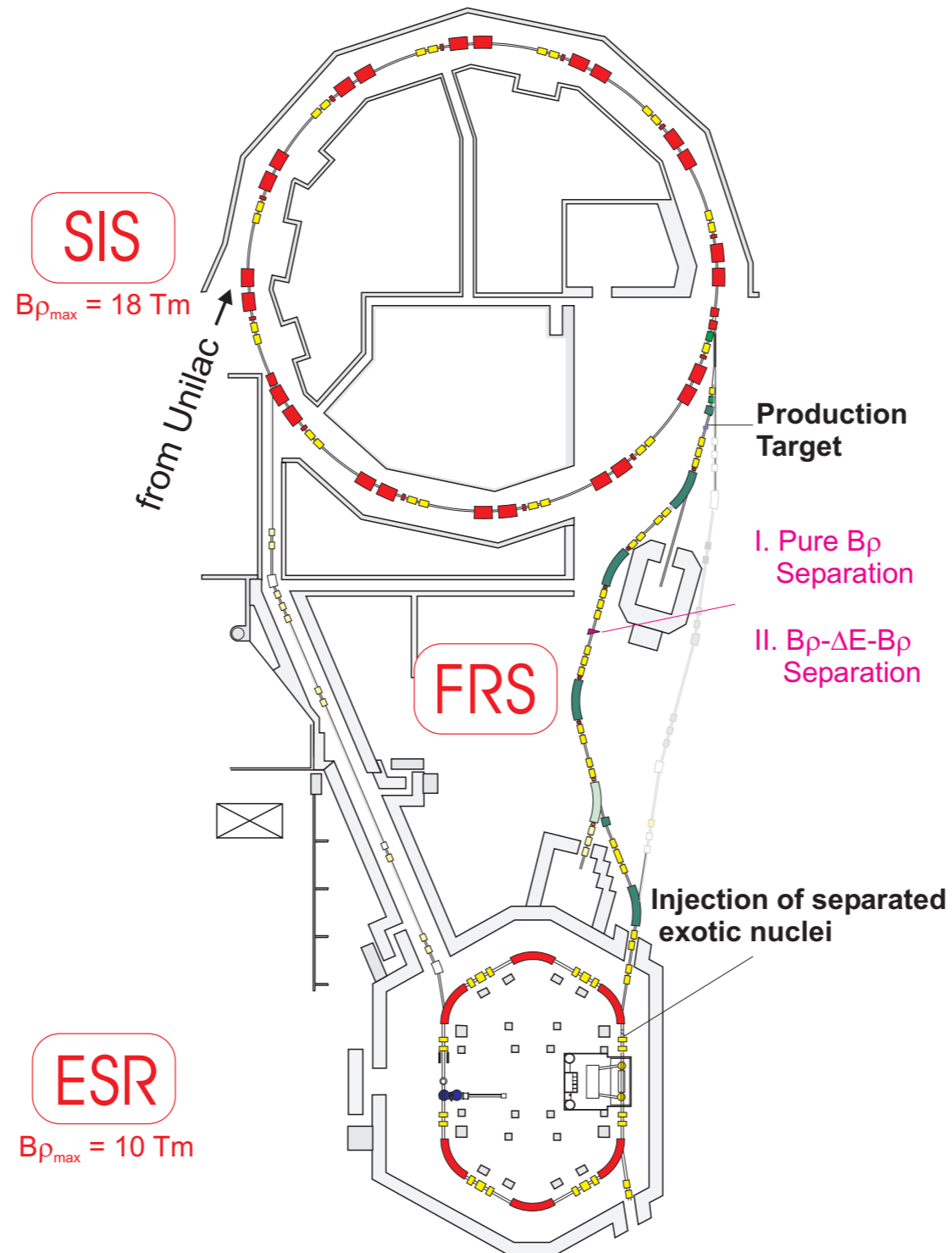
F. Bosch, Lect. Notes Phys. 651, 137 (2004)

$T_{1/2} > 1s$

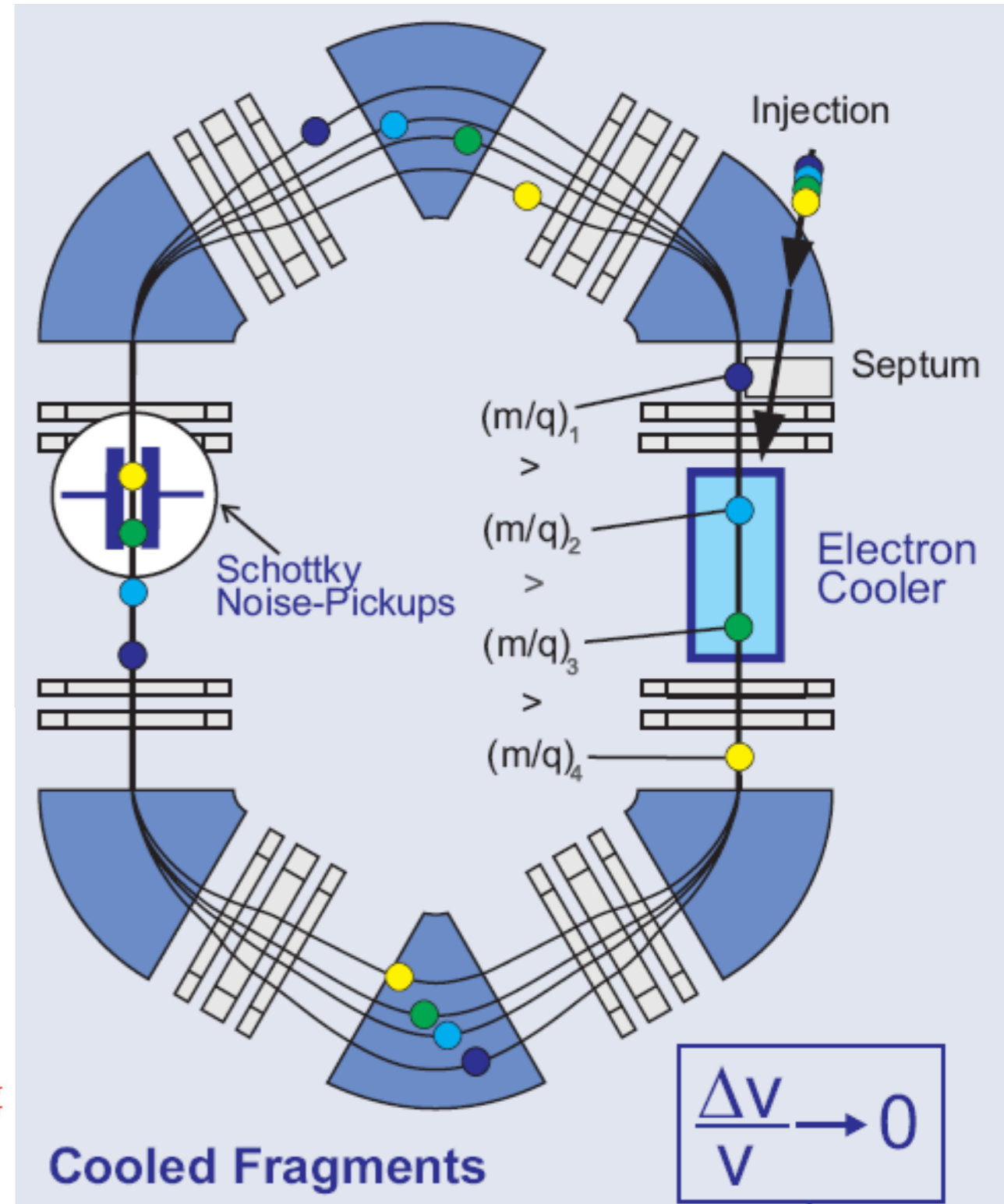
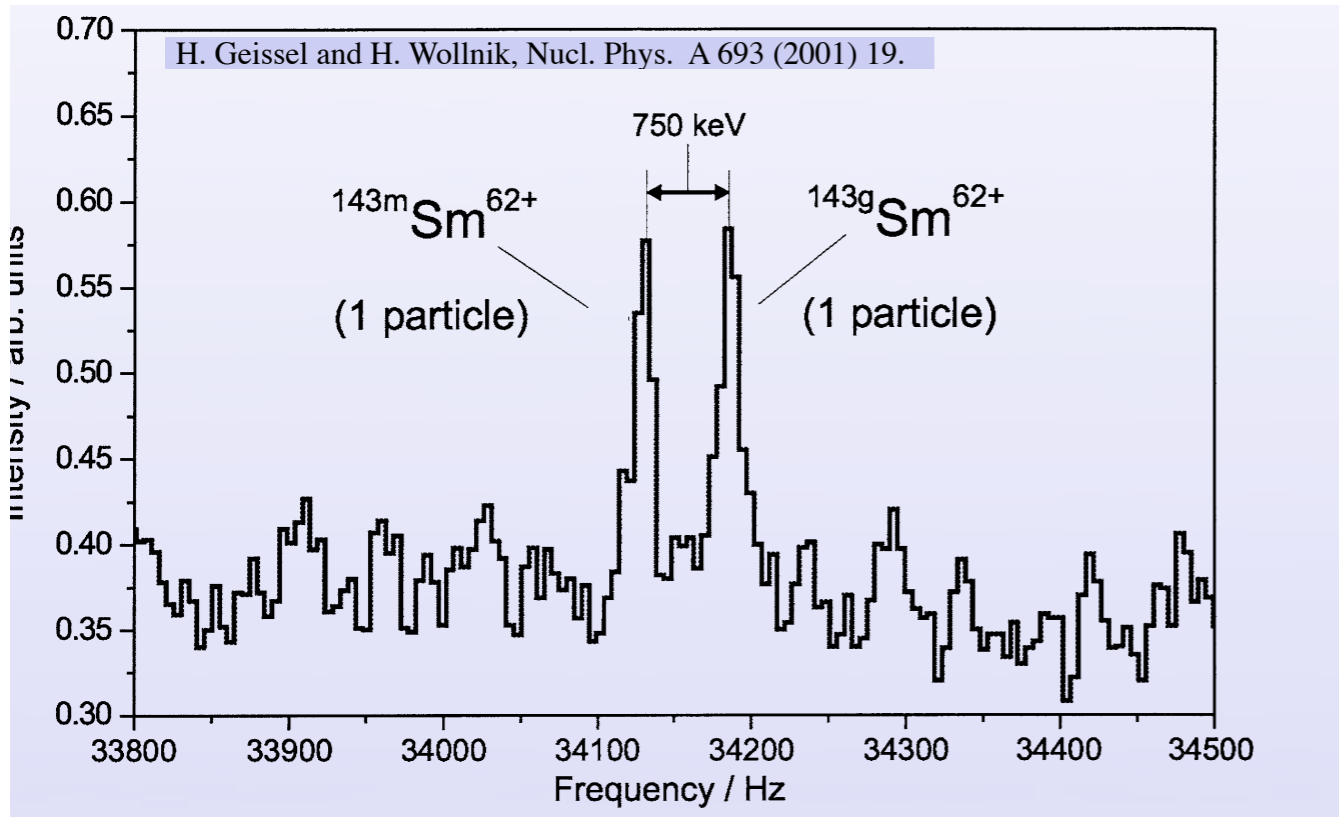


Mass Measurements with Storage Ring (ESR at GSI)

Mass and Lifetime Measurements
of Stored Exotic Nuclei

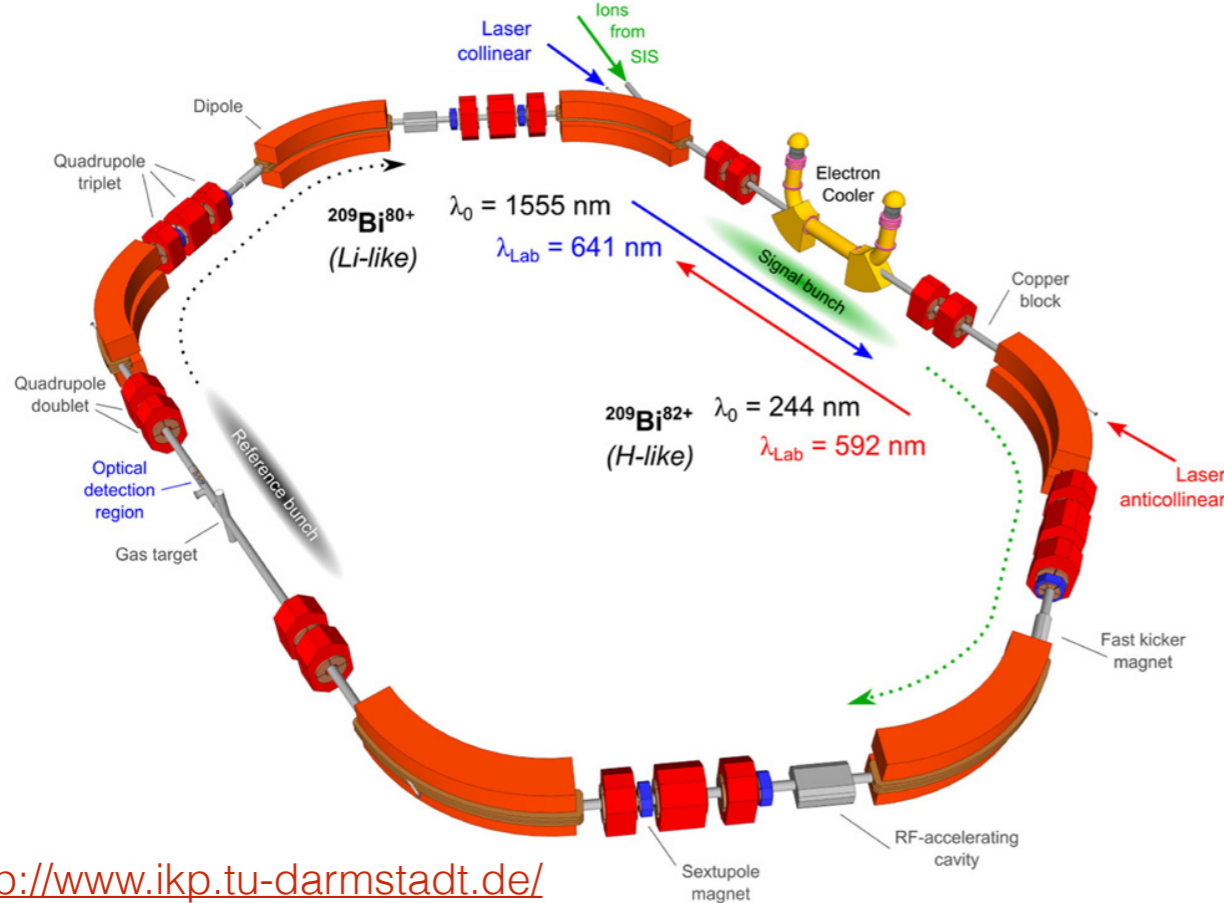


Mass Measurements with ESR



F. Bosch, Lect. Notes Phys. 651, 137 (2004)

$T_{1/2} > 1\text{ s}$



Lifetime Measurements with ESR

ESR

Injection of **charge-stripped** Mother nucleus

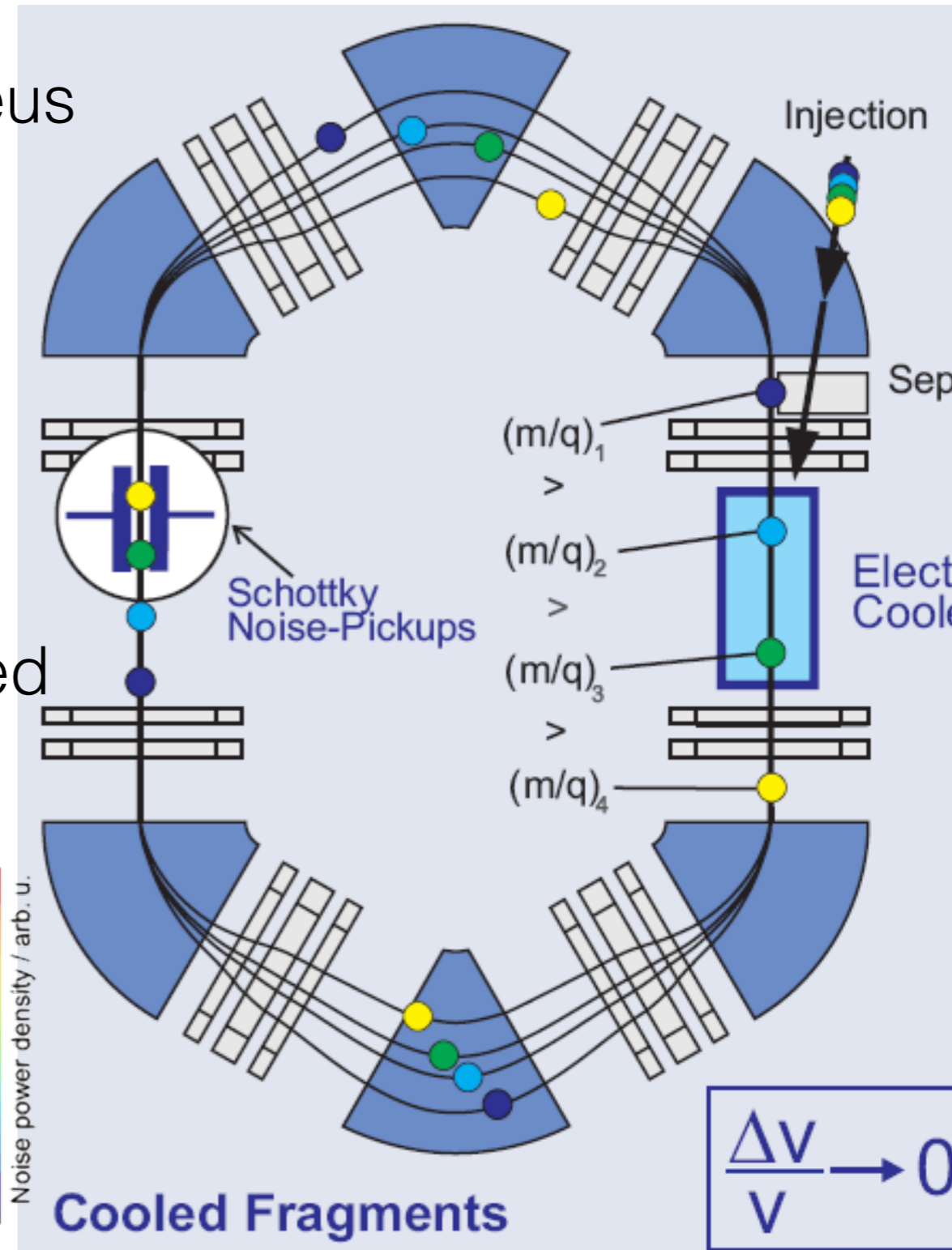
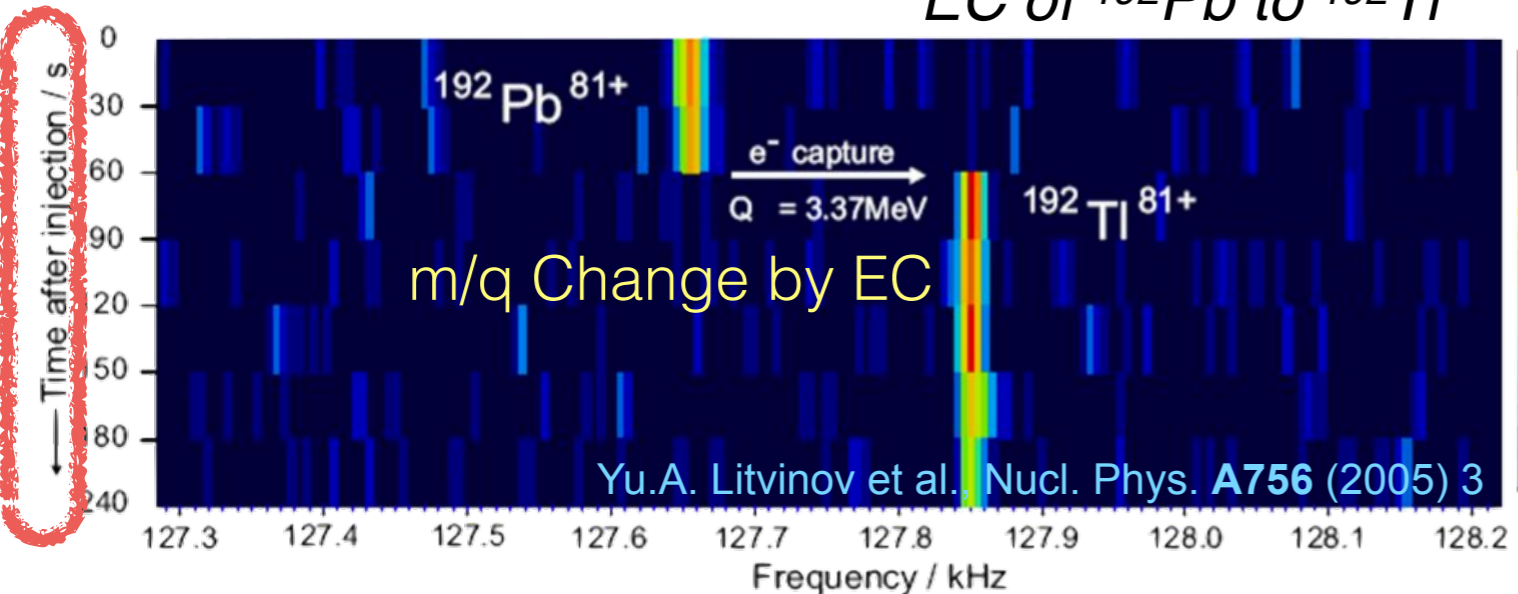
Storage it until it decays into daughter nucleus

Decay

Sudden Change of frequency
 $m_{\text{Mother}}/q \rightarrow m_{\text{Daughter}}/q$

Time from injection to decay \rightarrow recorded
Lifetime

EC of ^{192}Pb to ^{192}Tl



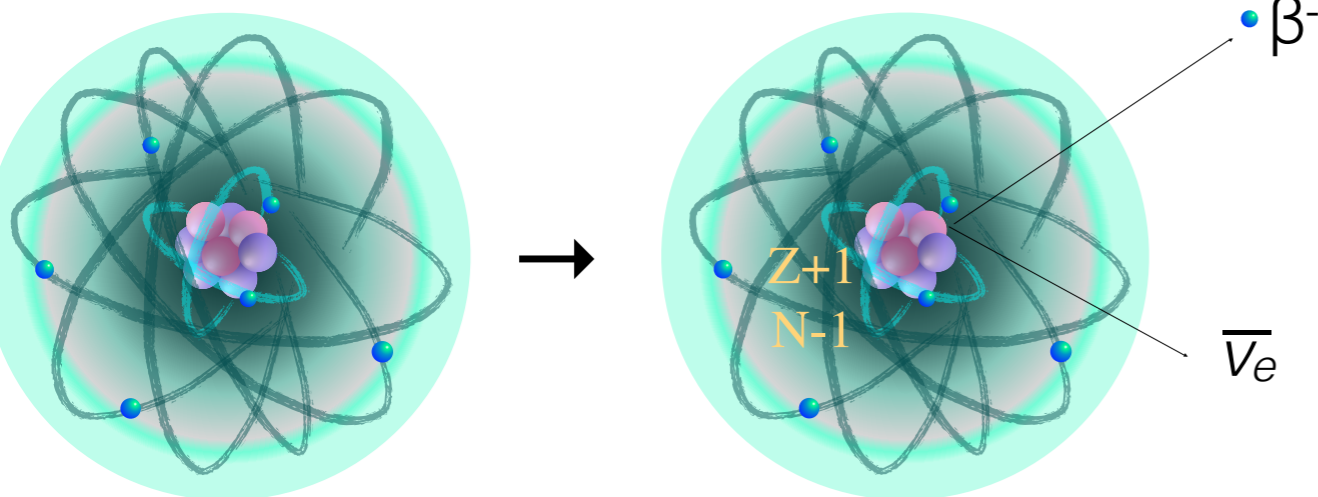
F. Bosch, Lect. Notes Phys. 651, 137 (2004)

Interesting Point :

Lifetime obtained using ESR = Lifetime of “Bare Nucleus”

Lifetime Measurements with ESR

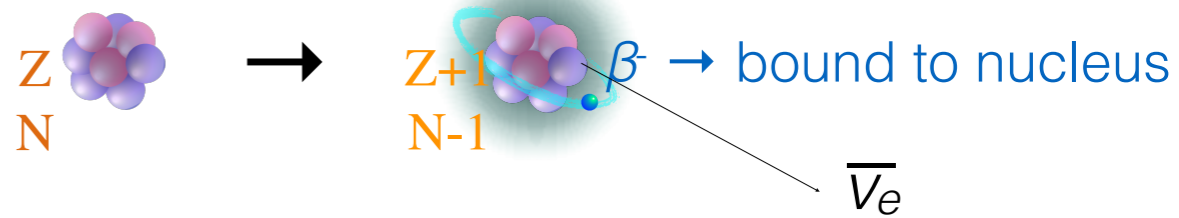
$$n \rightarrow p + \beta^- + \bar{\nu}_e$$



β^- decay : Nucleus in Atom

$$n \rightarrow p + \beta^- (\text{bound}) + \bar{\nu}_e$$

Bound-state β Decay



Bare Nucleus

“Eon Clock” for the age of Universe

$^{187}\text{Re}/^{187}\text{Os}$ (β -decay; $T_{1/2} = 42 \times 10^9$ y)

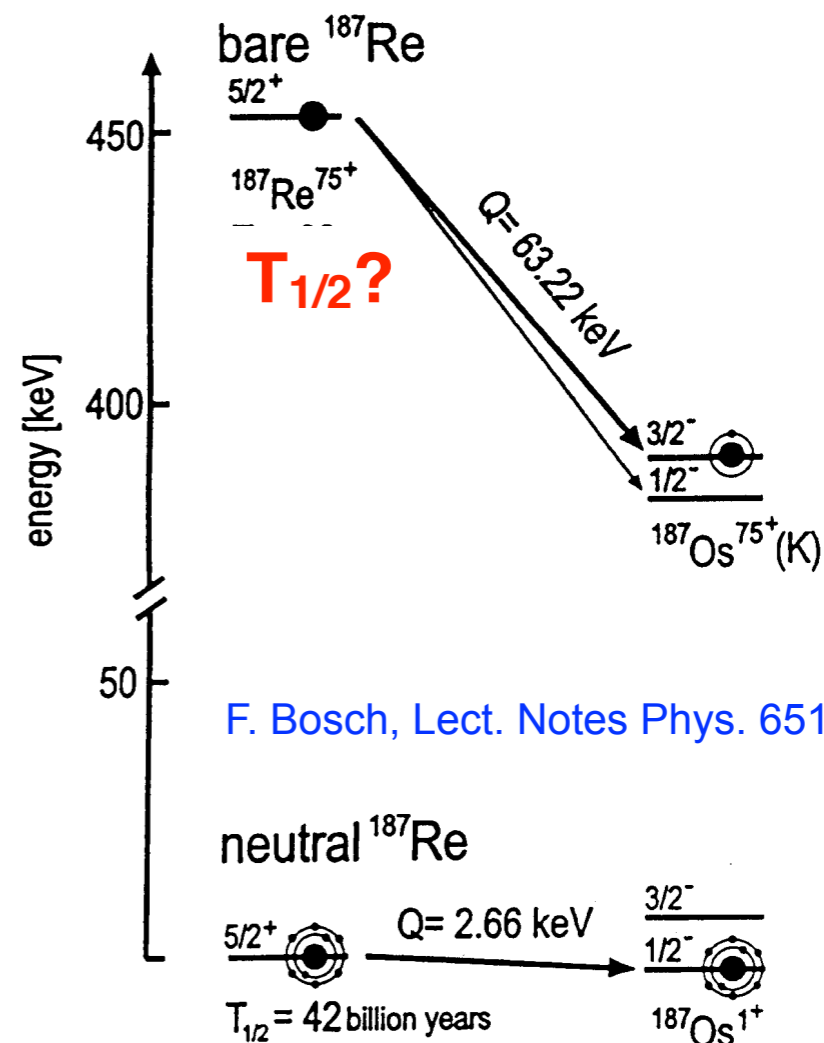
In hot stellar plasma, ^{187}Re could be bare

Bound-state β Decay



Different Q value from electron binding energy and different $T_{1/2}$

How long is the $T_{1/2}$ of bare ^{187}Re ?

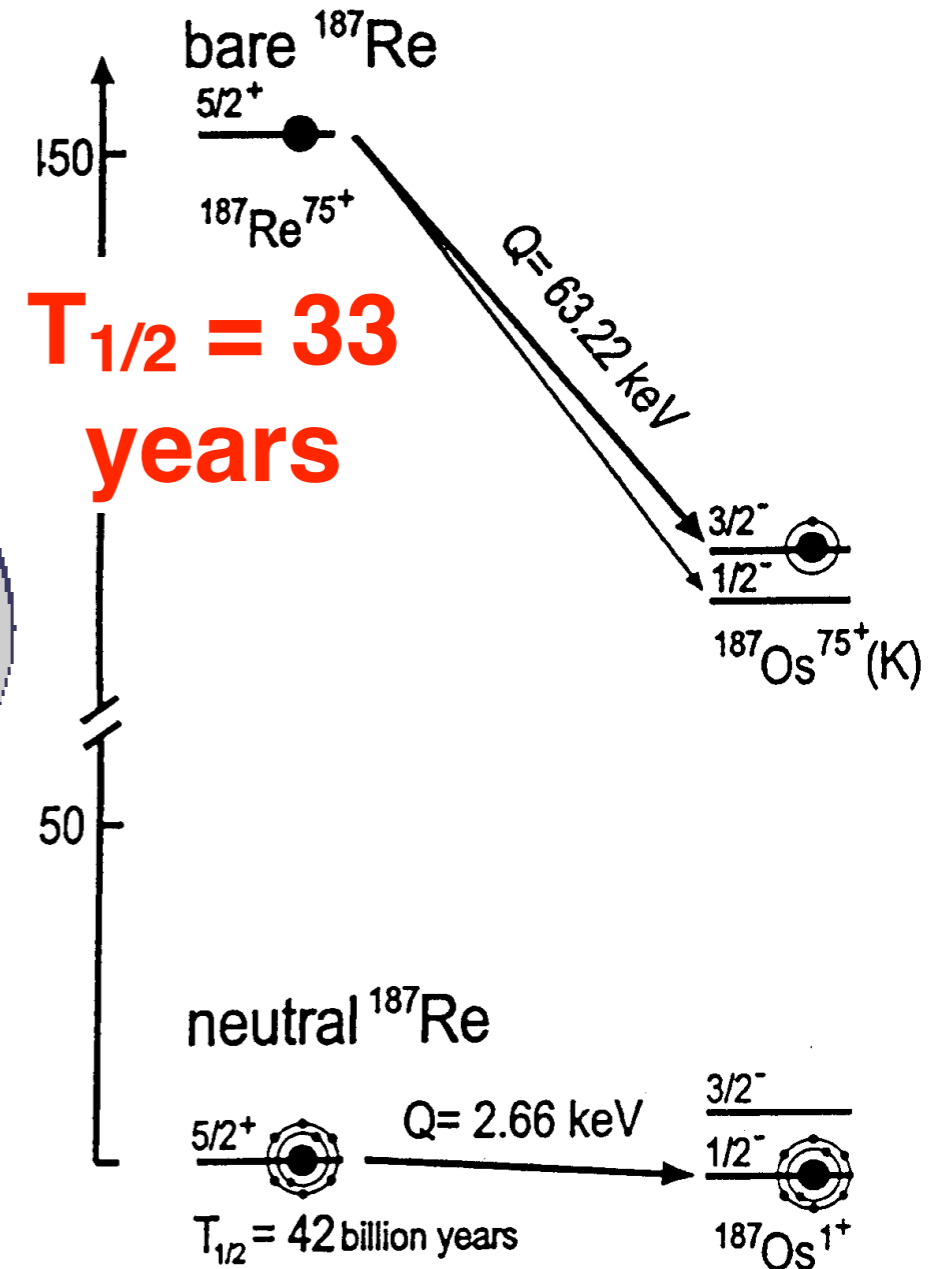
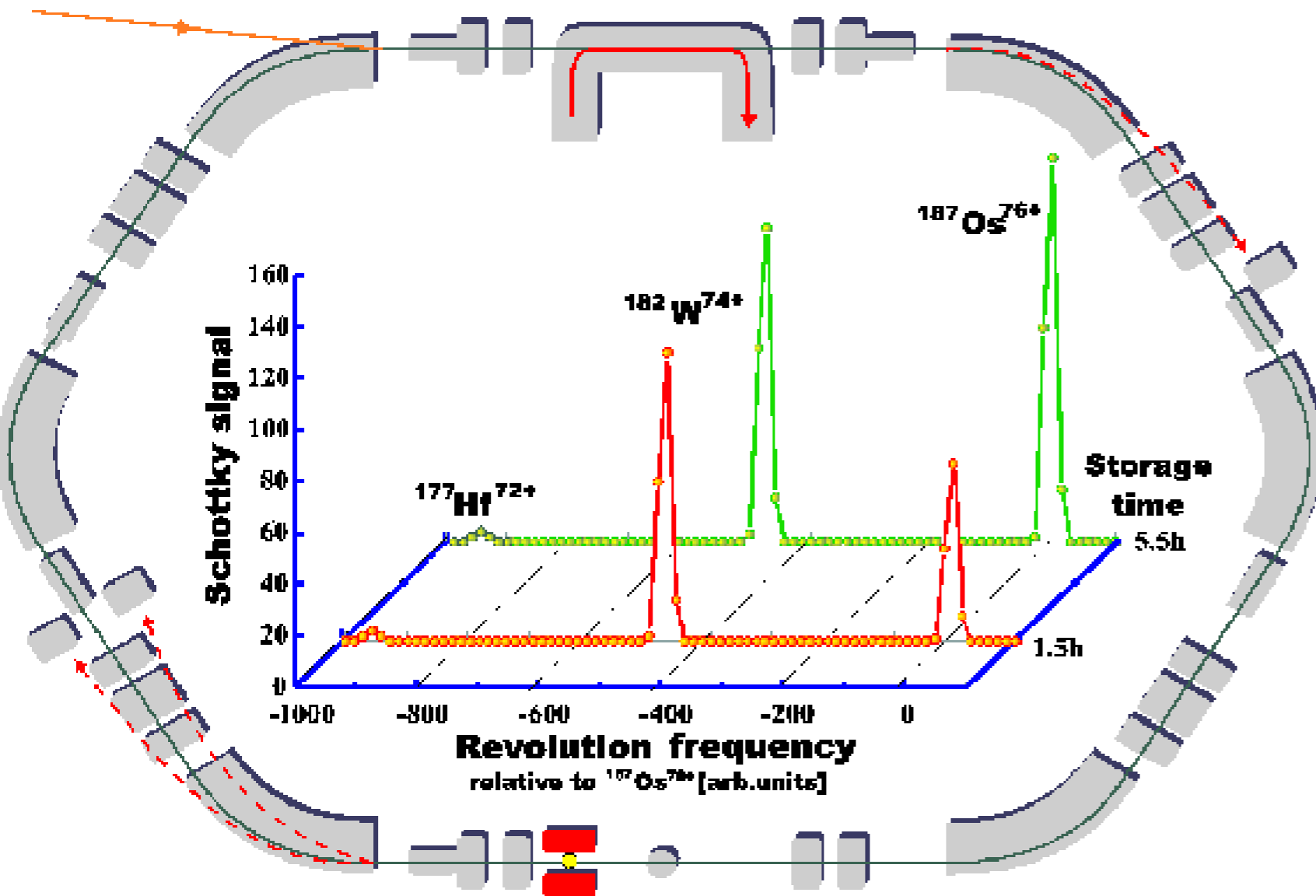


F. Bosch, Lect. Notes Phys. 651, 137 (2004)

Lifetime Measurements of ^{187}Re

F. Bosch, Lect. Notes Phys. 651, 137 (2004)

Beam : 400 MeV/u full-strip ^{187}Re Primary Beam from SIS



10^8 stored primary bare ^{187}Re ions, a few hundred ^{187}Os ions per hour were generated. From those numbers the impressively short half-life of 33 years for bare ^{187}Re has been determined.

Lifetime Measurements at RIBF

EURICA Project

Lifetime Measurements at RIBF

EURICA Project

PID tag : BigRIPS and ZDS

Beta-ray detection

WAS3ABi

- DSSSD

→ Tag the daughter which emits beta, from position information

- Measurements of beta-emitted time

→ **Lifetime**

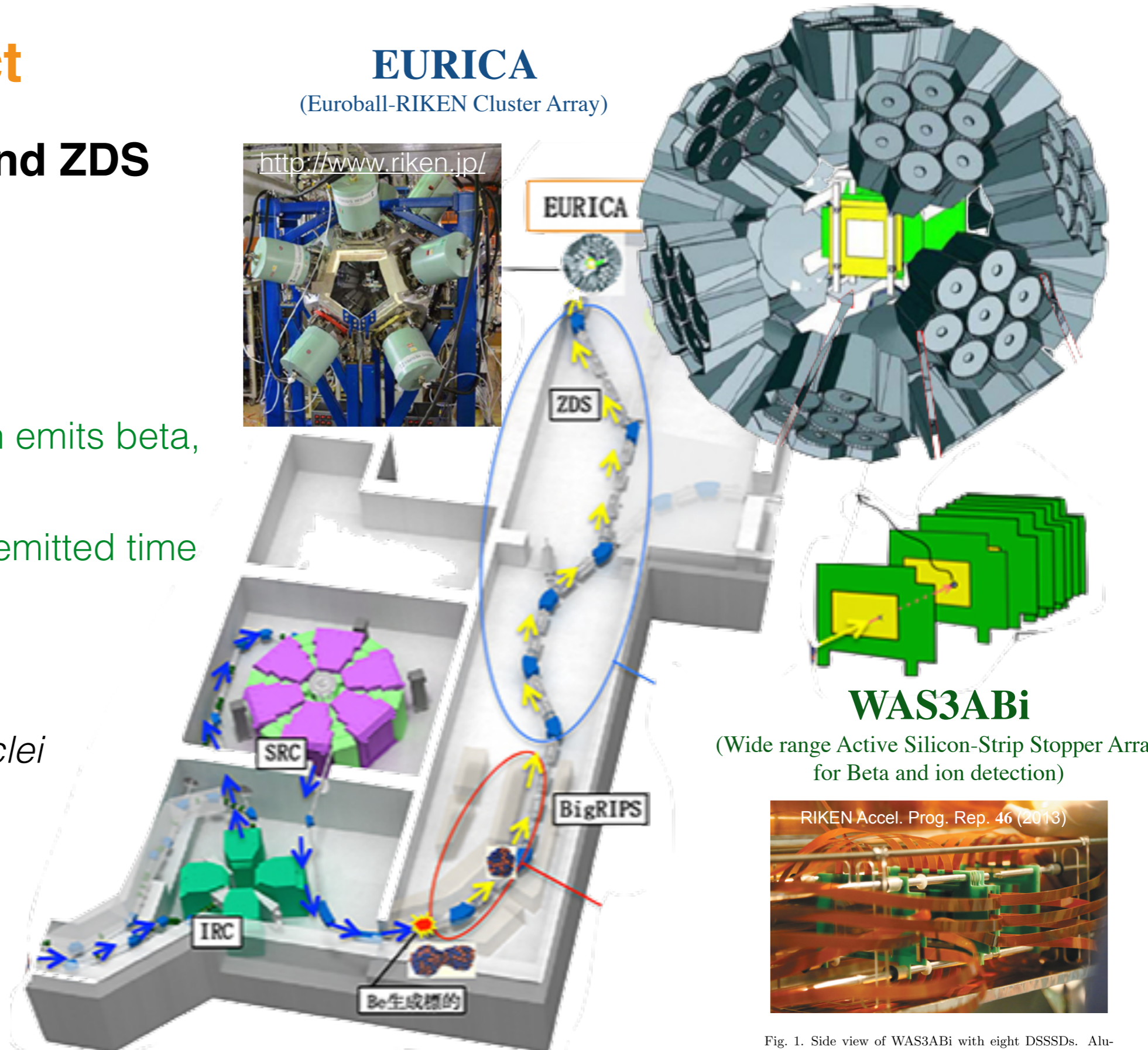
Gamma-ray from Isomer and Daughter Nuclei

EURICA

- Isomer tag
- β delayed gamma

EURICA

(Euroball-RIKEN Cluster Array)



WAS3ABi
(Wide range Active Silicon-Strip Stopper Array for Beta and ion detection)

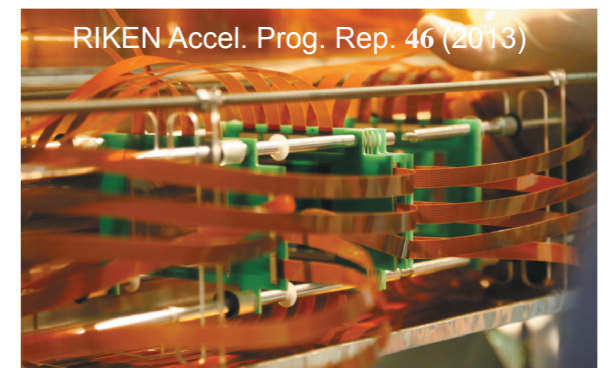


Fig. 1. Side view of WAS3ABi with eight DSSSDs. Aluminum rods were disconnected at the central position for the installation of the DSSSDs⁴.

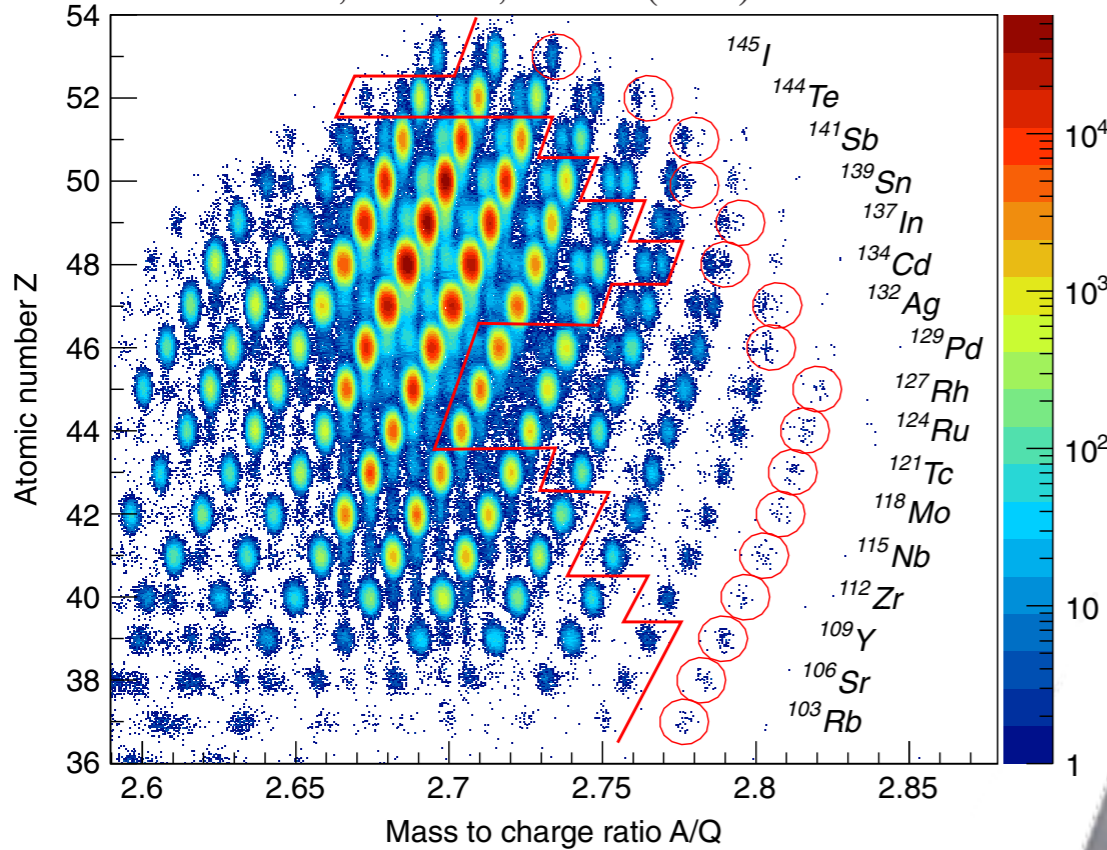
<http://www.riken.jp/>

Lifetime Measurements at RIBF

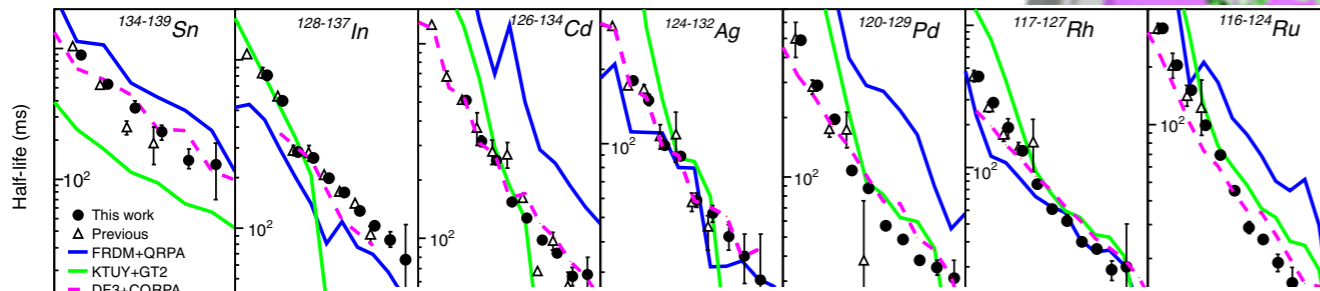
EURICA Project

Beams Implanted on WAS3ABi

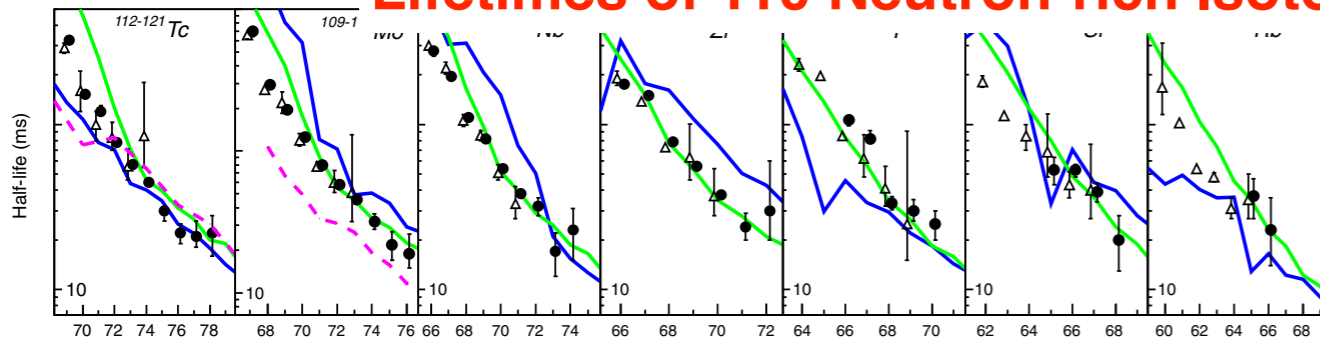
G. Lorusso et al., PRL 114, 192501 (2015)



G. Lorusso et al., PRL 114, 192501 (2015)

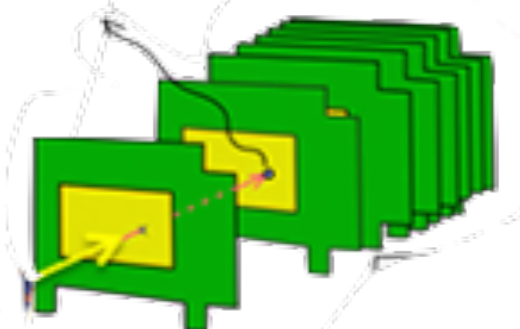
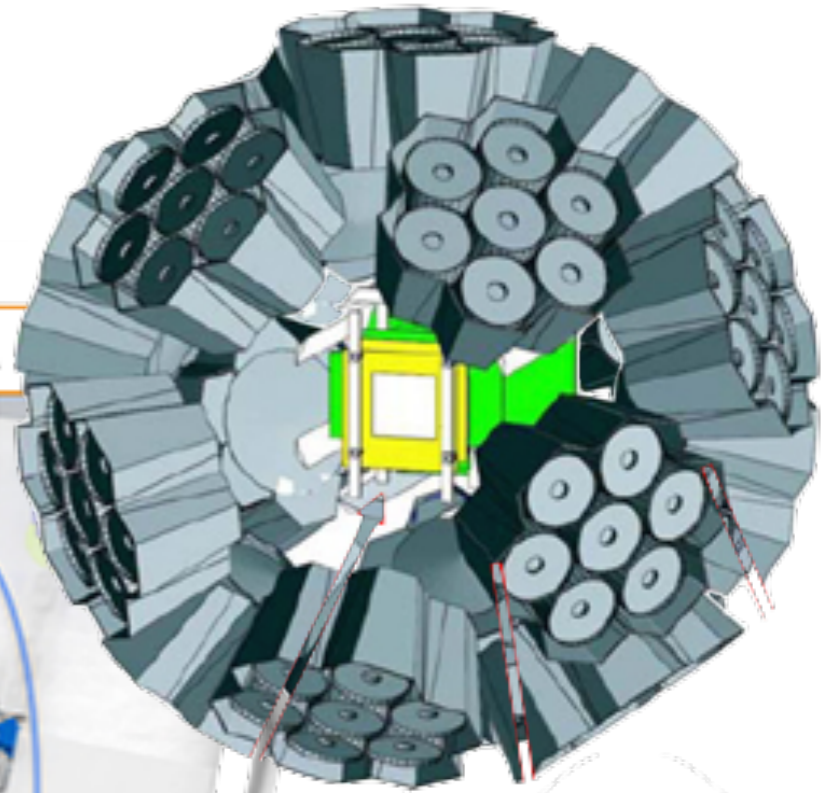


Lifetimes of 110 Neutron-rich Isotopes have been determined!



EURICA

(Euroball-RIKEN Cluster Array)



WAS3ABi

(Wide range Active Silicon-Strip Stopper Array for Beta and ion detection)

RIKEN Accel. Prog. Rep. 46 (2013)

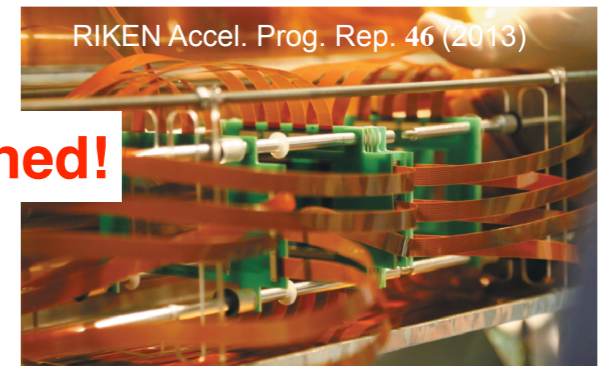


Fig. 1. Side view of WAS3ABi with eight DSSSDs. Aluminum rods were disconnected at the central position for the installation of the DSSSDs⁴⁾.

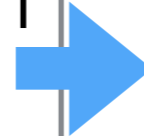
riken.jp/

Nuclear Radii Measurements

Nuclear Radii Measurements

Stable Nuclei :

Electron Scattering Experiment
X-ray Measurements Muonic Atom

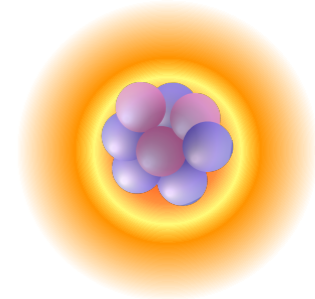


Unstable Nuclei :

Isotope shift Measurements

Charge Radii

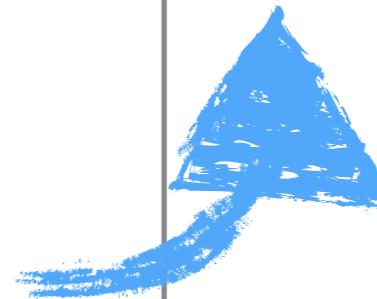
Sensitive to the
Coulomb Potential of
Protons



Stable Nuclei or Unstable Nuclei

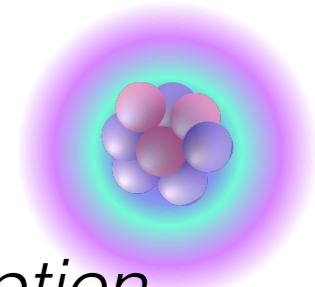
Nuclear Reaction

Elastic Scattering Cross Section
Total Reaction Cross Section



Matter Radii

Sensitive to the
Nuclear Potential of
Nucleons



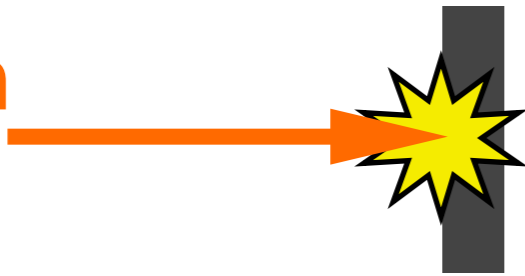
With some model assumption

Reaction Cross Section and Nuclear Size

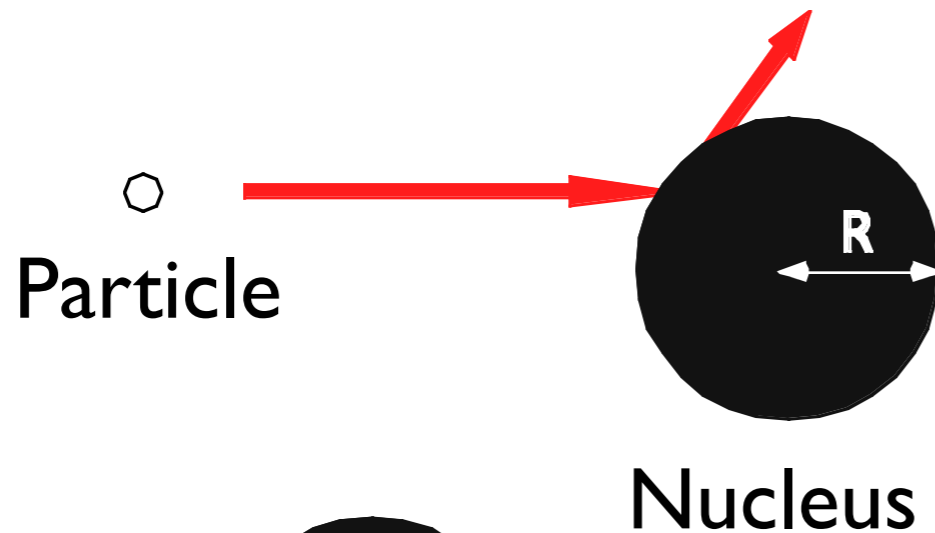
Reaction Cross Section · · · Nuclear Reaction Probability

$$\sigma_R = \sigma_{\text{tot}} - \sigma_{\text{el}}$$

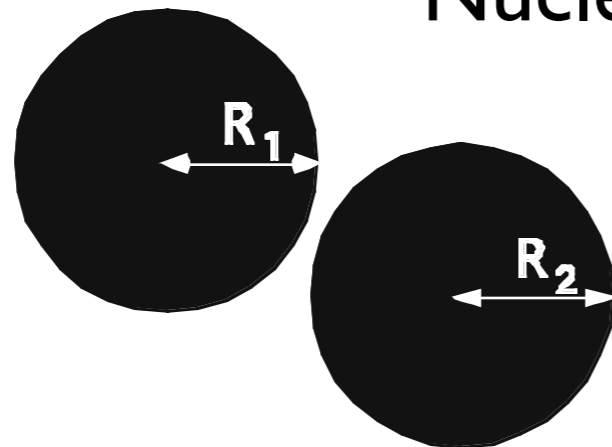
RI Beam



Experiment · · · Bombarding Target with RI Beam



$$\sigma_R \propto \pi R^2$$



$$\sigma_R \propto \pi (R_1^2 + R_2^2)$$

σ_R · · · Size of Nuclei

Reaction Cross Section *Interaction Cross Section*

$$\sigma_R = \sigma_{\text{tot}} - \sigma_{\text{el}}$$

$$\sigma_I = \sigma_R - \sigma_{\text{inel}}$$

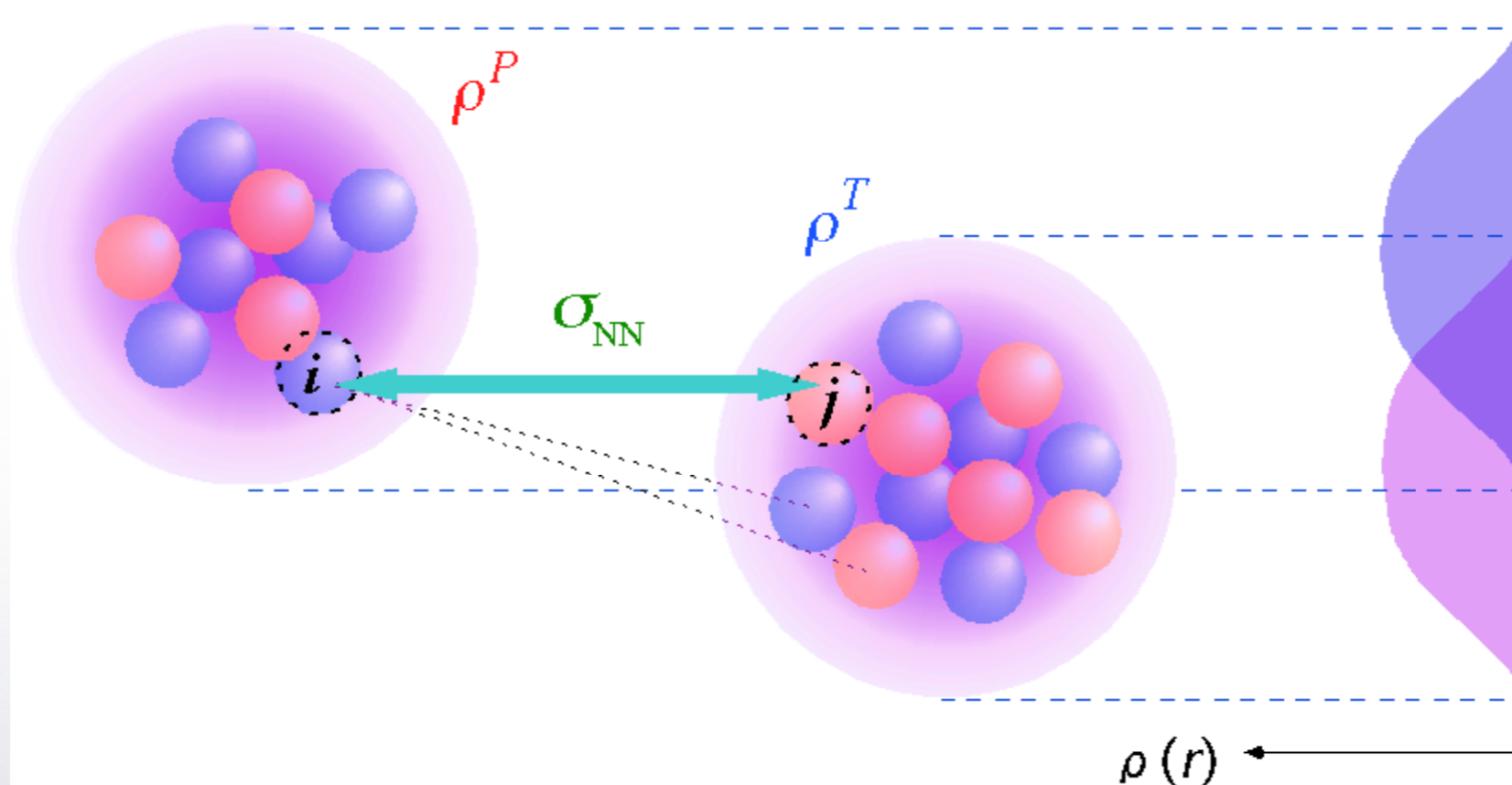
$\sigma_I \approx \sigma_R$ at high energies

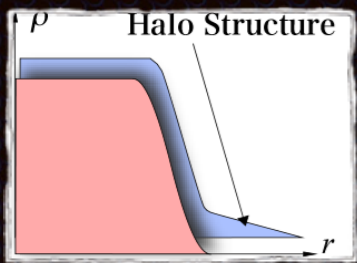
Glauber Model

$$\sigma_R = \int db \left[1 - \exp \left(- \int d^2r \sum_{i,j} \sigma_{NN}(E) \rho_z^{P_i}(\mathbf{r}) \rho_z^{T_j}(\mathbf{r} - \mathbf{b}) \right) \right]$$

- σ_{NN} Nucleon-Nucleon Total Cross Section
- ρ^P Projectile Nucleon Density Distribution
- ρ^T Target Nucleon Density Distribution

σ_I or σ_R \longleftrightarrow Nuclear Size
Glauber Calculation

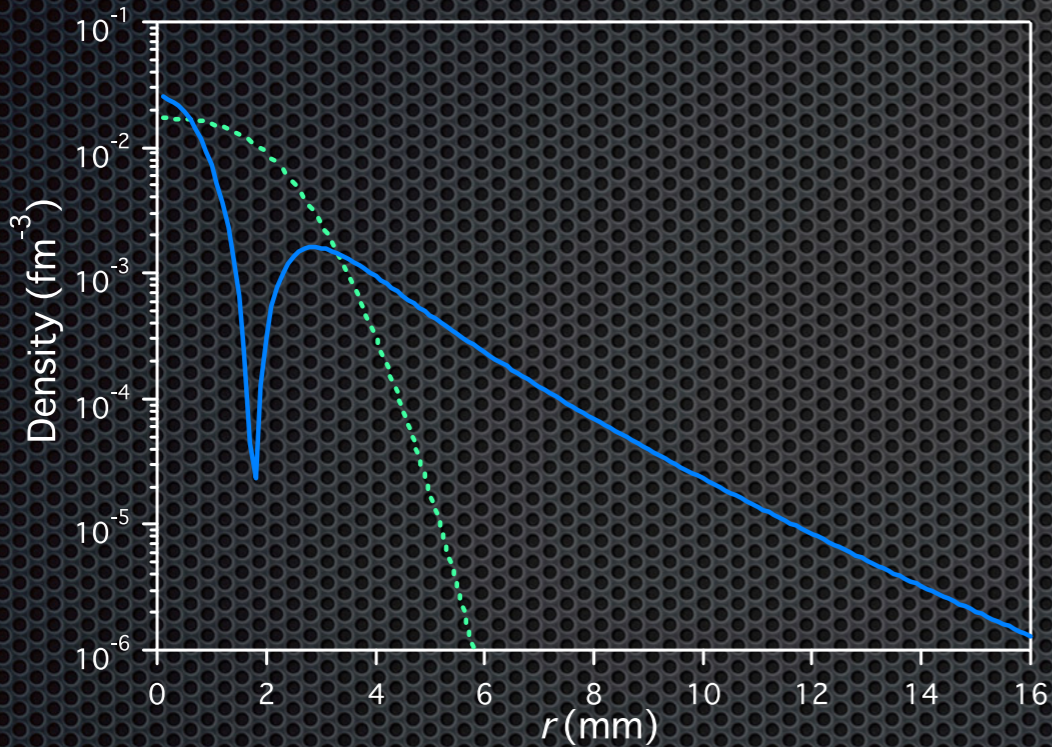
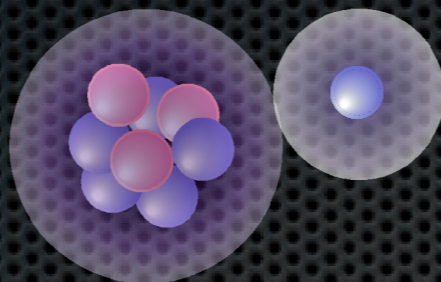




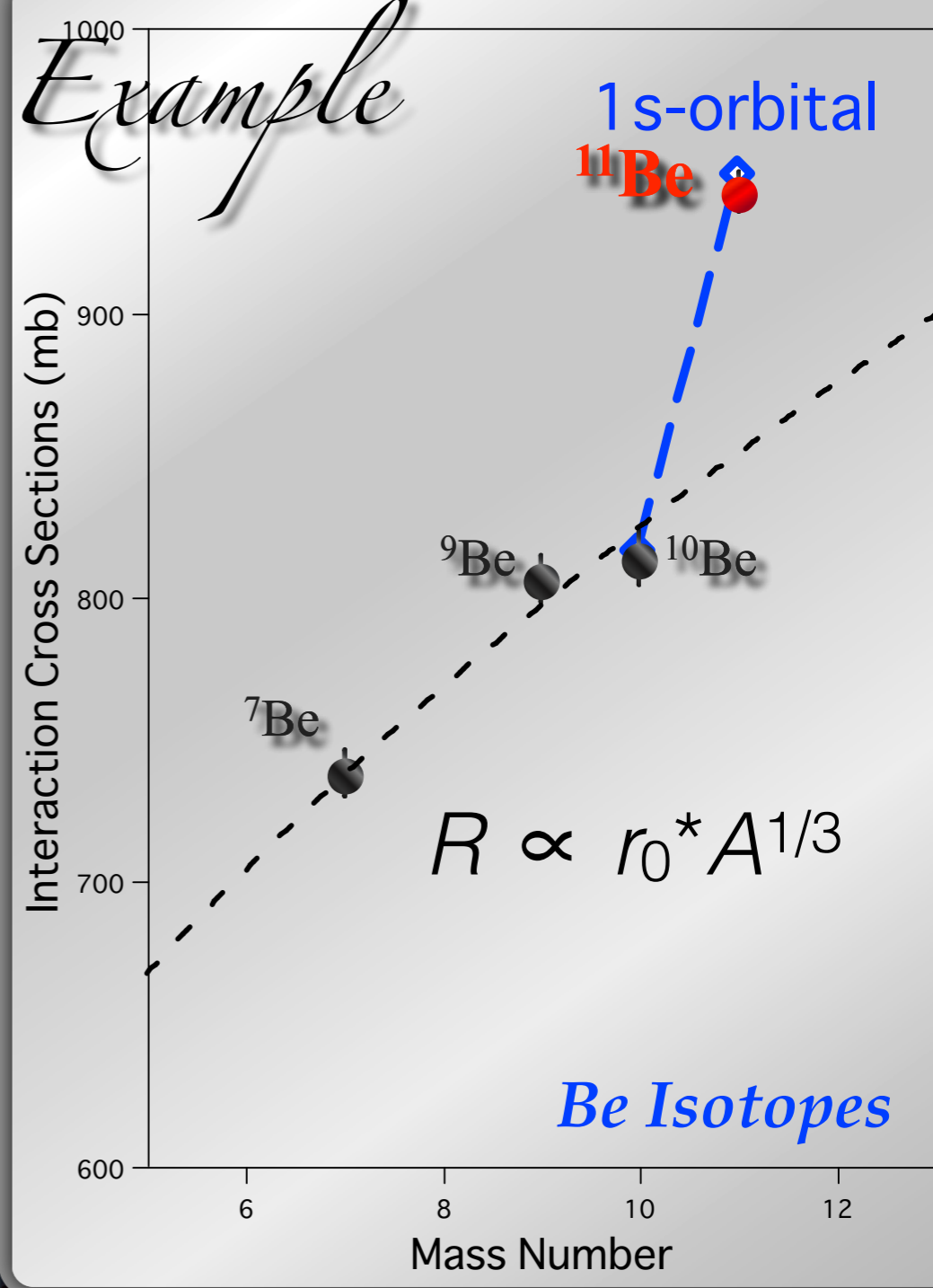
Nuclear Size and Halo Structure

Neutron-Rich Nucleus ^{11}Be

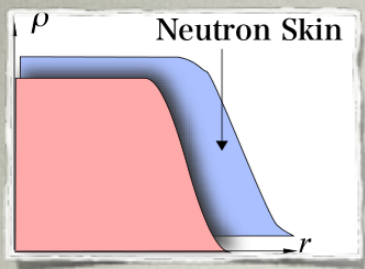
$$Z = 4, N = 7$$



Example

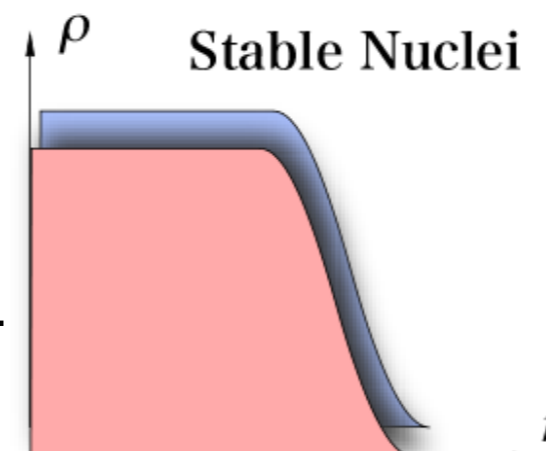
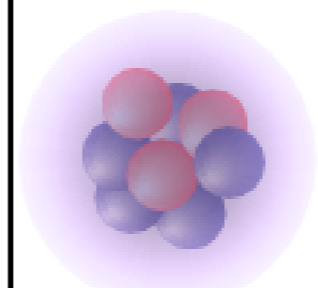
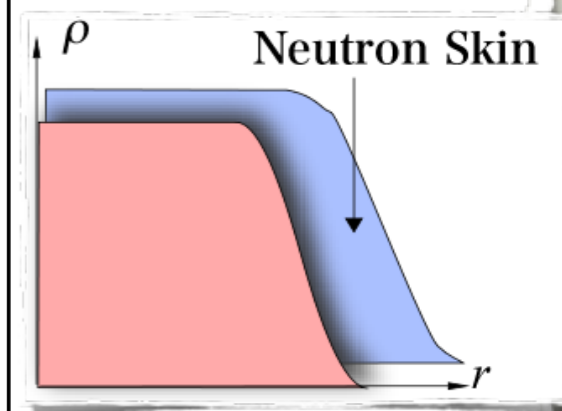
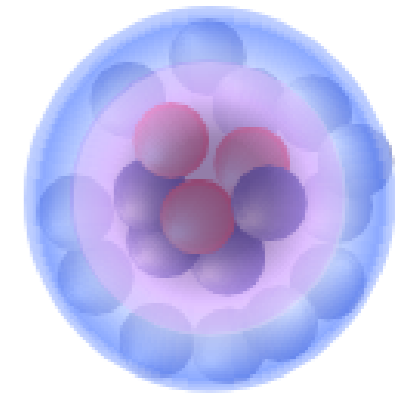
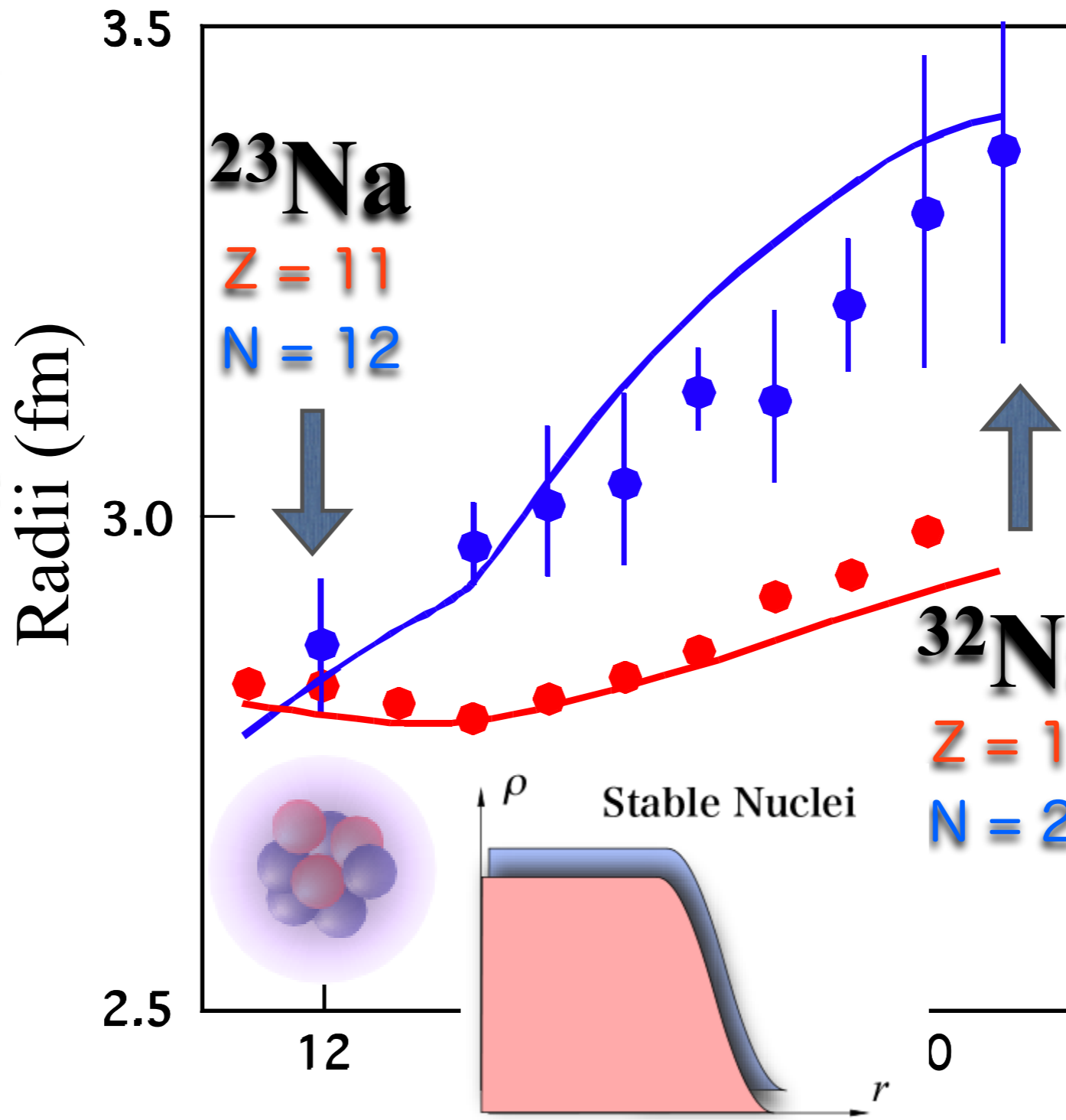


Nuclear Size and Skin formation



Nuclear Size of Na Isotopes

T. Suzuki et al., PRL 75 (1995) 3241



^{32}Na
 $Z = 11$
 $N = 21$

Theoretical Model

σ_1
↓

Neutron Radii

Comparison with known

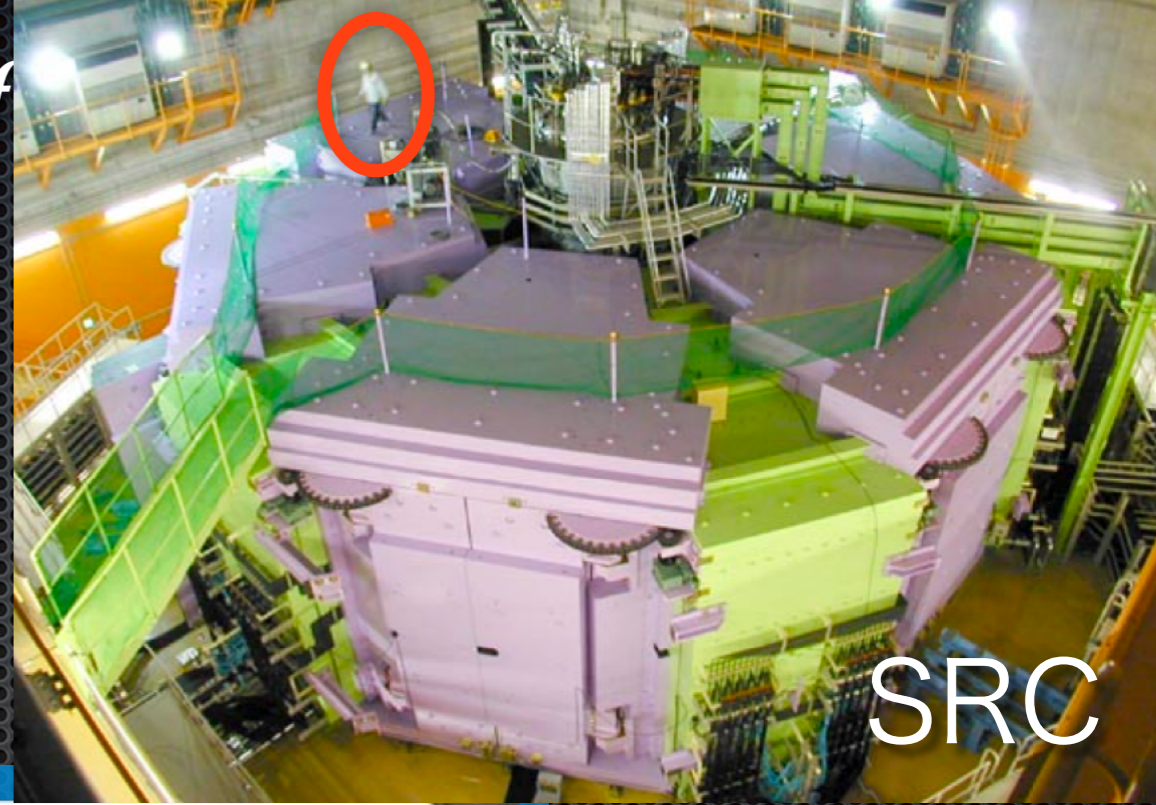
Proton Radii
(Laser spectroscopy)

Experiment : Measurements of

RIBF, RIKEN

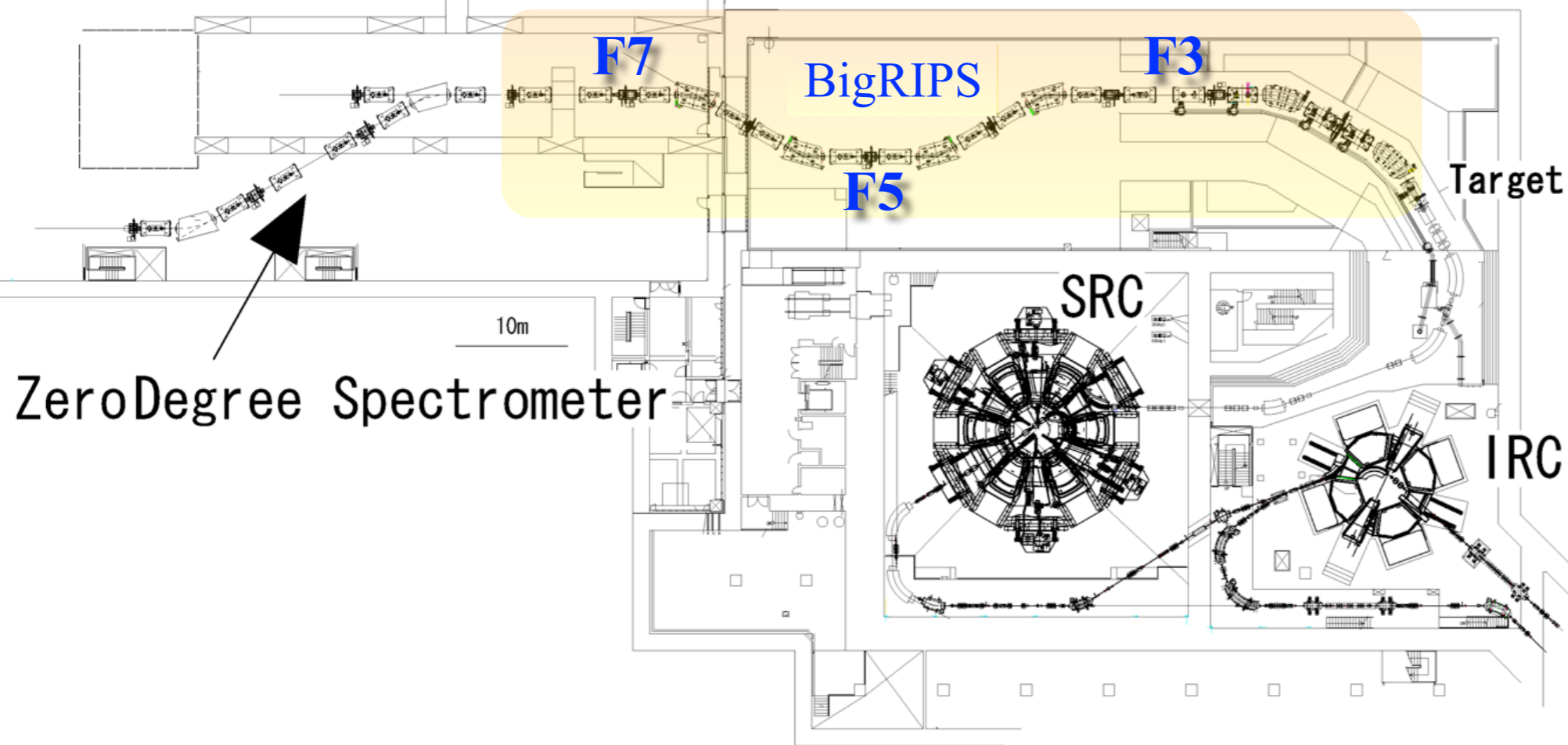
Primary Beam ^{48}Ca

Secondary Beam $^{20-32}\text{Ne}$

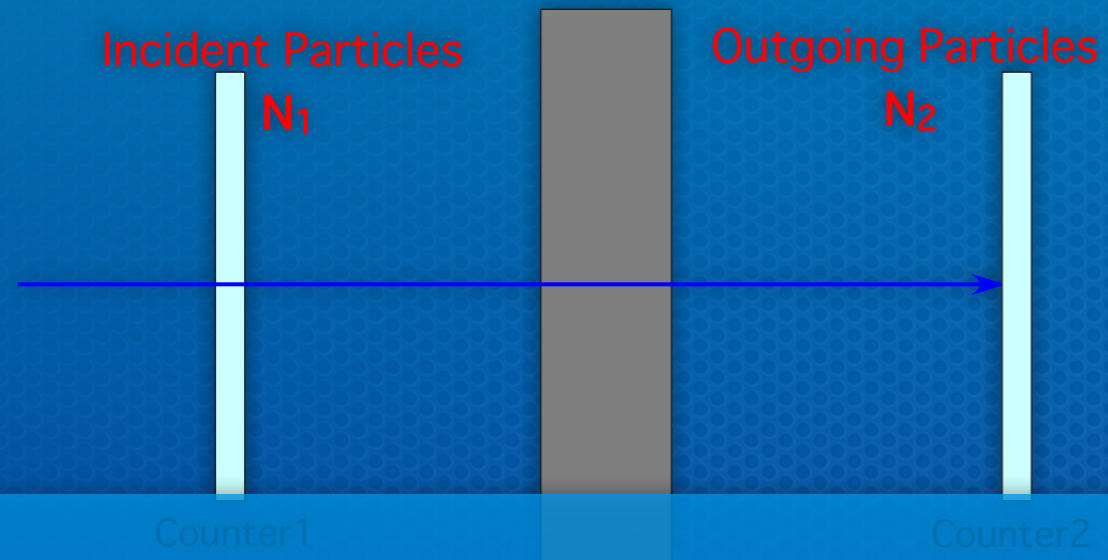


σ_I were measured using BigRIPS (F3 - F7)

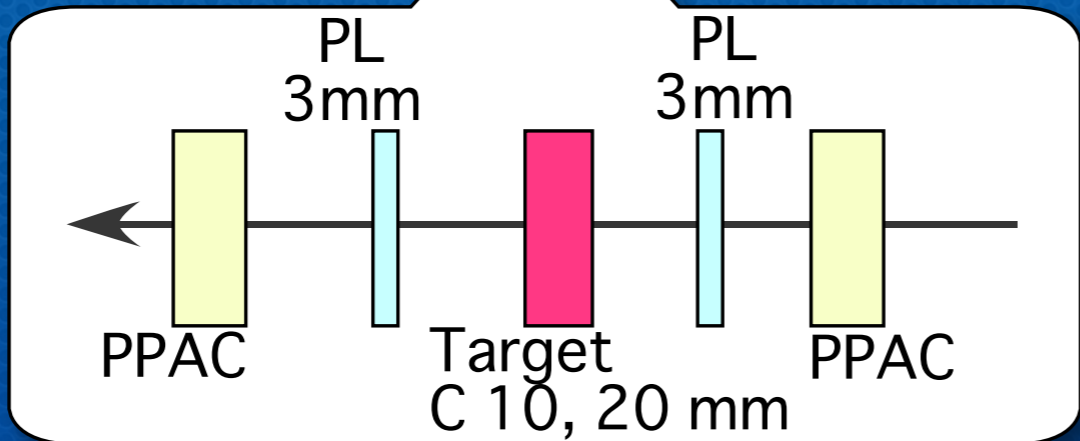
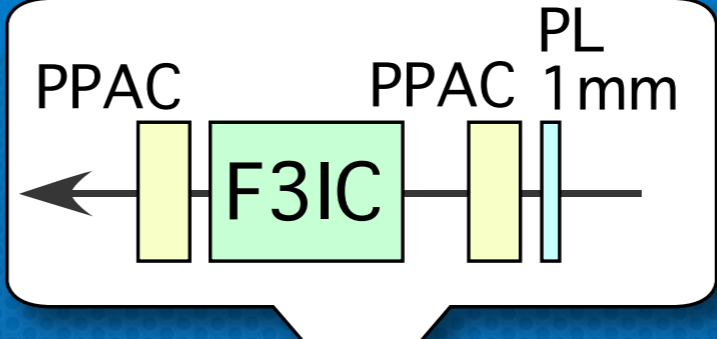
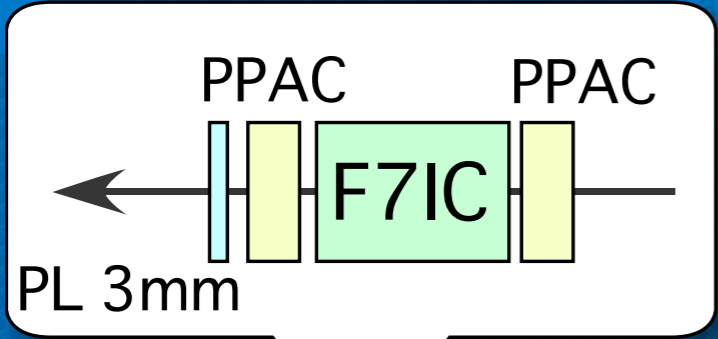
RI-beam delivery line



Transmission Method



$$\sigma_R = -\frac{1}{t} \ln\left(\frac{N_2}{N_1}\right)$$



$$B\rho \propto \frac{A}{Z} \cdot v$$

$$TOF \propto 1/v$$

$$\Delta E \propto z^2 / v^2$$

Neutron Skin Formation in Ne Isotopes

σ_I

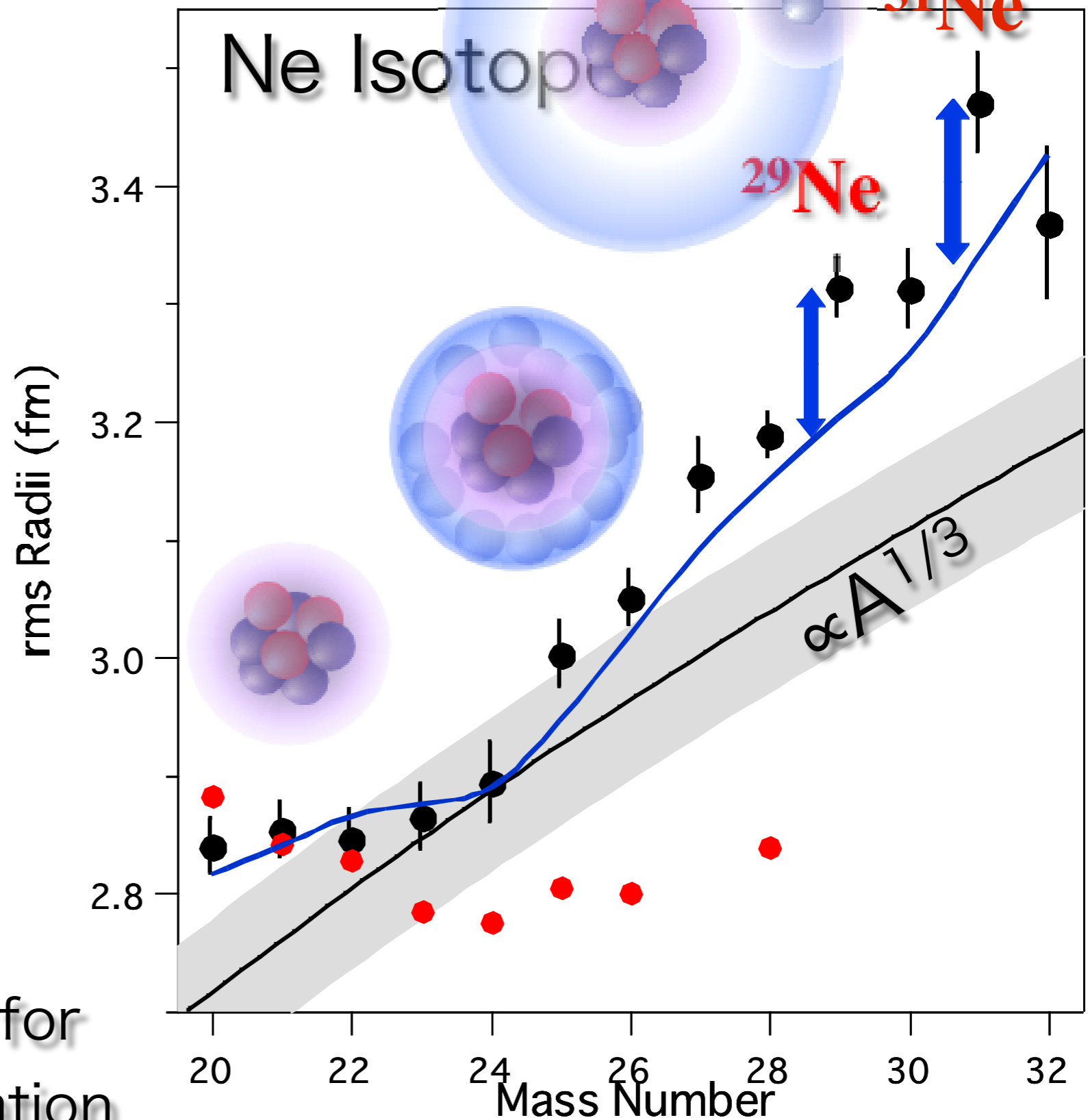


Matter Radii

**Comparison
with
known**

Proton Radii
(Laser spectroscopy)

Theoretical Model for
Neutron Skin Formation



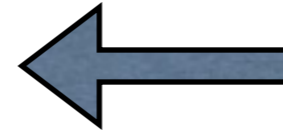
Near Future Experiment
Neutron Skin Determination for Ni isotope at RIBF

~ EOS for asymmetric nuclear matter ~

Study of Nuclear Matter from Symmetric to Asymmetric

Symmetric nuclear matter

- Saturation density $\sim 0.17 \text{ fm}^{-3}$
- Energy per particle $\sim -16 \text{ MeV}$
- Nearly incompressible



From Radii, Mass,
Collective Excitation data
for stable nuclei

Asymmetric nuclear matter

To describe the property of extreme matter

Neutron Star : Structure, Mass, Radius

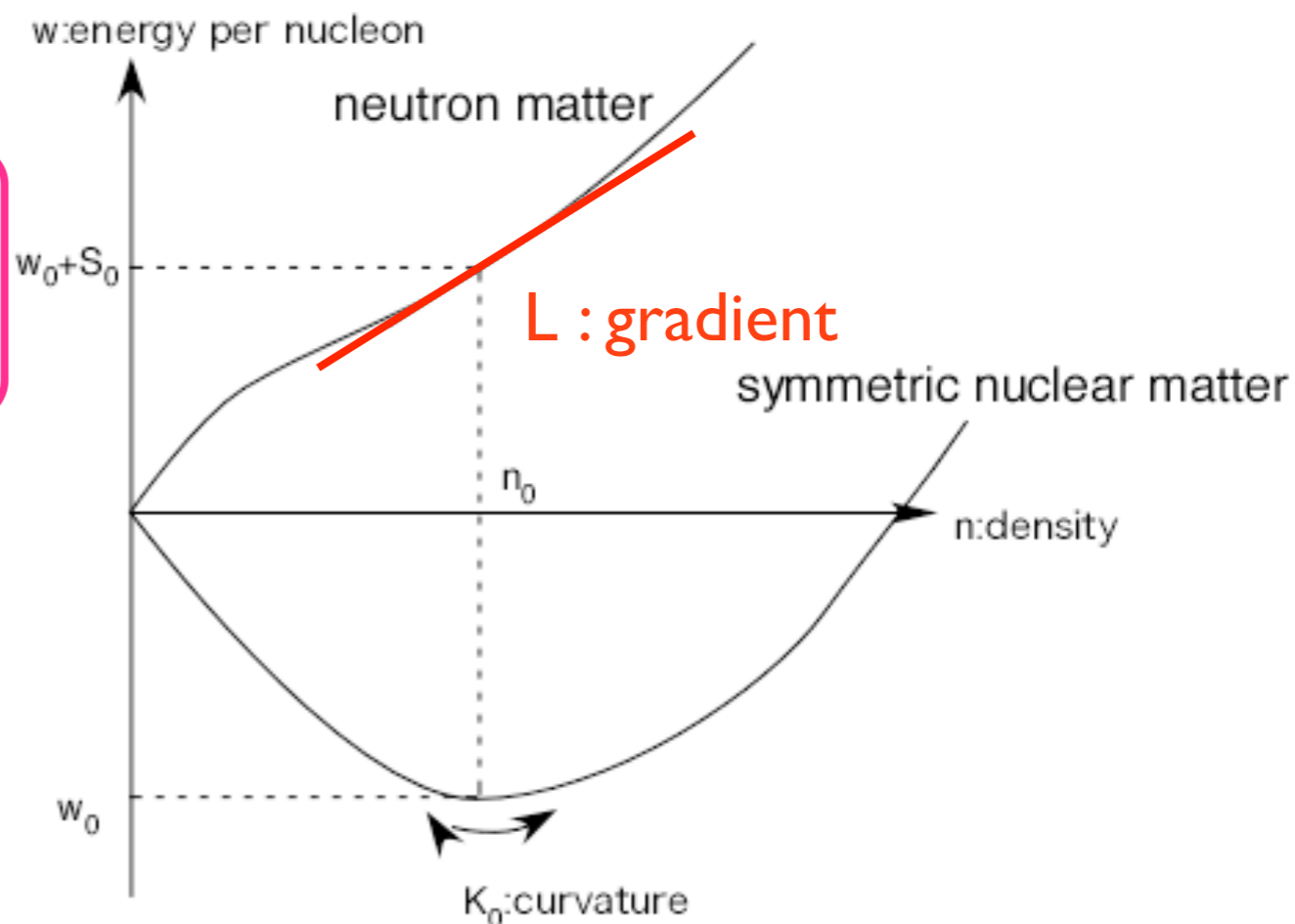
Energy per nucleon of nearly symmetric nuclear matter

$$w(n, \delta) \approx w_0 + \frac{K_0}{18n_0^2}(n - n_0)^2 + \delta^2 \left[S_0 + \frac{L}{3n_0}(n - n_0) \right]$$

n_0 : saturation density, w_0 : saturation energy
 K_0 : incompressibility, S_0 : symmetry energy at $n = n_0$
 $\delta = (N - Z)/A$

**L : Density derivative coefficient
of symmetry energy**

First key parameter L

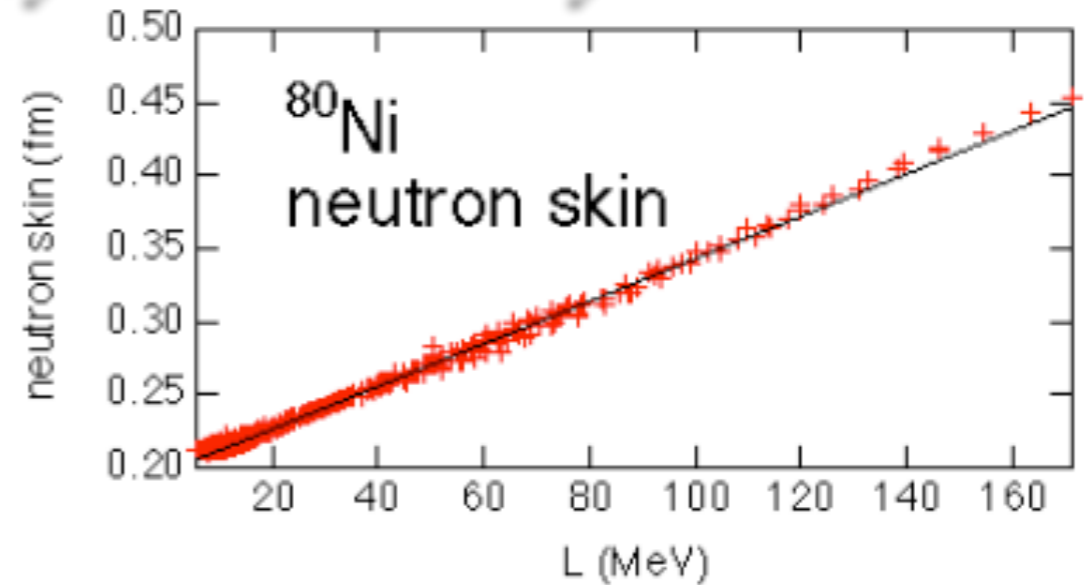


Study of Nuclear Matter from Symmetric to Asymmetric

How to know **L**?

Information from atomic nuclei

Many theories indicate strong correlation between neutron skin thickness and **L**



Kazuhiro Oyamatsu and Kei Iida Phys. Rev. C **81**, 054302 (2010)

One Simple correlation

M. Centelles et al., PRL **102**, 122502 (2009)

EOS around $N=Z$

$$w(n, \delta) \approx w_0 + \frac{K_0}{18n_0^2}(n - n_0)^2 + \delta^2 \left[S_0 + \frac{L}{3n_0}(n - n_0) \right]$$

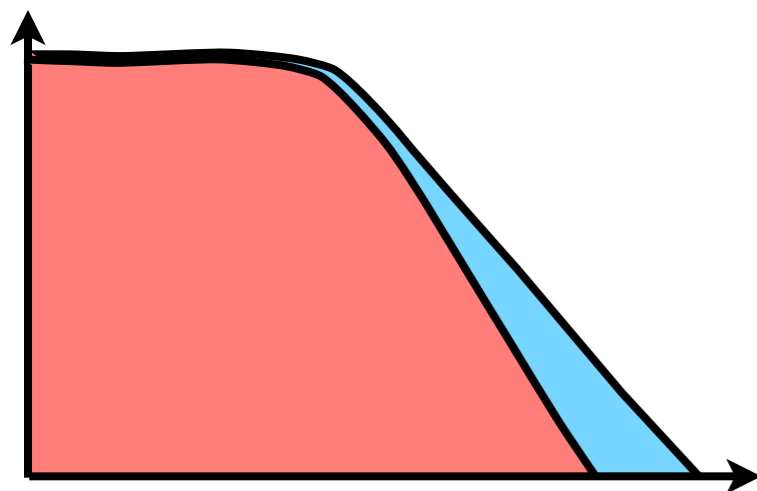
Droplet Model

$$E_B = a_V A - a_S A^{2/3} - a_C \frac{Z^2}{A^{1/3}} - a_A \frac{(A - 2Z)^2}{A} - \delta(A, Z)$$

When the density of nuclear matter is around nuclear surface density

Symmetry term $a_A \approx$ symmetry term of EOS

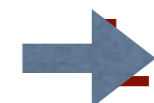
$\sim 0.1 \text{ fm}^{-3}$



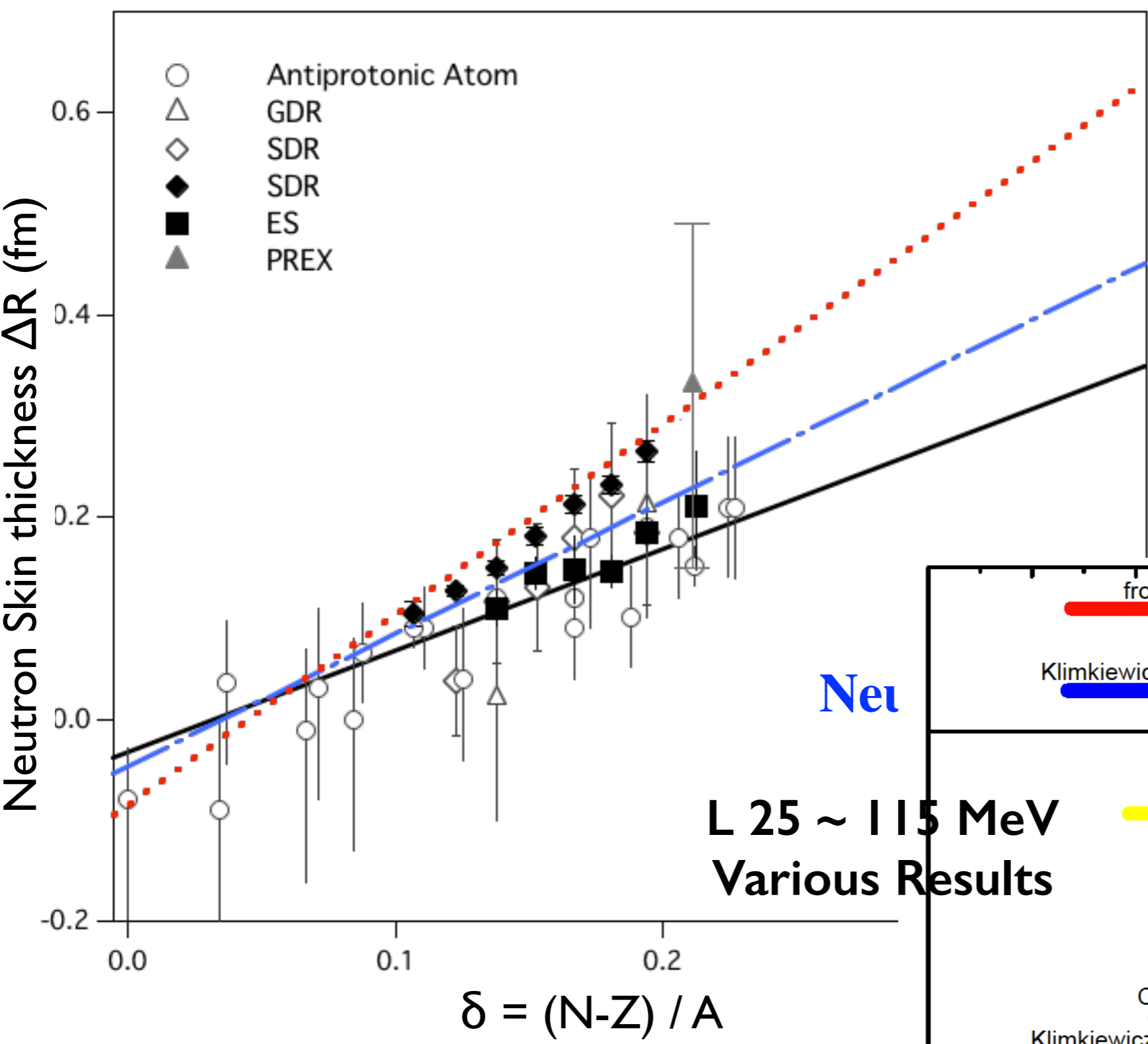
Neutron skin thickness $\Delta R \sim L \times \delta + \text{correction term}$

$$\delta = (N - Z)/A, \quad A > 40$$

Measurement of δ dependence of ΔR



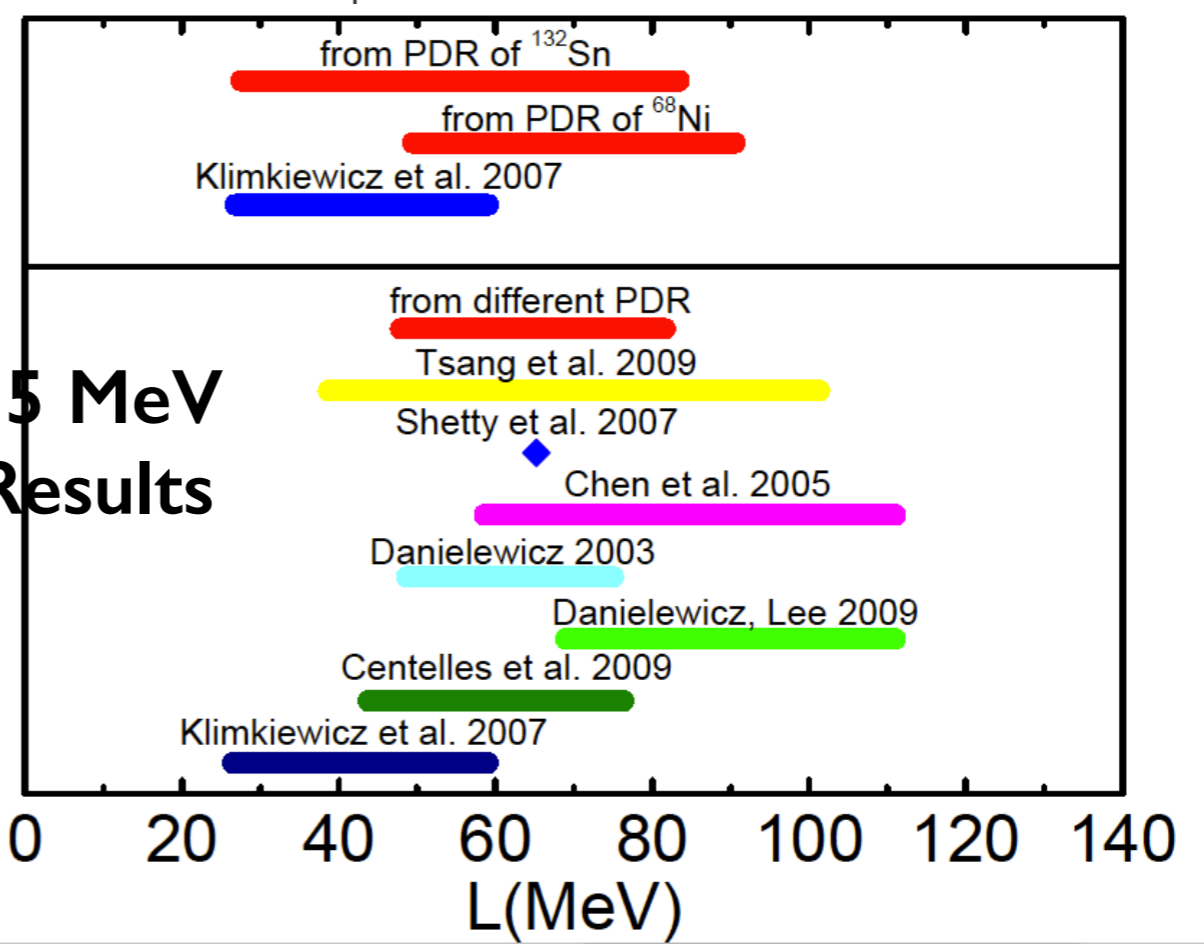
Neutron Skin Measurements and Models



Prediction from NL3 interaction
Neutron Star $R_{1.4M_{\odot}} = 15\text{km}$
 $M_{\text{max}} = 2.8 M_{\odot}$

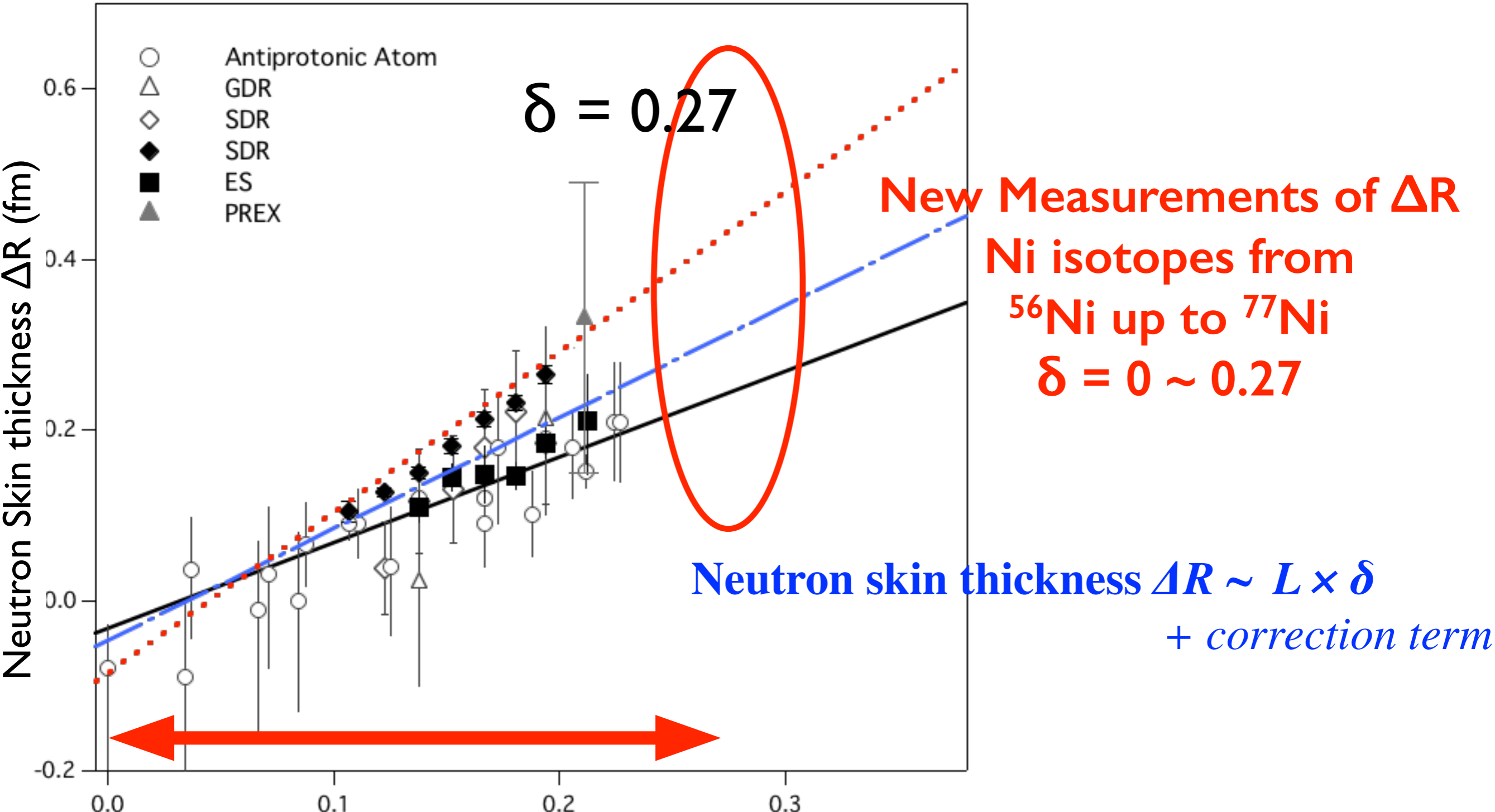
FSUGold interaction
Neutron Star $R_{1.4M_{\odot}} = 13\text{km}$
 $M_{\text{max}} = 1.7 M_{\odot}$

M. Centelles et al.,
Analysis of L using
Antiprotonic Atom Data
from $A = 40 \sim 238$



Anti Protonic Atom : A. Trzcinska et al., Phys. Rev. Lett. **87** 082501 (2001).
 GDR : A. Krasznahorkay et al., Nucl. Phys. A 567 521 (1994).
 SDR : A. Krasznahorkay et al., Phys. Rev. Lett. **82**, 3216 (1999).
 SDR : A. Krasznahorkay et al., Nucl. Phys. A 731 224 (2004).
 ES : S. Terashima et al., Phys. Rev. C 77 024317 (2008).
 ES : J. Zenihiro et al., Phys. Rev. C 82 044611 (2010).

Neutron Skin Measurements and Models



$$\delta = (N-Z) / A$$

$$\Delta R^2 = (A \cdot R_m^2 - Z \cdot R_p^2) / N$$

σ_I (Interaction cross section) \rightarrow Matter Radius

σ_{CC} (Charge changing cross section) \rightarrow Charge Radius